

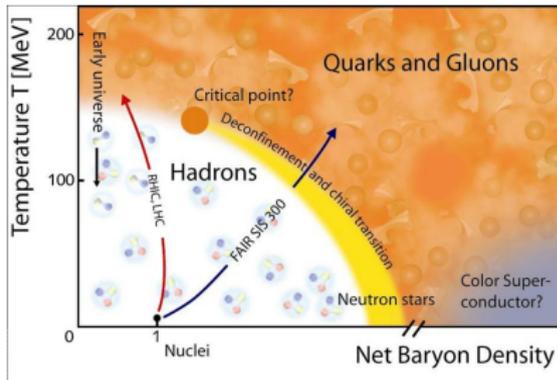
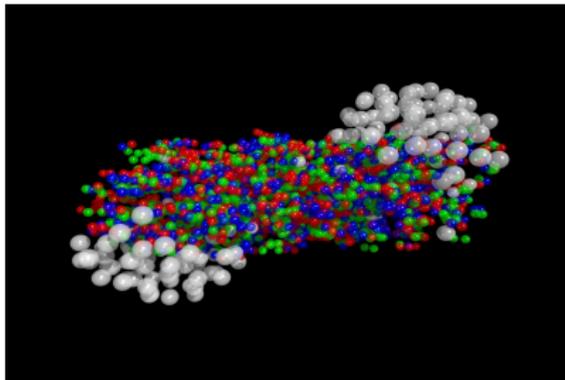
The Early Phase of High Energy Heavy Ion Collisions

Andreas Schäfer (Regensburg)

- Thermalization and entropy production
- 1. approach: Classical Field Theory
Kolmogorov-Sinai entropy and Husimi transformation
- 2. approach: AdS/CFT
- Conclusions

LHC: Investigating the phase diagram of QCD ?!?

p+p at 4 TeV + 4 TeV and p+Pb at 4 TeV + 1.58 A*TeV
and Pb+Pb at 1.38 A TeV + 1.38 A TeV



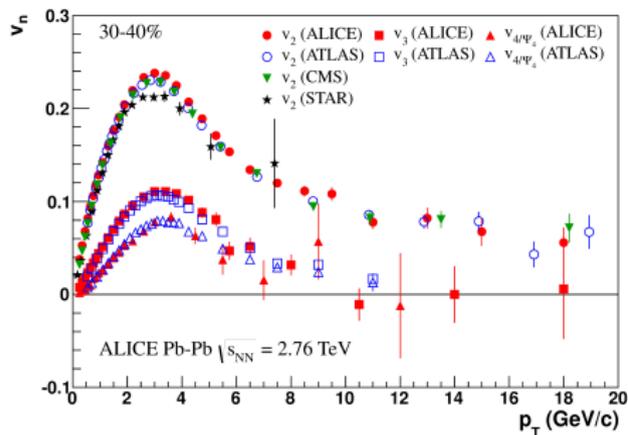
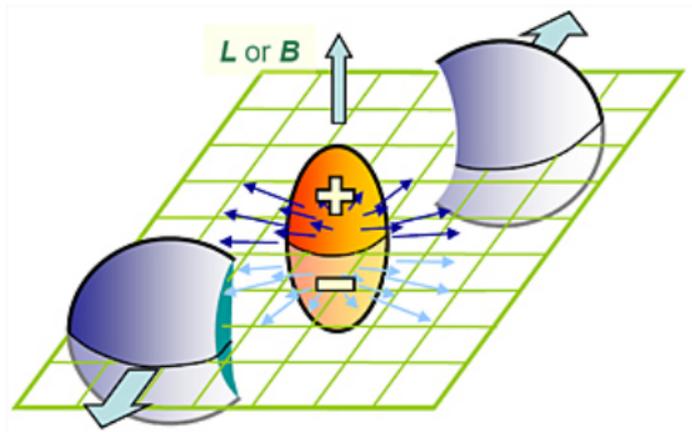
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GSI

Key question:

Does one reach thermal equilibrium fast enough to really probe the quark gluon plasma ?

Observable: Elliptic flow $v_n \sim \cos(n\phi)$

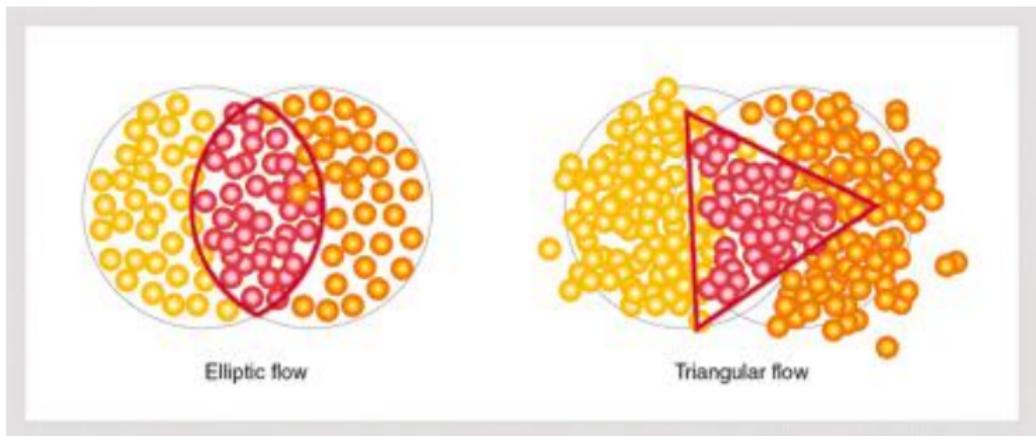


The importance of fluctuations: The system is not large

Example: v_3 and higher moments of flow, see below

$$v_n(p_T, \eta) = \langle \cos[n(\phi - \Psi_n)] \rangle$$

Ψ_n : True reaction plane orientation ϕ : measured angle of hadron



The ATLAS analysis from 1305.2942

The v_n describe the observed particle distributions

The eccentricities ϵ_n describe the initial state prior to hydrodynamic expansion (r, ϕ transverse coordinates)

$$\epsilon_n = \frac{\sqrt{\langle r^n \cos n\phi \rangle^2 + \langle r^n \sin n\phi \rangle^2}}{\langle r^n \rangle}$$

generic assumption: The v_n scale linearly with the ϵ_n .

In the following: $\langle v_n \rangle$ mean; σ_{v_n} standard deviation of v_n from $\langle v_n \rangle$

One assumes Gaussian fluctuations around a collective flow vector $\vec{v}_n = (v_n \cos(n\Phi_n), v_n \sin(n\Phi_n))$

$$p(\vec{v}_n) = \frac{1}{2\pi\delta_{v_n}^2} e^{-\frac{(\vec{v}_n - \vec{v}_n^{\text{RP}})^2}{2\delta_{v_n}^2}}$$

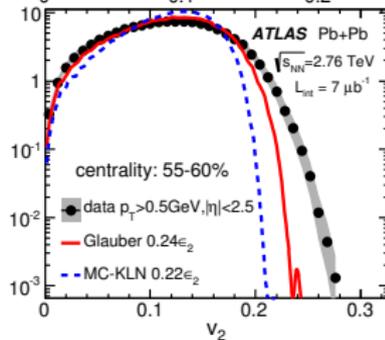
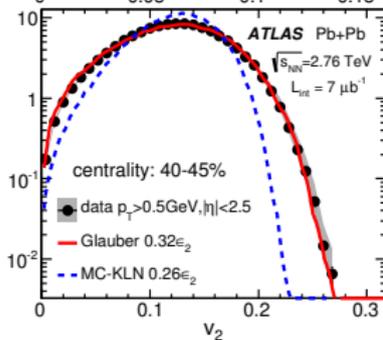
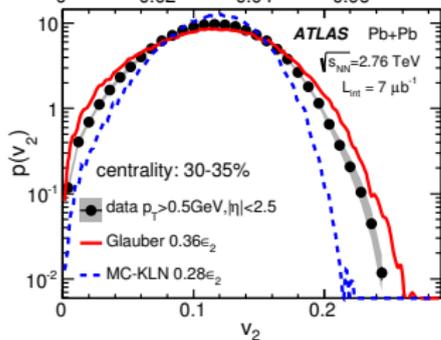
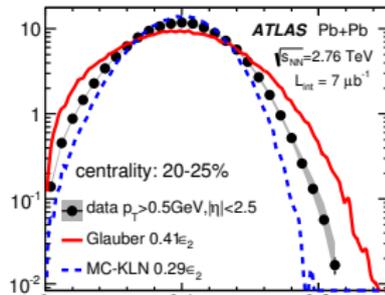
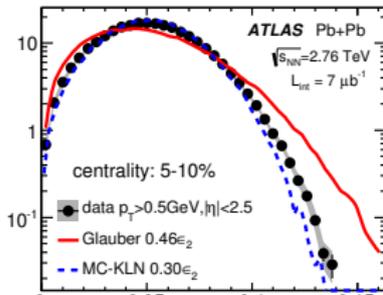
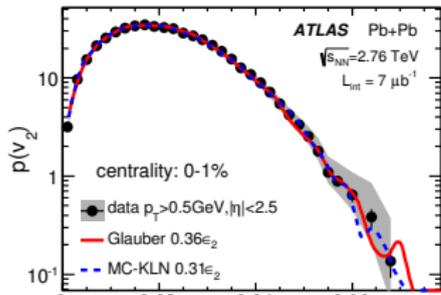
after integration over the azimuthal angle

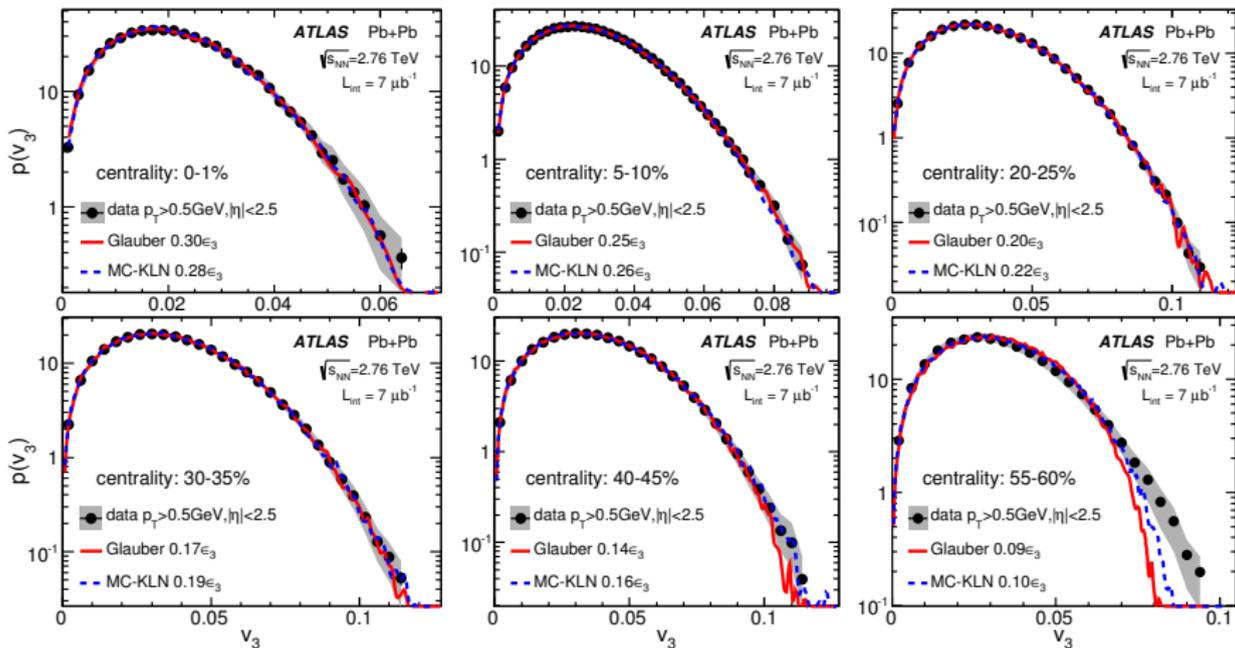
$$p(v_n) = \frac{v_n}{\delta_{v_n}^2} e^{-\frac{(v_n)^2 + (v_n^{\text{RP}})^2}{2\delta_{v_n}^2}} I_0\left(\frac{v_n^{\text{RP}} v_n}{\delta_{v_n}^2}\right)$$

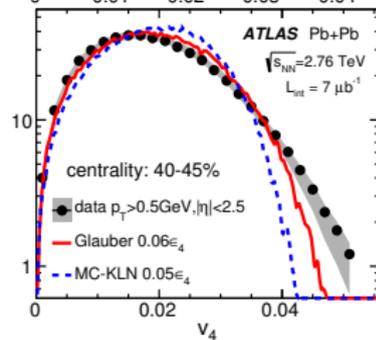
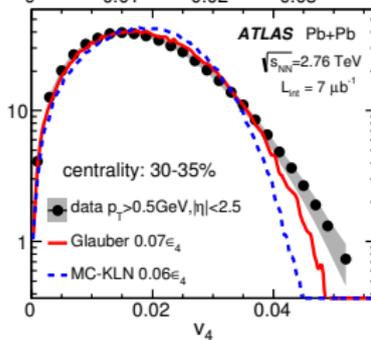
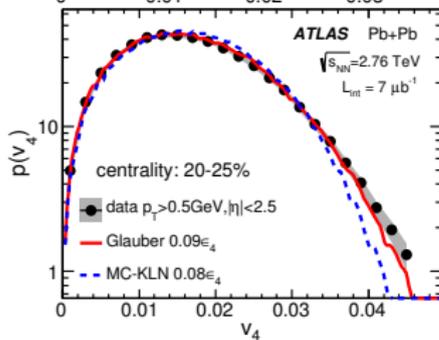
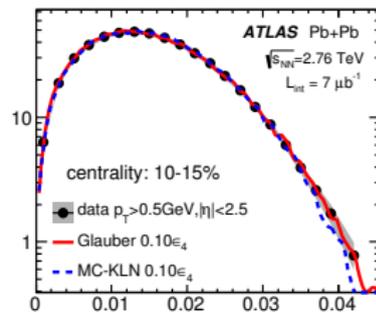
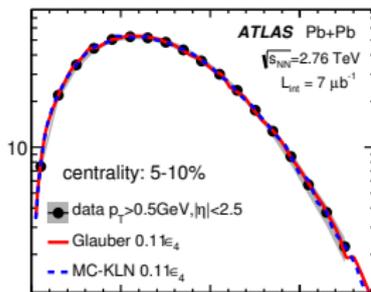
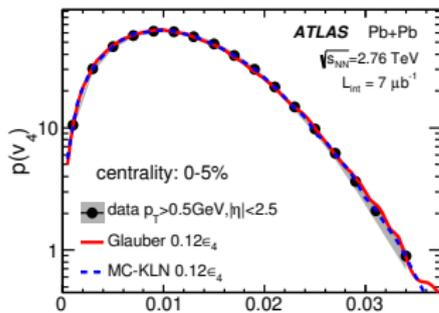
"fluctuations only" implies $v_n^{\text{RP}} = 0$ and $\sigma_{v_n} / \langle v_n \rangle = 0.523$

Very careful systematic studies are performed

Results in comparison with a Glauber Model and a CGC model (MC-KLN)





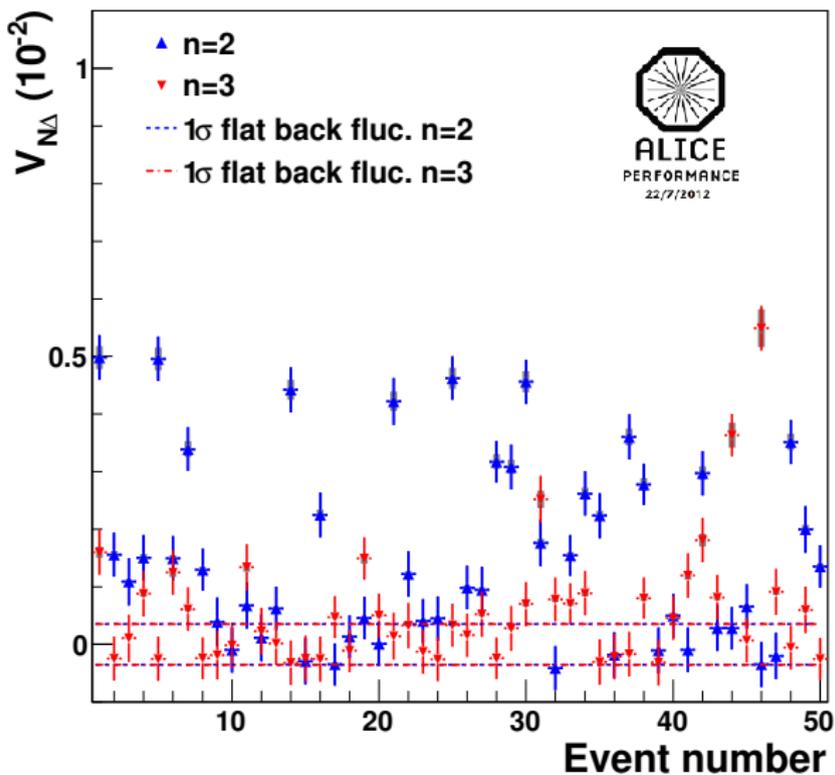


ATLAS:

Both models fail to describe the experimental data consistently over most of the measured centrality range

The ALICE analysis from 1301.6084

Pb-Pb $\sqrt{s_{NN}} = 2.76$ TeV, centrality 4-5%

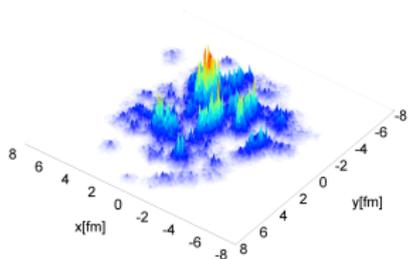


On an event-by-event basis one observes far larger fluctuations than expected

The two possible interpretations are:

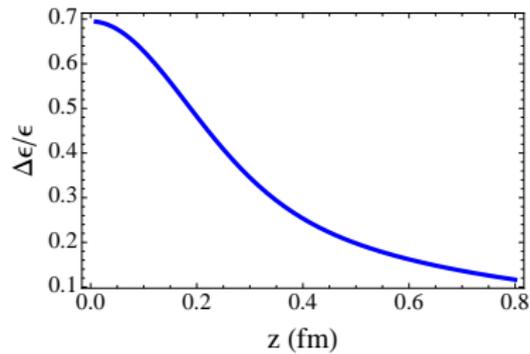
- different initial distributions, i.e. different ϵ_n
- or strongly non-linear dynamics, i.e. no proportionality between v_n and ϵ_n

Theory actually predicts large fluctuations



B. Schenke, Tribedy, Venugopalan 1206.6805

energy density at 0.2 fm/c



B.M. & A.S. 1111.3347

How is thermalization at all possible ?

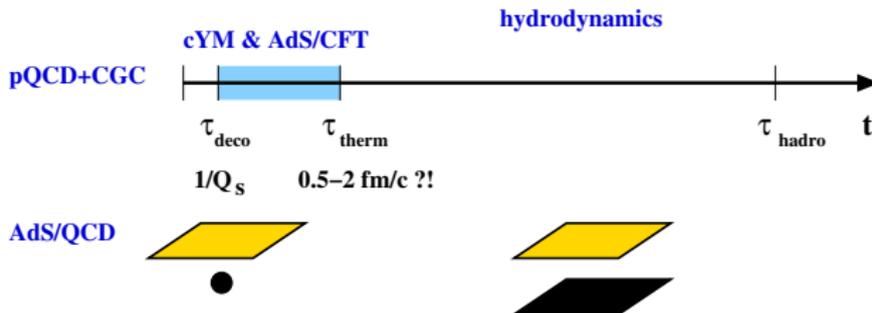
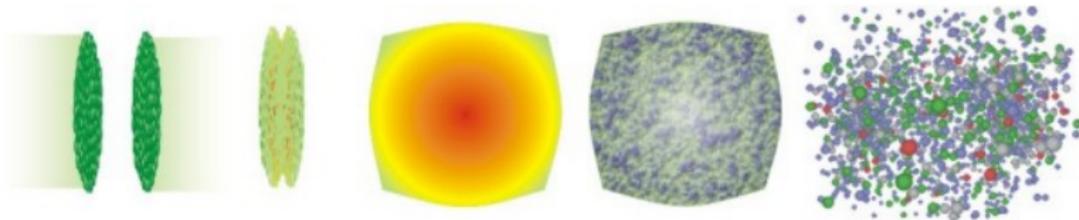
Thermalization requires massive entropy production at early stages ($< 1 \text{ fm}/c$)

But QCD is time-reversal invariant. Entropy is typically produced by measurements \Rightarrow Coarse-graining.

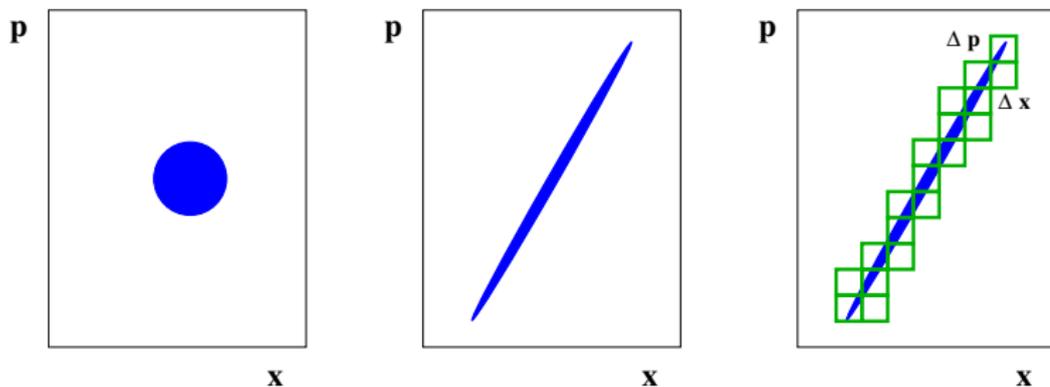
Q1: How can entropy be produced at all before any measurement takes place ?

Q2: How can it be produced so fast ?

Different stages of entropy production in a HIC

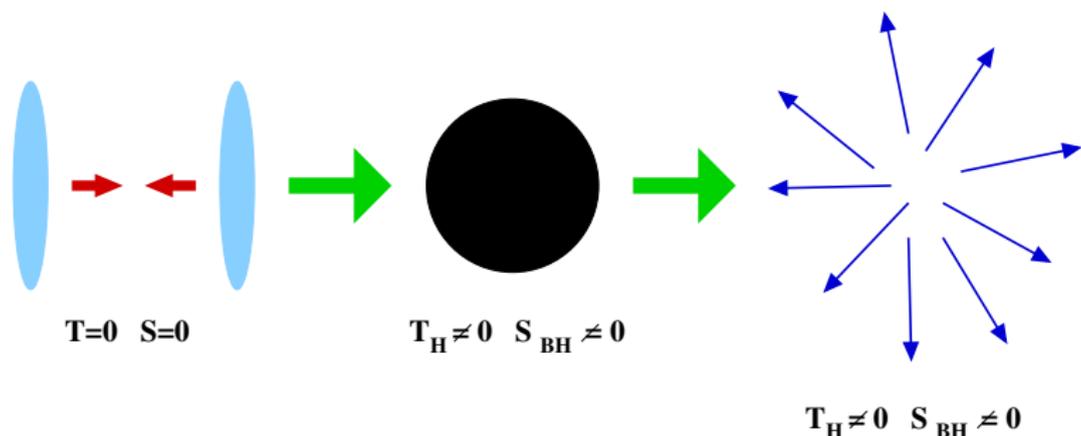


Coarse graining in non-linear mechanics



- Under time evolution the phase space volume is conserved
- The finite resolution of any measurement implies an increase in phase space volume

BH formation at LHC ?!?



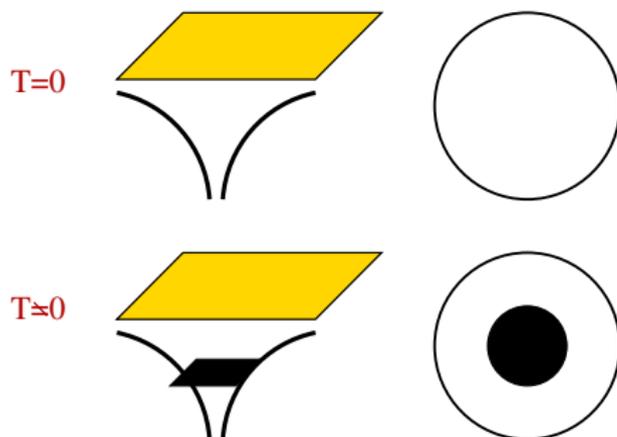
If the S-Matrix is unitary $S_{initial} = S_{final}$.

Entropy production corresponds to information loss.

Q1: Polchinski et al.: Unmeasurable information is lost information, Hawking radiation is not “really” a thermal but an entangled state

Holography: The Bekenstein entropy $S_B = k_B A / (4\ell_P^2)$ is proportional to the surface not the volume of a BH

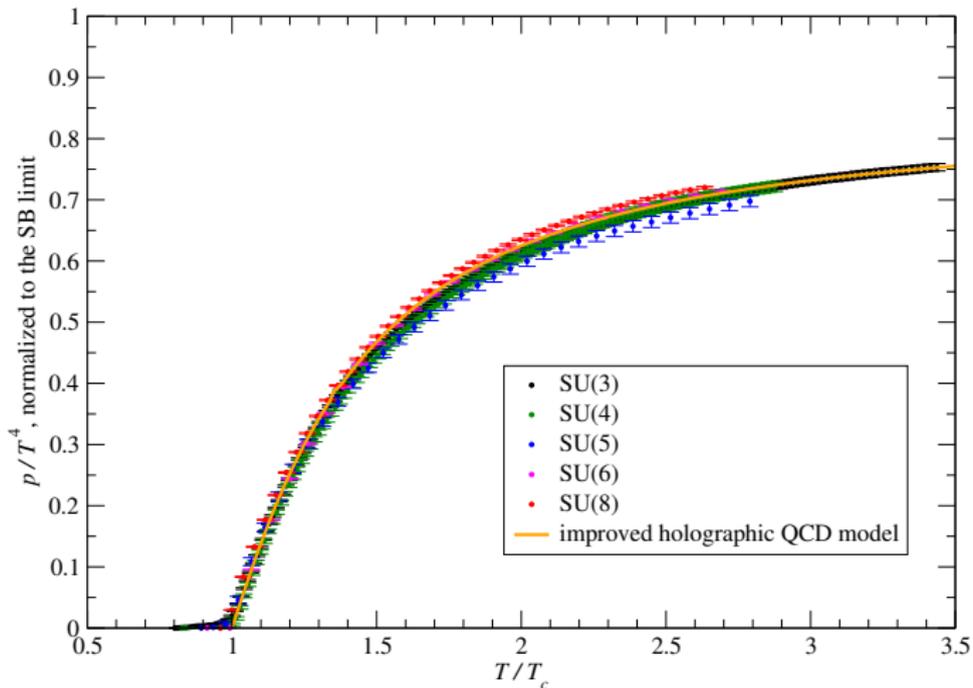
The dictionary



Hawking temperature = boundary temperature
(endless debates on hep-th for the non-equilibrium situation)

BH entropy problem = QCD entropy problem

Pressure



AdS/CFT is only valid for large $N \Rightarrow$ large N lattice simulations

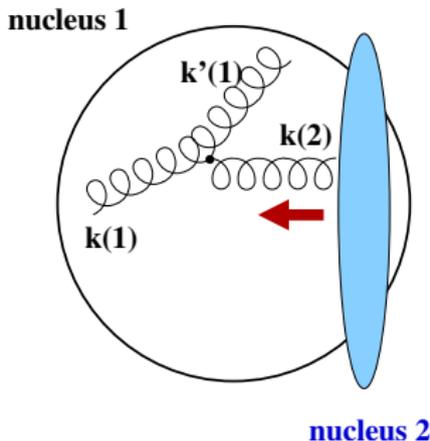
M. Panero, 0907.3719

I shall concentrate on the second phase but as a reminder:

1. Phase: Interaction between gluons from different nuclei;
pQCD evolution of the density matrix D .

decay time of the ratio $\text{Tr}D^2/(\text{Tr}D)^2$

generated entropy from $S(\tau_f) = \text{Tr} \{D(\tau_f) \log D(\tau_f)\}$



This turned out to be a very hard calculation giving a finite result only for (Λ IR cutoff of CGC model)

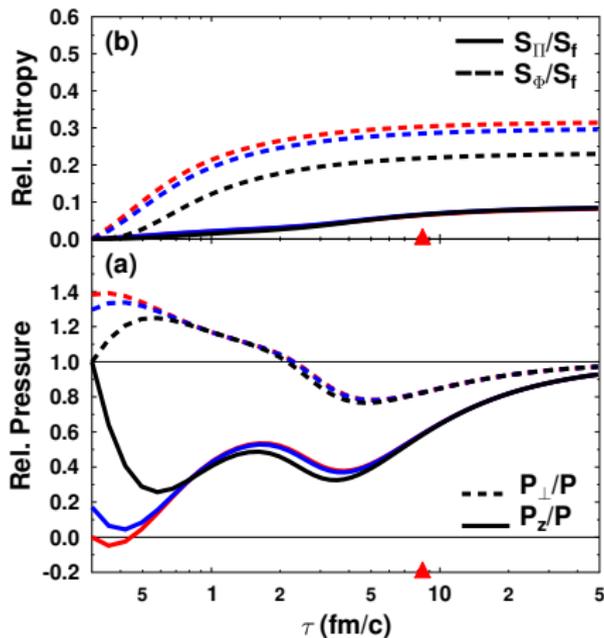
$$g^4 \rightarrow g^2(\Lambda^2) g^2(1/x^2)$$

which is exactly what was found by Kovchegov & Weigert and Balitsky.

Q1: Entropy is generated (at mid-rapidity) by mixing gluons from the two nuclei

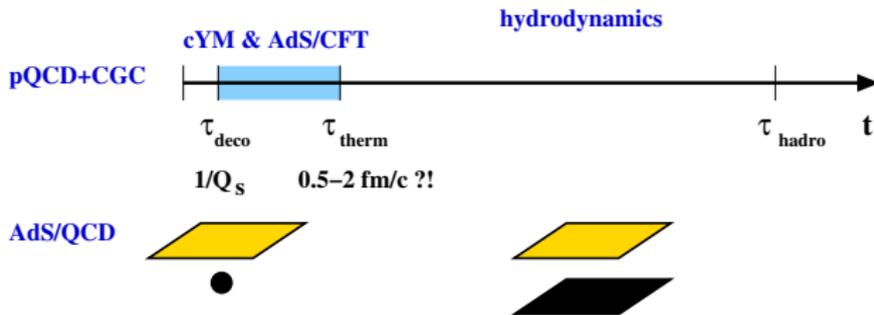
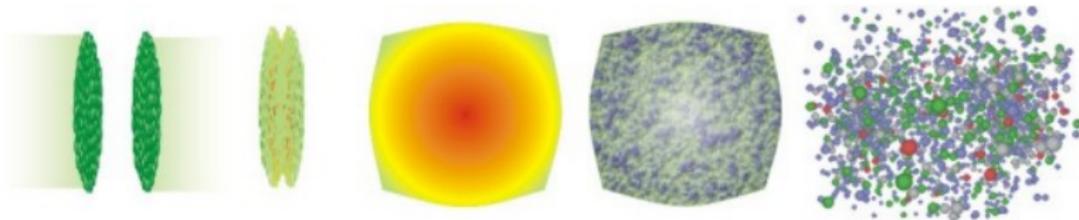
Q2: This process is very fast (0.3 fm/c) but only \sim 30-40 % of the needed entropy is generated

3. Phase: Viscous hydrodynamics



Conclusion: Nearly ideal hydrodynamics. Only little entropy is produced. Entropy production needs 1-2 fm/c)

Different stages of entropy production in a HIC



1. approach: Classical Field Theory

Non-linear Dynamics and Quantum Decoherence

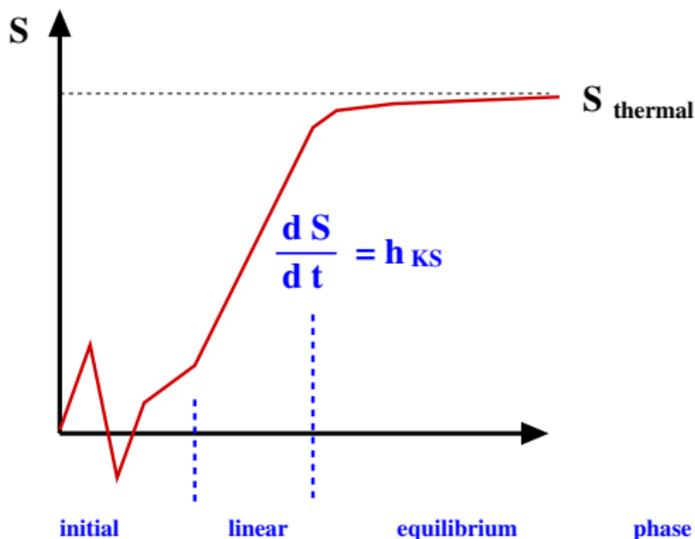
Q1: Husimi coarse graining

The Lyapunov exponents of a classical theory are determined numerically. The Kolmogorov-Sinai entropy is defined as

$$h_{KS} = \sum_{i, \lambda_i > 0} \lambda_i$$

The “Kolmogorov-Sinai entropy” is not entropy, but an entropy growth rate.

a generic picture



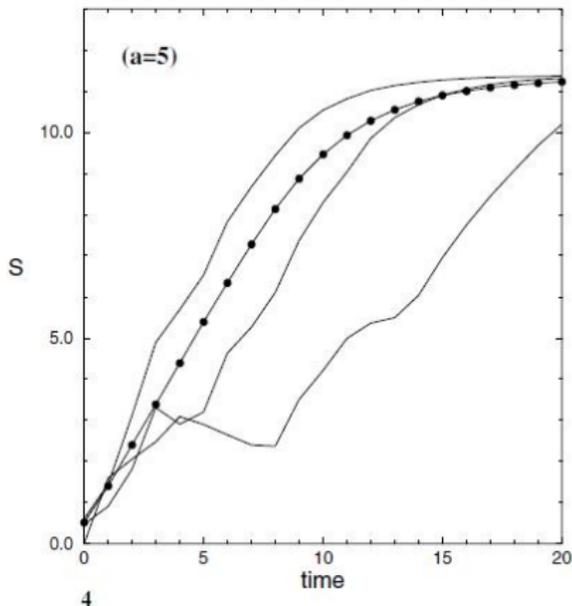
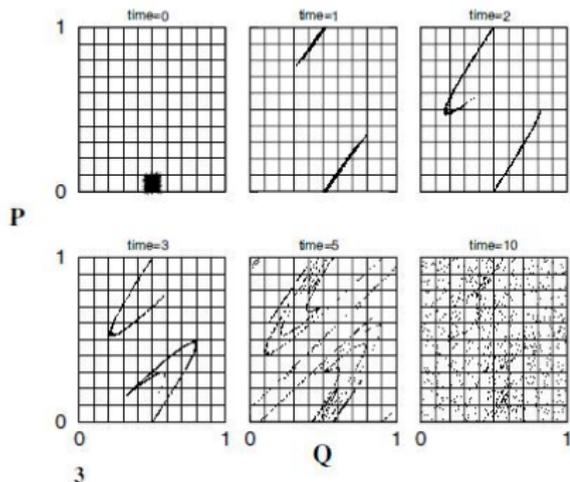
In the linear phase: $\frac{dS}{dt} = h_{KS}$

h_{KS} seems to play the same role in quantum theories !

Example: Standard map

M. Baranger, Chaos, Solitons and Fractals 13(2002)471

$$q, p \rightarrow Q, P : \quad P = p + \frac{q \sin(2\pi q)}{2\pi} \pmod{1}; \quad Q = q + P \pmod{1}$$



Husimi function and Wehrl entropy

The Husimi function

$$H_{\Delta}(p, x; t) = \int \frac{dp' dx'}{\pi \hbar} \exp\left(-\frac{1}{\hbar \Delta}(p - p')^2 - \frac{\Delta}{\hbar}(x - x')^2\right) W(p', x'; t)$$

The Wehrl entropy

$$S_{H, \Delta}(t) = - \int \frac{dp dx}{2\pi \hbar} H_{\Delta}(p, x; t) \ln H_{\Delta}(p, x; t); \quad \lim_{t \rightarrow \infty} \frac{dS_{H, \Delta}}{dt} = h_{KS}$$

Crucial assumption: Classical and quantum system have similar entropy growth rate

The Husimi function takes into account that the uncertainty principle implies coarse graining independent of the actual measurement.

It is easiest explained for a simple example: The inverse 1-dim oscillator.

$$\hat{\mathcal{H}} = \frac{1}{2}\hat{p}^2 - \frac{1}{2}\lambda^2\hat{x}^2$$

Initial state: Gaussian wave packet of width $\sqrt{\hbar/\omega}$:

$$\langle x|\psi_0\rangle = \left(\frac{\omega}{\pi\hbar}\right)^{1/4} e^{-\omega x^2/2\hbar}$$

The Wigner function associated with the density matrix $\hat{\rho}$ is defined as

$$W(p, x; t) = \int du e^{\frac{i}{\hbar}pu} \langle x - \frac{u}{2} | \hat{\rho}(t) | x + \frac{u}{2} \rangle$$

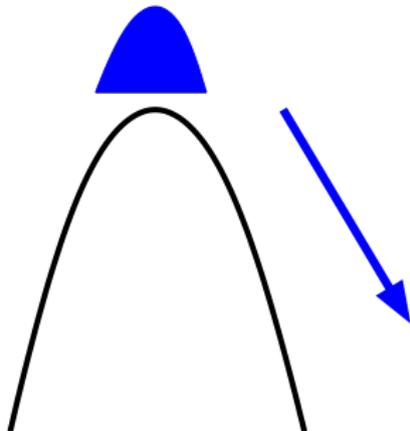
$$\int \frac{dp dx}{2\pi\hbar} W(p, x; t) = \text{Tr} [\hat{\rho}] = 1$$

$$\int \frac{dp dx}{2\pi\hbar} [W(p, x; t)]^2 = \text{Tr} [\hat{\rho}^2] \leq 1$$

The Wigner function is constant along the classical path.

$$x = x_0 \cosh \lambda t + \frac{p_0}{\lambda} \sinh \lambda t, \quad p = \lambda x_0 \sinh \lambda t + p_0 \cosh \lambda t,$$

Thus there is only one positive Lyapunov exponent, namely λ
and $h_{KS} = \lambda$



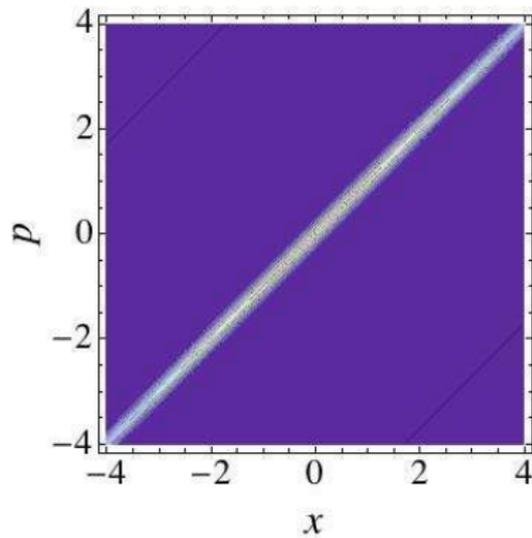
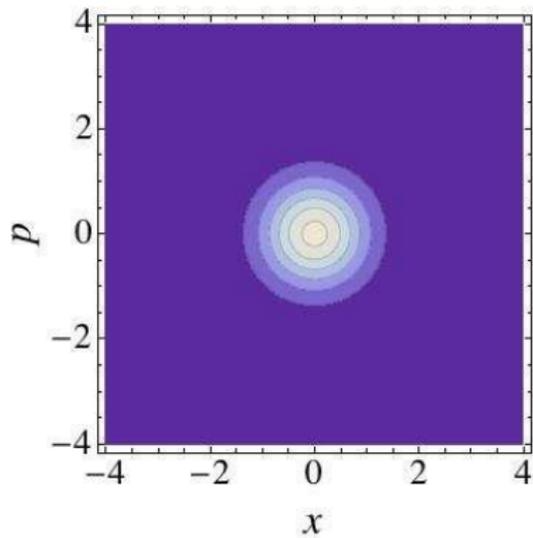
The Husimi function

$$H_{\Delta}(p, x; t) = \int \frac{dp' dx'}{\pi \hbar} \exp\left(-\frac{1}{\hbar \Delta}(p - p')^2 - \frac{\Delta}{\hbar}(x - x')^2\right) W(p', x'; t)$$

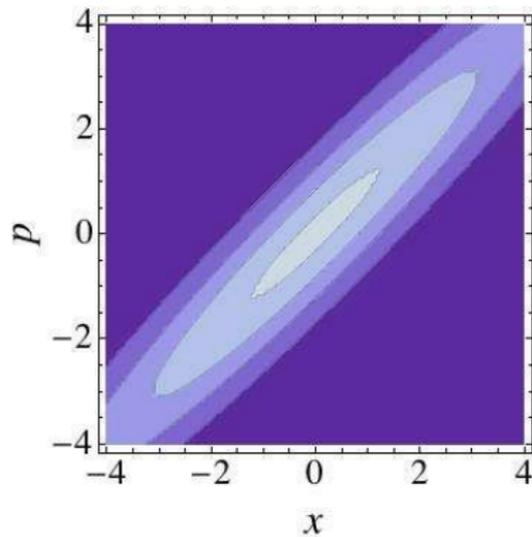
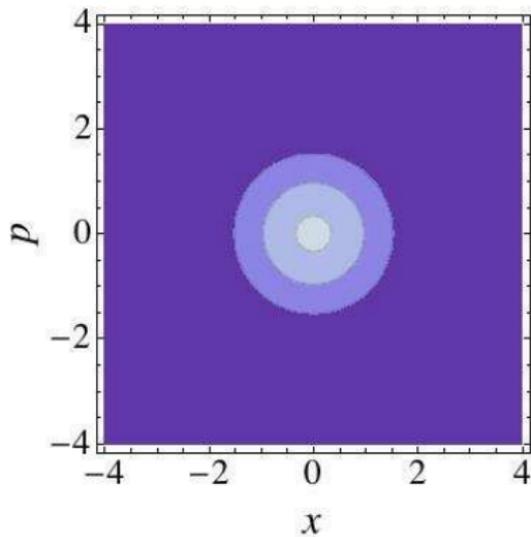
is non-negative, in contrast to the Wigner function. Δ is an arbitrary parameter corresponding to the actual measurement.

For this simple case the differential equation for W and the integration can be done analytically

The Wigner function for $t = 0$ and $t = 2/\lambda$



The Husimi function for $t = 0$ and $t = 2/\lambda$ for $\Delta = \lambda$. The phase space volume increases due to smearing.



For the Wehrl entropy

$$S_{H,\Delta}(t) = - \int \frac{dp dx}{2\pi\hbar} H_{\Delta}(p, x; t) \ln H_{\Delta}(p, x; t)$$

one finds

$$\lim_{t \rightarrow \infty} \frac{dS_{H,\Delta}}{dt} = \lambda = h_{KS}$$

independent of Δ !

Relation to Other Definitions of Entropy

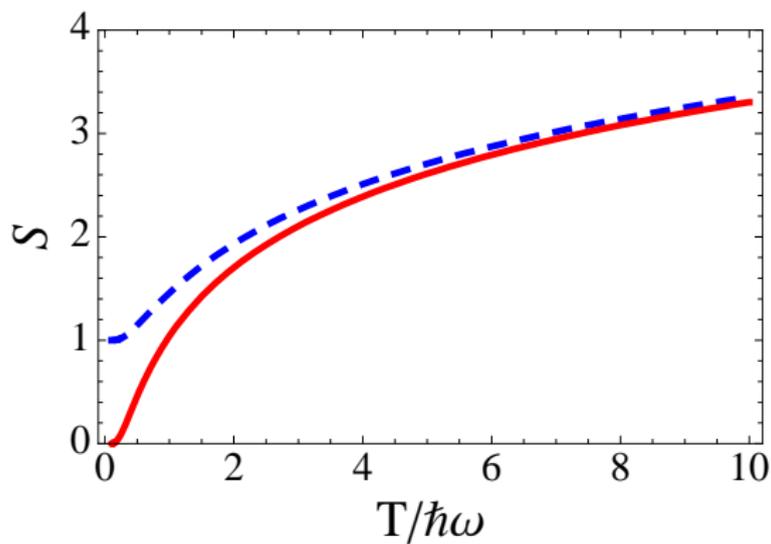
one harmonic oscillator; occupation probability of eigenstate $|n\rangle$:

$$w_n = e^{-n\beta\hbar\omega} / \mathcal{Z}_\beta$$
$$\mathcal{Z}_\beta = \sum_{n=0}^{\infty} e^{-n\beta\hbar\omega} = (1 - e^{-\beta\hbar\omega})^{-1}$$

von Neumann entropy

$$\begin{aligned} S_{\text{vN}} &\equiv - \sum_{n=0}^{\infty} w_n \ln w_n = \frac{\beta\hbar\omega}{e^{\beta\hbar\omega} - 1} - \ln(1 - e^{-\beta\hbar\omega}) \\ &= -\bar{n} \ln \bar{n} + (\bar{n} + 1) \ln(\bar{n} + 1) \end{aligned}$$

S_{VN} (solid) versus S_H (dashed)



$$H = \frac{1}{2} \sum_{x,a,i} E_i^a(x)^2 + \frac{1}{4} \sum_{x,a,i,j} F_{ij}^a(x)^2$$

$$F_{ij}^a(x) = \partial_i A_j^a(x) - \partial_j A_i^a(x) + \sum_{b,c} f^{abc} A_i^b(x) A_j^c(x)$$

$$\delta \dot{X}(t) = \mathcal{H} \delta X(t)$$

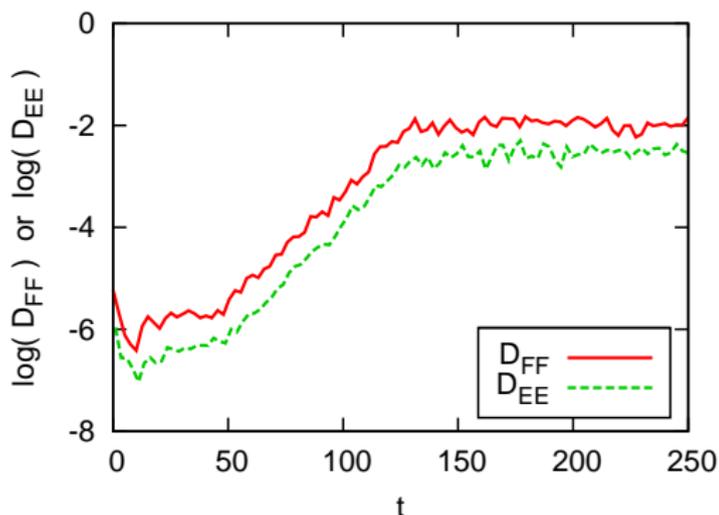
$$\dot{A}_i^a(x) = E_i^a(x)$$

$$\dot{E}_i^a(x) = \sum_j \partial_j F_{ji}^a(x) + \sum_{b,c,j} f^{abc} A_j^b(x) F_{ji}^c(x)$$

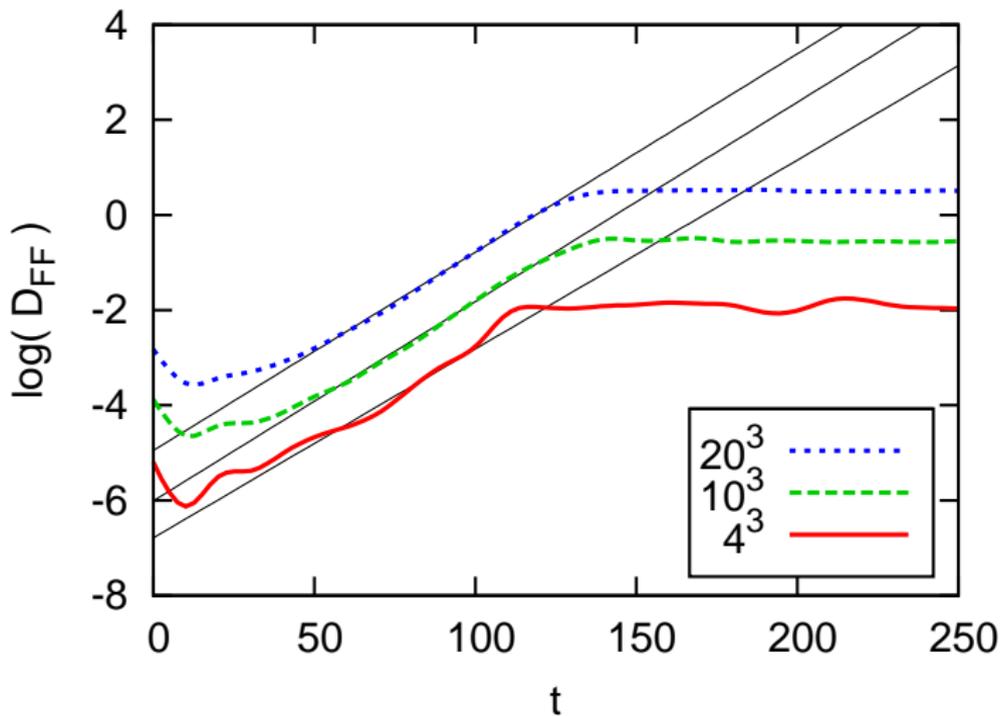
Different distance measures give the same result.

$$D_{EE} = \sqrt{\sum_x \left\{ \sum_{a,i} E_i^a(x)^2 - \sum_{a,i} E_i'^a(x)^2 \right\}^2}$$

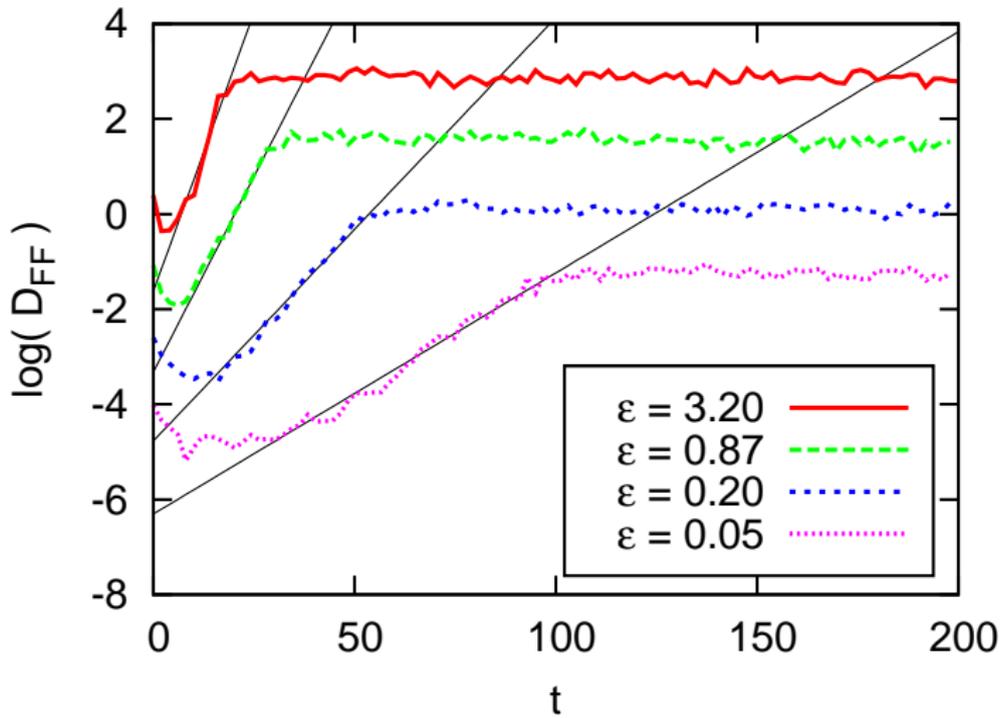
$$D_{FF} = \sqrt{\sum_x \left\{ \sum_{a,i,j} F_{ij}^a(x)^2 - \sum_{a,i,j} F_{ij}'^a(x)^2 \right\}^2}$$



Time evolution in SU(2) simulation on 4^3 , 10^3 , and 20^3 lattices with the same energy density



Time evolution in SU(3) simulation on a 4^3 lattice



Classical YM theories are UV divergent

⇒ The lattice constant has the physical meaning of a screening length

Scale setting can be done by different means:
equation the energy densities on the lattice

$$\epsilon_{cl}(T) = \frac{\epsilon^L}{a^4 g^2}$$

$$\epsilon^L = 2(N_c^2 - 1) \frac{1}{L^3} \sum_{\mathbf{k}} |\mathbf{k}| \frac{T^L}{|\mathbf{k}|}$$

$$\epsilon(T) = 2(N_c^2 - 1) \int \frac{d^3k}{(2\pi)^3} \frac{|\mathbf{k}|}{e^{|\mathbf{k}|/T} - 1}$$

or choosing the damping length of SU(N) gauge theories

All approaches give

$$a = \frac{\theta}{T} \sim \epsilon^{-1/4}$$

with θ of order unity

The thermalization time is estimated from

$$\tau_{\text{eq}} = \frac{\Delta s}{s_{KS}} + \tau_{\text{delay}}$$

The delay time is found to be of the order

$$\frac{1}{6(N_c^2 - 1)L^3} e^{\lambda_{\text{max}} \tau_{\text{delay}}} \approx 1$$

Q2: numerically we find

$$\tau_{\text{eq}} \approx 2 \text{ fm}/c$$

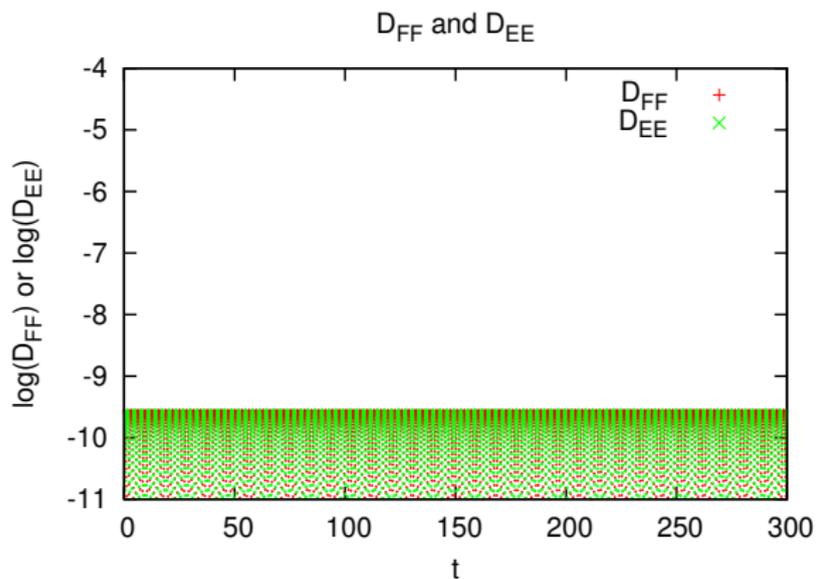
with substantial theoretical uncertainties

⇒ This would suggest substantial corrections due to incomplete thermalization for the LHC experiments.

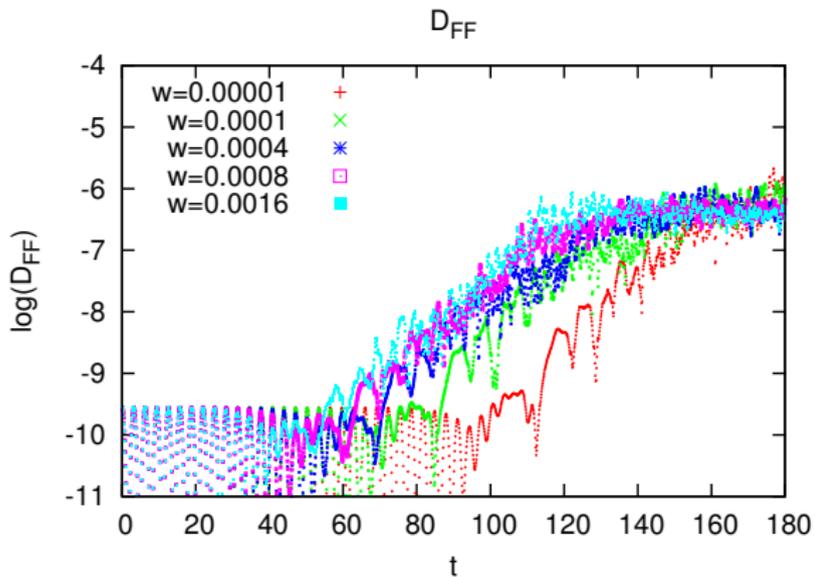
Initial fluctuations might make $\leq 1 \text{ fm}/c$ possible

We tested whether more realistic initial gauge field configurations lead to smaller τ_{eq} .

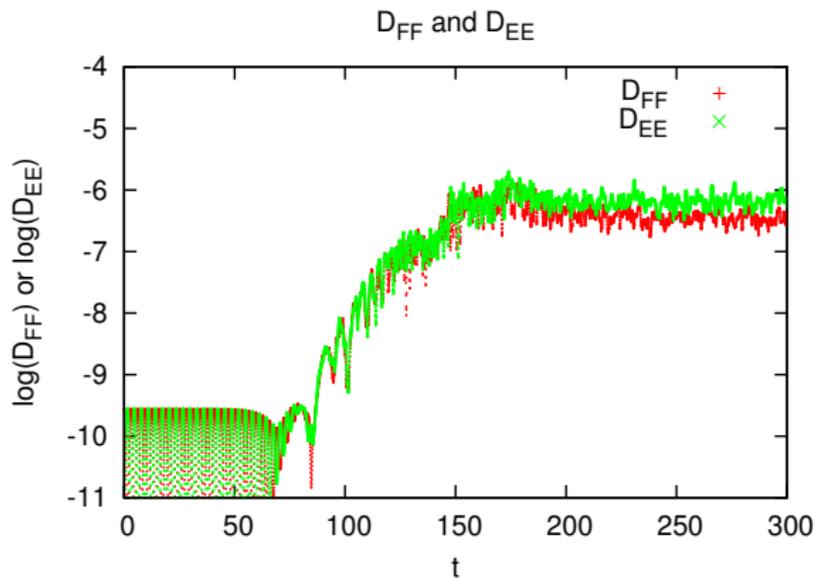
$$A_i^a(\vec{r}) = \delta_{i2} \left[\epsilon_1 \sin\left(\frac{2x\pi}{N_x}\right) + \epsilon_2 \sin\left(\frac{2nx\pi}{N_x}\right) \sin\left(\frac{2mz\pi}{N_z}\right) \right] + \eta_i^a(\vec{r})$$
$$E_i^a(\vec{r}) = 0$$

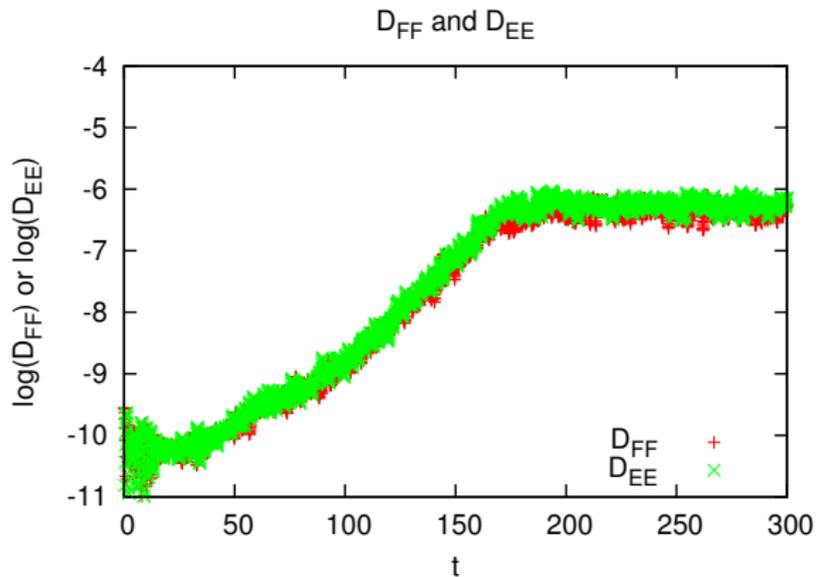


CGC-type field without initial fluctuations \Rightarrow purely numerical fluctuations are under control.

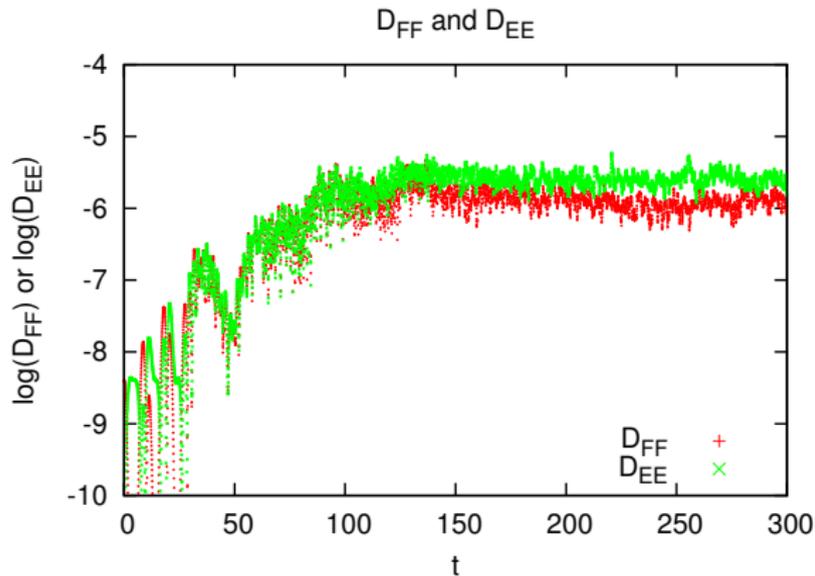


effect of increasing fluctuations





For comparison: Behavior for fluctuations only



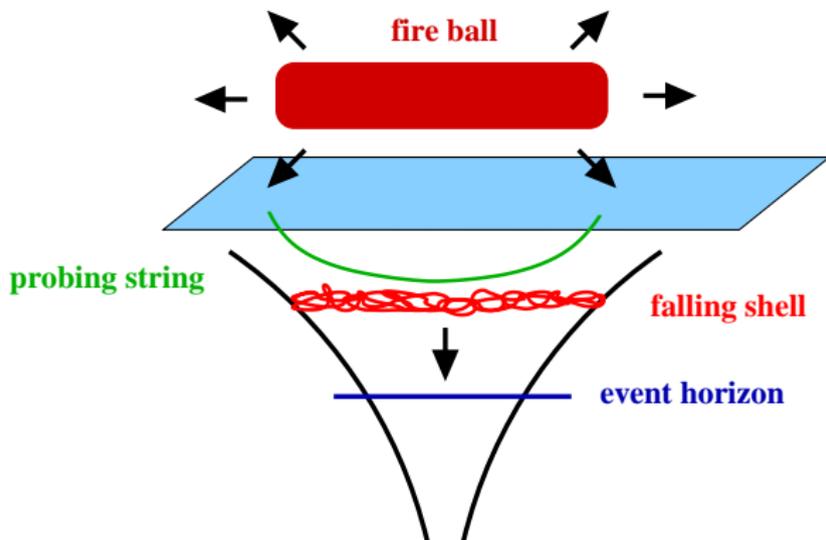
Constant background plus fluctuations

Upshot: The behavior is indeed generic

τ_{delay} depends on details

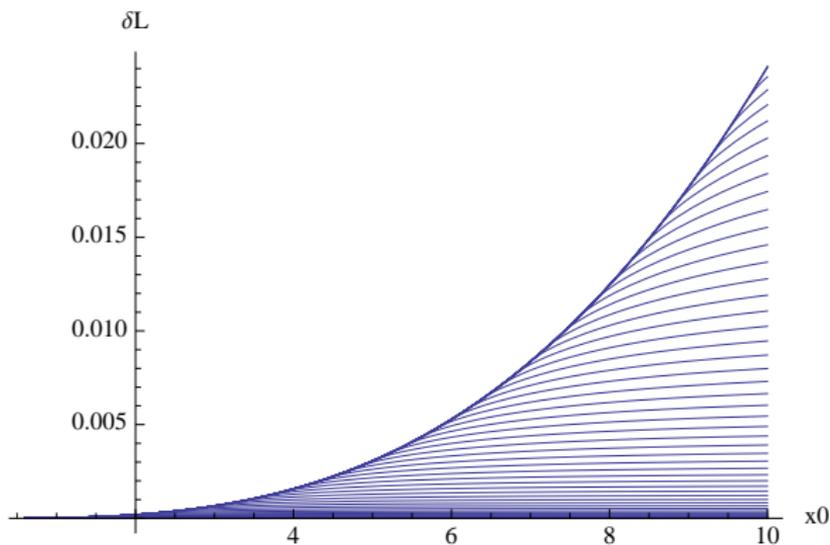
2. approach: Equilibration times from AdS/CFT

Idea: Probe black brane formation with a string or membrane.
PRL: 1012.4753 PRD: 1103.2683

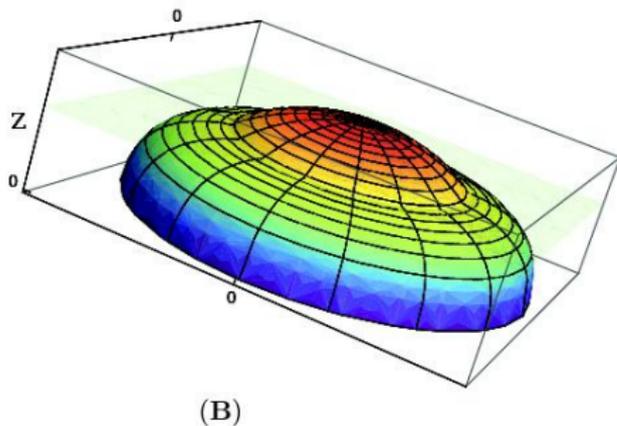
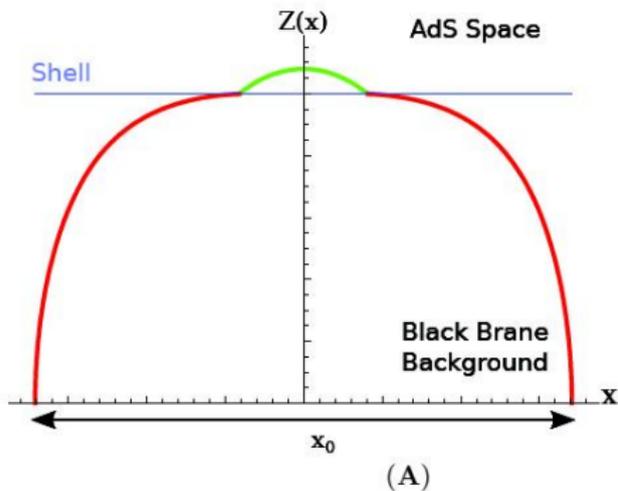


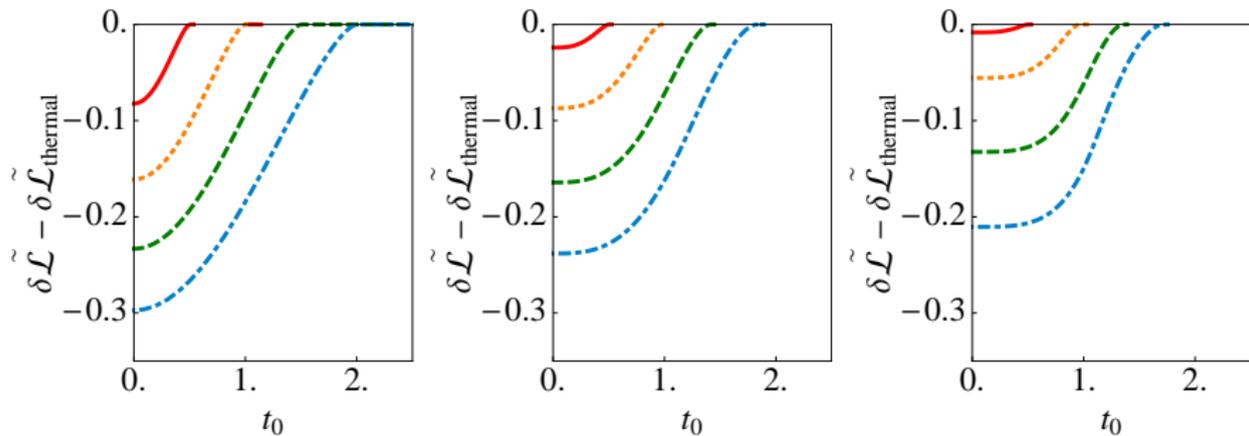
The change in geodesic length is sensitive to equal time correlators of high dimension gluonic operators.

$$\frac{\langle \mathcal{O}(t_{shell}, x) \mathcal{O}(t_{shell}, 0) \rangle_{shell}}{\langle \mathcal{O}(t_{shell}, x) \mathcal{O}(t_{shell}, 0) \rangle_{AdS}} \approx e^{-\Delta \delta L(t_{shell}, x)}$$

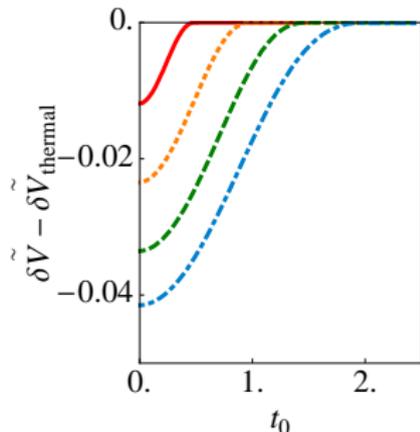
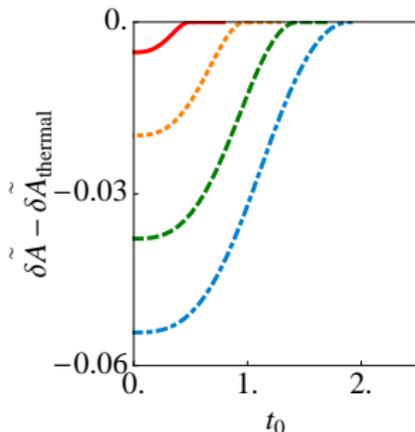
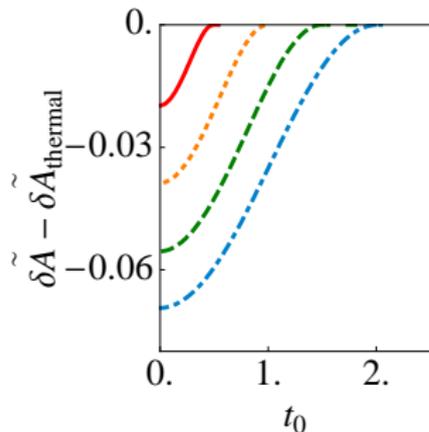


We solved analytically and numerically different cases:
 $AdS_3 \sim CFT(1 + 1)$, $AdS_4 \sim CFT(1 + 2)$, $AdS_5 \sim CFT(1 + 3)$
and analysed how the length of the geodesic/the area of the surface approaches its thermal value, as a function of ℓ and t_0 .





$\delta\tilde{\mathcal{L}} - \delta\tilde{\mathcal{L}}_{\text{thermal}}$ ($\tilde{\mathcal{L}} \equiv \mathcal{L}/\ell$) for $d = 2, 3, 4$ (left, right, middle) and $\ell = 1, 2, 3, 4$ (top to bottom curve).



$\delta\tilde{\mathcal{A}} - \delta\tilde{\mathcal{A}}_{\text{thermal}}$ ($\tilde{\mathcal{A}} \equiv \mathcal{A}/\pi R^2$ and $\delta\tilde{\mathcal{V}} - \delta\tilde{\mathcal{V}}_{\text{thermal}}$
 $(\tilde{\mathcal{V}} \equiv V/(4\pi R^3/3))$ as a function of t_0 for radii $R = 0.5, 1, 1.5, 2$
 (top to bottom) in $d = 3$ (left panel) and $d = 4$ (middle and right
 panel) CFT

Observations

- Thermalization is approached as fast as compatible with causality.

Note: speed of light (gluons), not speed of sound (density fluctuations)

For heavy ion collisions this implies

$$\tau \sim 1/(2Q_s) \sim 0.1 \text{ fm}/c$$

Clear contradiction to YM result ?

- Short distances thermalize first, top-down rather than bottom-up thermalization

Unavoidable in the AdS dual theory. A fundamental difference between strong and weak coupling ???

We included fluctuations in one spatial dimension of a scalar field on a 1+2 dimensional boundary and calculated in AdS_4 using perturbative expansions.

$$\phi(t, x) = \epsilon \left(1 + e^{-\mu^2 x^2} \right) e^{-(t-\nu)^2/\sigma^2}$$

We compared with

- free streaming
- second order viscous hydrodynamics

$$\epsilon(x, y) \Rightarrow \text{“local temperature”} \quad \epsilon = \frac{1}{8\pi G_N} \left(\frac{4}{3} \pi T \right)^3$$

$$\text{shear and bulk viscosity} \quad \eta = \frac{1}{16\pi G_N} \left(\frac{4}{3} \pi T \right)^2 \quad \text{and} \quad \zeta = 0$$

⇒ [Hubeny and Rangamani, 1006.3675] shear stress relaxation time $\tau_{\Pi} = \frac{3}{4\pi T} \left(1 + \frac{\gamma}{3} + \frac{\Gamma'(-1/3)}{3\Gamma(-1/3)} \right)$

“We” (i.e. primarily A. Bernamonti and F. Galli) solve

$$D = -u_t \partial_t + u_x \partial_x$$

$$M^{\alpha\beta} = \frac{\partial^{\langle\alpha} u^{\beta\rangle}}{\theta}$$

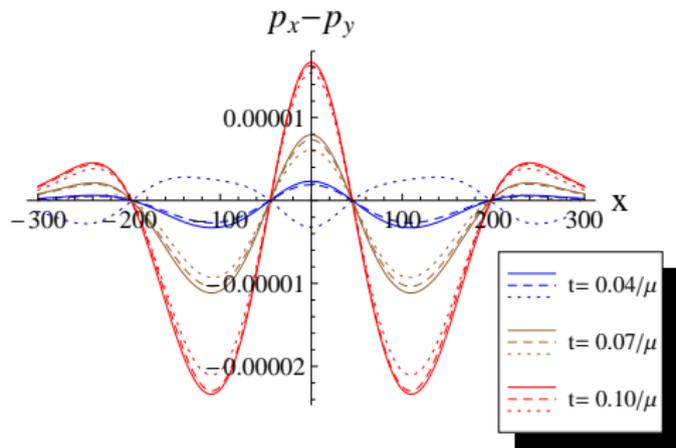
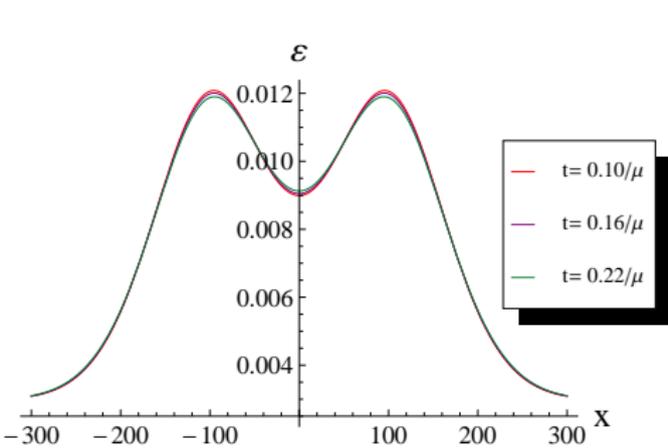
$$\theta = \partial_{\rho} u^{\rho}$$

$$\sigma^{\alpha\beta} = P^{\alpha\rho} P^{\beta\sigma} \left(\partial_{\langle\rho} u_{\sigma\rangle} - \frac{1}{2} P_{\rho\sigma} \theta \right) \quad \text{shear tensor}$$

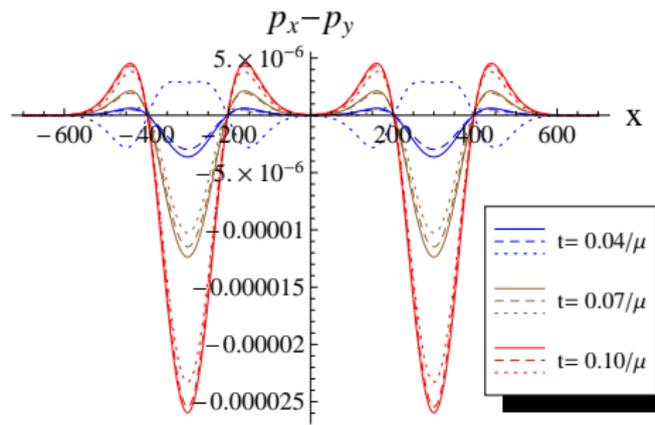
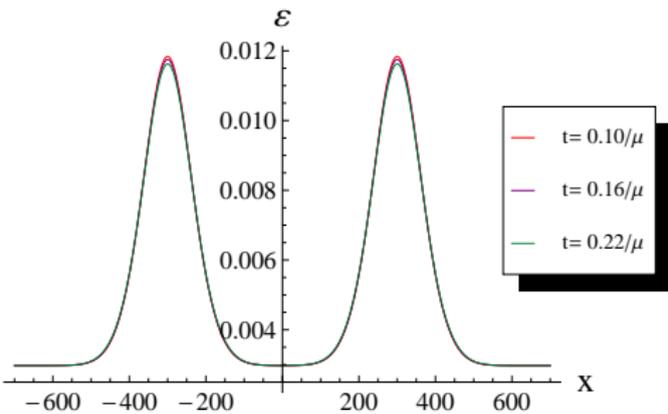
$$D\sigma^{\alpha\beta} = 2(D\theta)M^{\alpha\beta} + 2\theta DM^{\alpha\beta}$$

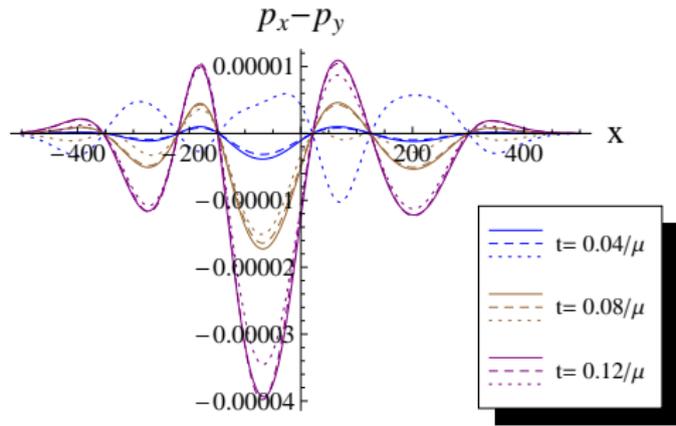
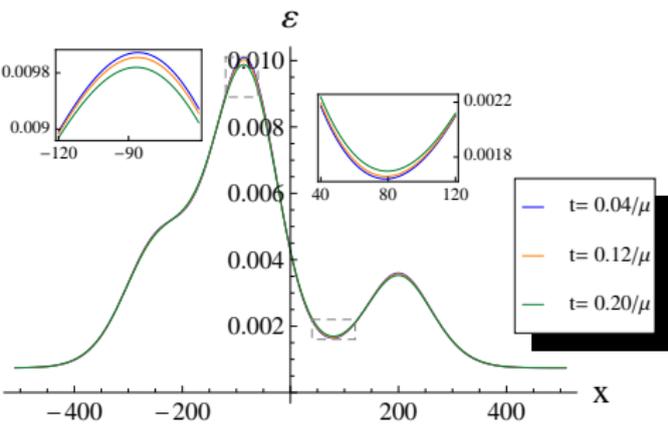
$$\Pi_{(2)}^{\alpha\beta} = \eta\tau_{\Pi} \left[\langle D\sigma^{\alpha\beta} + \frac{1}{2}\sigma^{\alpha\beta}\theta \rangle + \dots \right]$$

From the stress tensor Π one can read off $p_x - p_y$



Conclusion: hydroization is reached long before equilibration





phase	Q1:mechanism	Q2: τ_{therm}	ΔS
1.) pQCD	mixing	0.3 fm/c	30-40 %
2.) cYM	Husimi trafo	2 fm/c	40-50%
2.) AdS/CFT	Bekenstein	0.1 \Rightarrow ? fm/c	40-50 %
3.) hydrodynamics	viscosity	1-2 fm/c	10-20 %
4.) hadron gas	fragmentation	1 fm/c	10 %

Interpretation: “hydroization” time $\ll 1$ fm/c, equilibration time 2 fm/c

The same picture emerges from detailed numerical simulations which are the way to go

⇒ Michal Heller's talk.

Michal gave us a great hands-on introduction into numerical AdS calculations in Regensburg Thank you again !!

Our string collaboration will now also start with numerical simulations

van der Schee, Romatschke, Pratt, Hogg 1307.2539 and 1301.2635

In many aspects this is close to what one would like to do, but no fluctuations \Rightarrow no chance to describe data ?!?

- AdS/CFT analytic: short time expansion \Rightarrow starting condition in local rest frame

$$T^{\mu\nu} = \text{diag}(-\epsilon, P_T, P_T, -1.5P_T)$$

- AdS/CFT numerical solution of Einstein equations, enforcing convergence

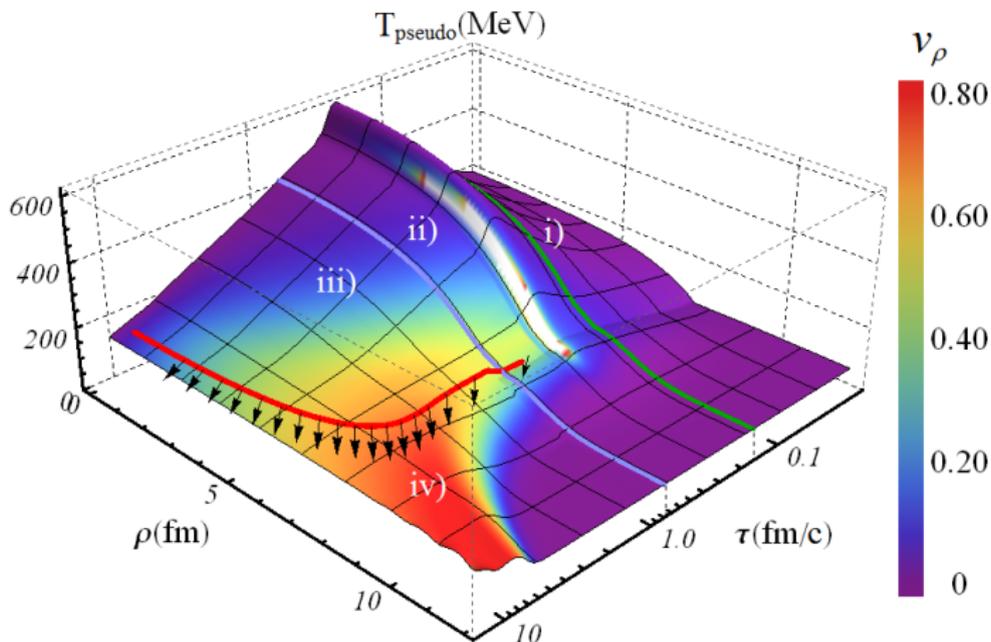
$$ds^2 = -Ad\tau^2 + \Sigma^2 \left(e^{-B-C} d\xi^2 + e^B d\rho^2 + e^C d\theta^2 \right) + 2drd\tau + 2Fd\rho d\tau$$

$$B(r, \tau, \rho) \rightarrow B_0(r, \tau, \rho) + \sum_{i=0}^6 \frac{b_i(\tau, \rho) r^{-i}}{1 + \sigma^7 r^{-7}}$$

- change at some time to a hydro code
- hadronization by MC code describing particle scattering

The general scheme

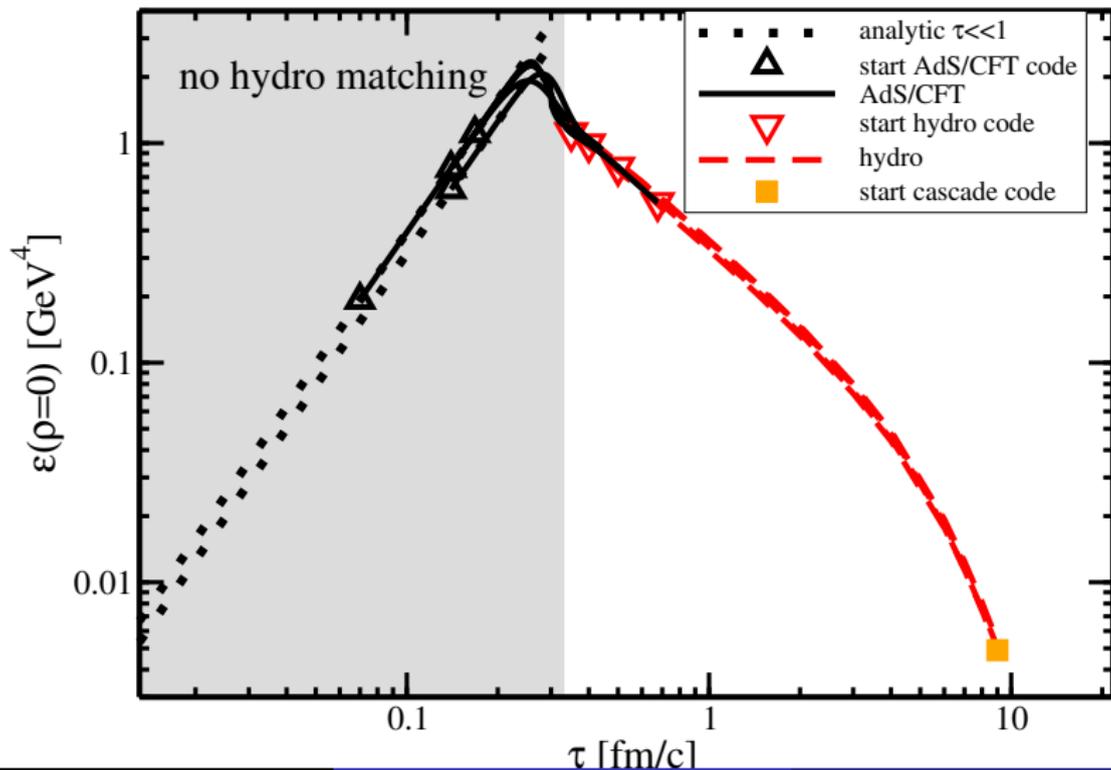
i) early time expansion ii) numerical AdS iii) hydro iv) kinetic theory



Time evolution of energy density at center of fireball for different values of σ , τ_{init} and τ_{hydro} .

Central Energy Density

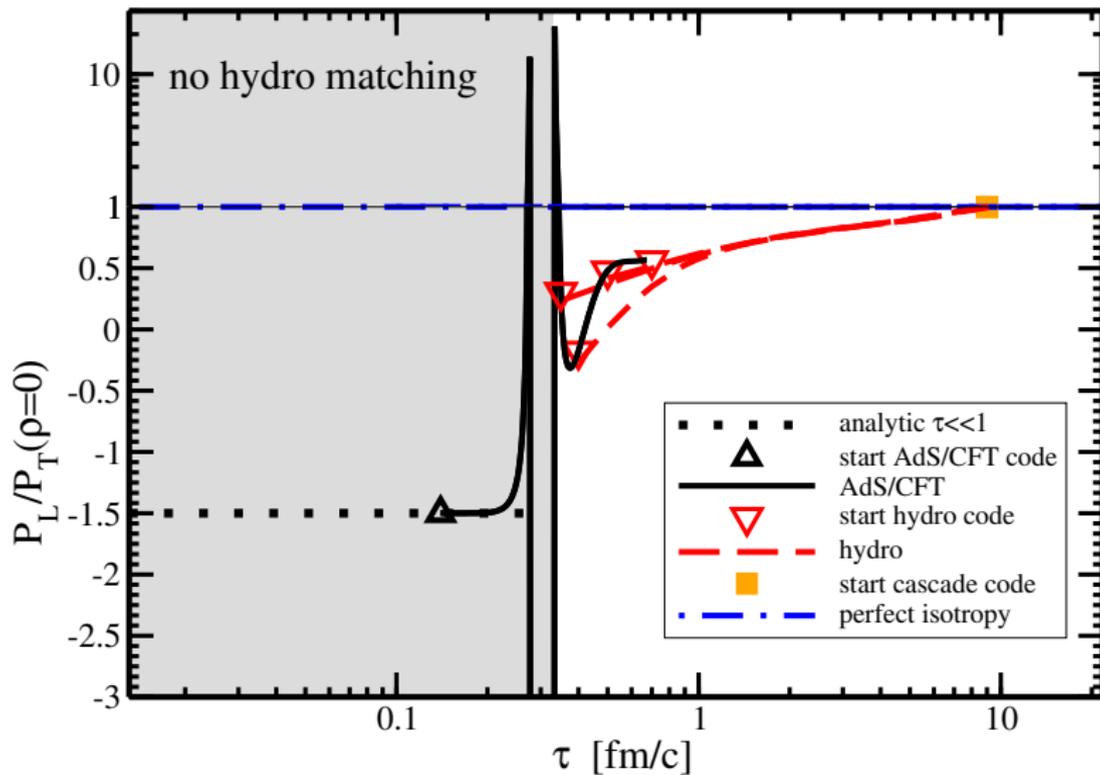
Pb+Pb @ $\sqrt{s} = 2.76$ TeV



P_L/P_T at center of fireball for different values of τ_{hydro}

Pressure Anisotropy

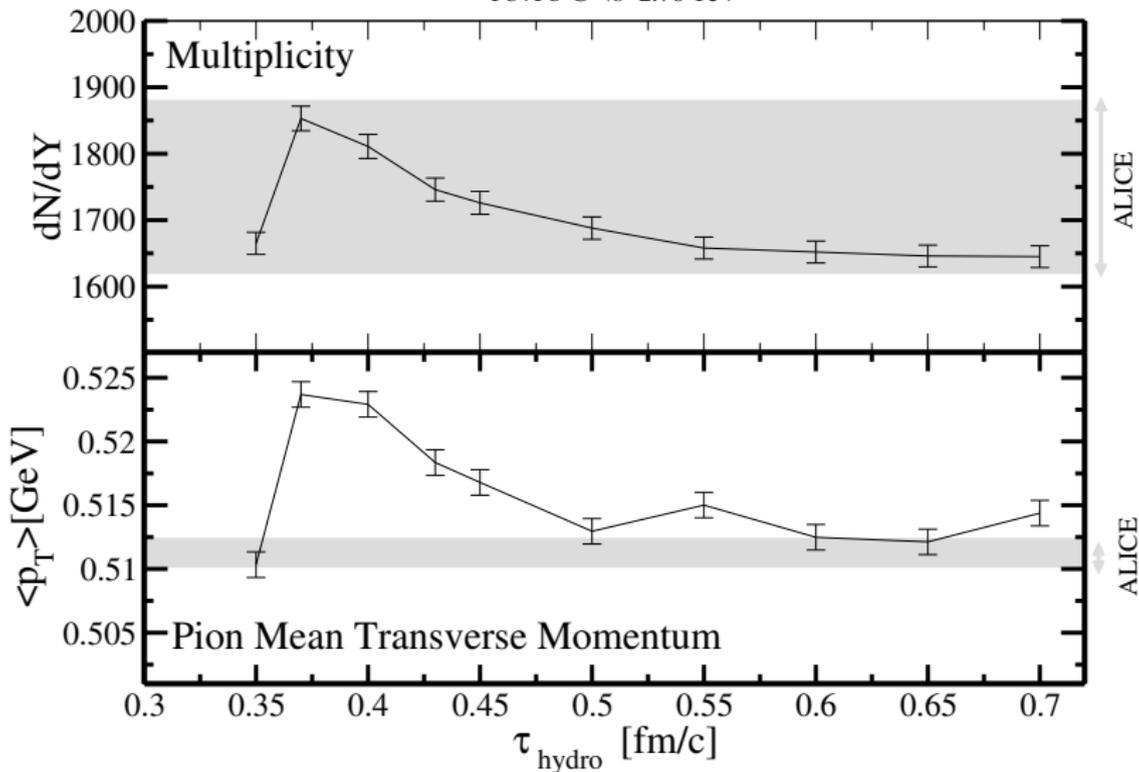
Pb+Pb @ $\sqrt{s} = 2.76$ TeV



Comparison with ALICE data

Effect of choosing τ_{hydro}

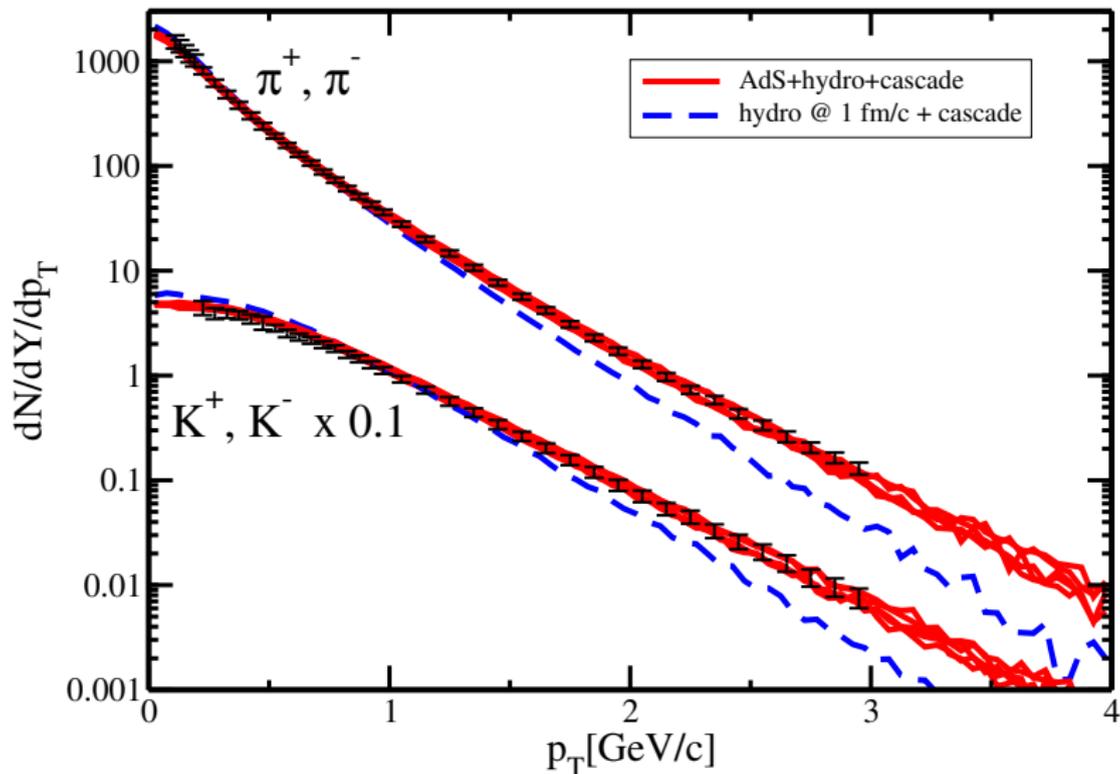
Pb+Pb @ $\sqrt{s}=2.76$ TeV



Comparison with ALICE data

Light Particle Spectra

Pb+Pb @ $\sqrt{s} = 2.76$ TeV



Comments:

The results look a bit too good to be true to me

- It is hard to believe that the instability induced by $P_L/P_T = -1.5$ is free of uncertainties, e.g. ,

$$T_{tt} = \delta(t+z)\epsilon_0 \int_{-\infty}^{\infty} dz \left[1 + e^{(\sqrt{\rho^2+z^2}-R)/a} \right]^{-1}$$

is schematic

- Exact agreement with data without including fluctuations is very surprising

Note: During the oscillatory phase free-streaming is not a valid description in this calculation



This is something we might be able to improve upon

Conclusions

- Understanding entropy production, hydroization and thermalization in HICs is a problem of fundamental importance.
- Thermalization via non-linear dynamics and coarse graining with \hbar , including initial fluctuations, needs $\tau \approx 1 - 2 \text{ fm}/c$.
- Thermalization for strong coupling as described by AdS/CFT is top-down and very fast $\tau \approx 0.1 \text{ fm}/c$. Initial fluctuations change this conclusion.
- Perhaps the initial magnetic field helps, see Gergo's talk
- The practical aim: Obtain precise results for e.g. η

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