Open quantum systems:

From Brownian motion to quantum gravity

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The big ticket questions



Return to the basics (I am not very smart)

Modeling dissipative dynamics: the Langevin equation



Many applications: Brownian motion, Johnson noise, etc.

Dissipation also plays a role in theories at local equilibrium (hydrodynamics)

$$T_{\mu\nu} = \underbrace{(\epsilon + p) u_{\mu}u_{\nu} + p g_{\mu\nu}}_{\text{Ideal fluid}} - \underbrace{\eta \partial_{(\mu}u_{\nu)}}_{\text{Dissipative terms}} - \underbrace{\zeta (\partial \cdot u) (g_{\mu\nu} + u_{\mu}u_{\nu})}_{\text{Dissipative terms}} + \cdots$$

but what about stochastic fluctuation?

Microscopic origin of dissipation

The two situations before are **effective** descriptions.



A solution: open quantum systems

Consider a QM system composed of two distinct parts:



with unitary time evolution $\rho(t) = e^{-iHt} \rho(0) e^{iHt}$. Could be taken to be pure... but don't. The Hilbert space is taken to be of the form

Spanned by
$$\{|x\rangle\}$$

 $\mathcal{H} = \mathcal{H}_A \bigotimes \mathcal{H}_B$
Spanned by $\{|R\rangle\}$

The state at t = 0 can be taken to be a tensor product state

$$\label{eq:rho} \begin{split} \rho(0) = & \rho_A \; (0) \otimes \; \rho_B(0) \\ & \hfill \hfill \\ \end{split}$$

 Very simple (SRE) state: $I(A,B) = S(AB) - S(A) - S(B) = 0$

The matrix elements of $\rho(t)$ in the given basis are

$$\langle x, R | \rho(t) | y, Q \rangle \xrightarrow{K(x, r; t | x', R'; 0)}$$

$$\langle x, R | \rho(t) | y, Q \rangle \xrightarrow{=} \sum_{x' \in \mathcal{X}, R | e^{-iHt} | x', R' \rangle} \langle x', R' | \rho(0) | y', Q' \rangle$$

$$\times \langle y', Q' | e^{iHt} | y, Q \rangle$$

$$\bigwedge_{K^*(y, Q; t | y', Q'; 0)}$$

and the propagators admit a path integral representation

$$K(x,r;t|x',R';0) = \int_{\tilde{x}(0)=x',\tilde{R}(0)=R'}^{\tilde{x}(t)=x,\tilde{R}(t)=R} D\tilde{x}D\tilde{R} e^{iS[\tilde{x},\tilde{R}]}$$

Mind the limits

All together

$$\begin{array}{l} & \overbrace{} \text{Difference between } K \text{ and } K^{*} \\ \hline \\ \langle x, R | \, \rho(t) \, | y, Q \rangle = \int D \tilde{x} D \tilde{R} \, D \tilde{y} D \tilde{R} \, e^{i S[\tilde{x}, \tilde{R}] \, - \, i S[\tilde{y}, \tilde{Q}]} \\ & \times \langle x', R' | \, \rho(0) \, | y', Q' \rangle \end{array}$$

or abusing short-hand

$$\rho(t) = \int_{\rho(0)} DX_R DX_L e^{iS[X_R] - iS[X_L]}$$

$$X_R = \{\tilde{x}, \tilde{R}\}$$

$$X_L = \{\tilde{y}, \tilde{Q}\}$$

We can also use diagrams:



and it is clear also how to compute correlators



This is known as the Schwinger-Keldysh contour.

This story can be extended for higher order correlators



► All this can be "easily" applied to QFT.

Some comments about SK contours

• At least double the DoF $X \rightarrow \{X_R, X_L\}$ (purification).

- ▶ In-in formalism: only input is $\rho(0)$ (out-of-equilibrium).
- Any operator ordering (retarded/advance correlators).

Compare against S-matrix calculations (in-out formalism).

$$\begin{array}{c} t \to \infty \\ & & \\ & & \\ & & \\ & & \\ & & \\ t \to -\infty \end{array} |out\rangle \sim a_i^{\dagger} |0\rangle \\ & & \\ &$$

Of course, we could also have



credit: Anonymous undergraduate.

Partial traces and influence functionals



We do not care about the full density matrix, $\rho(t)$, but just $\rho_A(t) = \text{Tr}_B \rho(t)$ where the environment is ignored^{American style}. In terms of matrix elements:

$$\begin{aligned} \langle x | \rho_A(t) | y \rangle &= \oint dR \, \langle x, R | \rho(t) | y, R \rangle \underbrace{\operatorname{Only A's DoF}}_{= \int dx' dy' \int_{x',y'}^{x,y} D\tilde{x} D\tilde{y} \, e^{iS_A[\tilde{x}] - iS_A[\tilde{y}]} \, \langle x' | \, \rho_A(0) \, | y' \rangle \\ &= \int dx' dy' \int_{x',y'}^{x,y} D\tilde{x} D\tilde{y} \, e^{iS_A[\tilde{x}] - iS_A[\tilde{y}]} \, \langle x' | \, \rho_A(0) \, | y' \rangle \\ & \underbrace{\operatorname{Interaction}}_{\times e^{iS_I[\tilde{x},\tilde{R}] - iS_I[\tilde{y},\tilde{Q}]}} \times \underbrace{\int_{R',Q'}^{R} D\tilde{R} D\tilde{Q} \, e^{iS_B[\tilde{R}] - iS_B[\tilde{Q}]}}_{\langle R' | \, \rho_B(0) \, | Q' \rangle} \end{aligned}$$

where we used $\rho(0) = \rho_A(0) \otimes \rho_B(0)$.

Or in much more compact notation

$$\rho_{A}(t) = \int_{\rho_{A}(0)} Dx_{R} Dx_{L} \exp\left[iS_{A}[x_{R}] - iS_{A}[x_{L}] + iF[x_{R}, x_{L}]\right]$$
Dissipation & noise

where we introduce the influence functional

$$e^{iF[x_R,x_L]} = \int_{\rho_B(0)} DQ_R DQ_L \, e^{iS_B[Q_R] - iS_B[Q_L] + iS_I[x_R,Q_R] - iS_I[x_L,Q_L]}$$

An old friend, rediscovered

Consider a very simple interaction:

$$S_I = \lambda \int_0^t dt' x(t') Q(t')$$

then

$$e^{iF[x_R,x_L]} = \int_{\rho_B(0)} DQ_R DQ_L \, e^{iS_B[Q_R] - iS_B[Q_L]}$$
$$\times e^{i\lambda \int dt'(x_R Q_R - x_L Q_L)}$$
$$= \left\langle e^{i\lambda \int dt'(x_R Q_R - x_L Q_L)} \right\rangle_{\rho_B}$$

The influence functional = generating functional of correlators.

Effective action for Fluctuating Hydro

A special case:
$$\rho_B = \frac{1}{Z(\beta)} e^{-\beta H}$$

- The effect of system A on B is to take it out of equilibrium!
- ▶ The response is captured by $e^{iF[x_R,x_L]}$, in the limit $\omega, |\vec{k}| \ll T.$
- From the influence functional we get $\langle T_{\mu\nu}T_{\rho\sigma}\cdots\rangle$

The only example ever: harmonic oscillators

Let us evaluate the influence functional for:

$$H = \frac{p^2}{2M} + V(x) + \frac{1}{2} \sum_k \left(\frac{P_k^2}{m} + m\omega_k^2 R_k\right) + x \sum_k \lambda_k R_k$$

we then evaluate

$$e^{iF[x_R,x_L]} = \int_{R',Q'}^{R} DQ_R DQ_L e^{\frac{i}{2}m\sum_k \int dt' \left(\dot{Q}_R^2 - \omega_k^2 Q_R^2 + \lambda_k x_R Q_R\right) - (R \to L)} \\ \times \langle R' | \rho_B(0) | Q' \rangle \\ \left(\langle R' | e^{-\beta H} | Q' \rangle = K(R'; -i\beta | Q'; 0) \right)$$



with a lot of insertions along the Lorentzian segments.

Doing the path integral



Final answer:

$$e^{iF[x_R,x_L]} = \exp\left[-\int d\tau ds \left(x_R(\tau) - x_L(\tau)\right) \times \left(\alpha(\tau - s)x_R(s) - \alpha^*(\tau - s)x_L(s)\right)\right]$$

where

$$\alpha(t) = \sum_{k} \frac{\lambda_k}{2m\omega_k} \left[e^{-i\omega_k t} + \frac{e^{-i\omega_k t} + e^{i\omega_k t}}{e^{\beta\omega_k} - 1} \right]$$

It is better to write all this as:

$$e^{iF[x_a, x_d]} = \exp\left[-\int d\tau ds \left(\alpha_R(\tau - s)x_d(\tau)x_d(s)\right) + 2i\alpha_I(\tau - s)x_d(\tau)x_d(s)\right)\right]$$

where

$$x_a \equiv \frac{1}{2} \left(x_R + x_L \right) , \quad x_d \equiv x_R - x_L$$

This term is important: KMS

$$\alpha_{R}(\tau) = \operatorname{Re}\left(\alpha(\tau)\right) = \sum_{k} \frac{\lambda_{k}^{2}}{2m\omega_{k}} \operatorname{coth}\left(\frac{\omega_{k}\beta}{2}\right) \cos\left(\omega_{k}t\right)$$
$$\alpha_{I}(\tau) = \operatorname{Im}\left(\alpha(\tau)\right) = \sum_{k} \frac{\lambda_{k}^{2}}{2m\omega_{k}} \sin\left(\omega_{k}t\right)$$

What it all means? Recovering Brownian motion

Ok... we got a big expression for the propagator of A:

$$J(x,y;t|x'y';0) = \int Dx_R Dx_L e^{iS[x_R] - iS[x_L] - i\int x_d \alpha_I x_a} - \int x_d \alpha_R x_d$$

How can we interpretate it?

Compare with expectation: system under random external force

$$J(x, y; t | x'y'; 0) = \int Dx_R Dx_L \frac{DF}{h} p(F) e^{iS[x_R] - iS[x_L] + i \int x_d F(t)}$$

Distribution of random force (take it gaussian)

Just as before, we change perspective

$$Z_F[x_d] = \int DF \, p(F) \, e^{i \int x_d F}$$
Generating functional of correlators, $\langle F(t_1)F(t_2)\cdots \rangle$

For gaussian distributions $p(F) \propto e^{-F^2/A}$:

$$Z_F[x_d] = e^{-\int d\tau ds \, x_d(\tau) A(\tau-s) x_d(s)}$$

where $\langle F(\tau)F(s)\rangle = A(\tau - s).$

We can compare the two expressions:

Pheno.

$$J(x, y; t | x'y'; 0) = \int Dx_R Dx_L e^{iS[x_R] - iS[x_L] - \int x_d Ax_d}$$

$$J(x, y; t | x'y'; 0) = \int Dx_R Dx_L e^{iS[x_R] - iS[x_L] - i\int x_d \alpha_I x_a - \int x_d \alpha_R x_d}$$

$$Additional part: micro. dynamics$$

and identify:

$$\langle F(t)F(0)\rangle = \alpha_R(t)$$

In the thermodynamic limit

$$\overbrace{\alpha_{R}(t)}^{\omega \ll T} \overbrace{\sim}^{T} \frac{T}{m} \sum_{k} \frac{\lambda_{k}^{2}}{\omega_{k}^{2}} \cos \omega_{k} t \sim \frac{T}{m} \int_{0}^{\infty} \rho(\omega) \frac{\lambda(\omega)^{2}}{\omega^{2}} \cos \omega t$$

$$\boxed{ \text{Large volume} }$$

With the right choice of $\rho(\omega)$:

$$\alpha_R(t) \sim \frac{2\eta T}{\pi} \frac{\sin \Omega t}{t} \sim 2\eta T \delta(t)$$

We recover Langevin from micro. dynamics!

A closer look: what about α_I

Once again, take the view of the enviroment:

$$Z[x_a, x_d] = \left\langle e^{i \int (Q_R x_R - Q_L x_L)} \right\rangle_\beta = \left\langle e^{i \int (Q_d x_a + Q_a x_d)} \right\rangle_\beta$$
$$= e^{-i \int x_d \alpha_I x_a - \int x_d \alpha_R x_d}$$

and taking functional derivatives:

Fluctuating force

$$\langle Q_{a}(t) Q_{a}(0) \rangle = \frac{1}{2} \langle \{Q(t), Q(0)\} \rangle = \frac{1}{i} \frac{\delta}{\delta x_{d}} \frac{1}{i} \frac{\delta}{\delta x_{d}} Z[x_{a}, x_{d}] = \alpha_{R}(t)$$

$$\langle Q_{a}(t) Q_{d}(0) \rangle = \langle [Q(t), Q(0)] \rangle = \frac{1}{i} \frac{\delta}{\delta x_{a}} \frac{1}{i} \frac{\delta}{\delta x_{d}} Z[x_{a}, x_{d}] = i\alpha_{I}(t)$$

Retarded response

Lessons learned

- Influence functional: key object in open quantum systems.
- The x_d x_d coefficient captures the fluctuations (they average to the semi-classical force in Langevin).
- The x_a x_d coefficient gives the response of the environment (dissipation, when conserved).
- Notice, there is no $x_a x_a$ term (unitarity).

Taking stock



Open quantum systems

The gravitational path integral

Old story: gravitational systems can be seen as thermal:



"Newer" story: the holographic GKPW dictionary

$$\left\langle e^{-\int J(x)\mathcal{O}(x)} \right\rangle_{\mathsf{CFT}} = \int_{\mathsf{AdS}} Dg D\phi \, e^{-S_E} \sim e^{-S^{\mathsf{on-shell}}[\phi \to J]}$$

Unpacking GKPW

The generating functional has two pieces of data

$$Z_{\mathsf{CFT}}[J] = \left\langle e^{-\int J(x) \mathcal{O}(x)} \right\rangle_{\rho}$$
State

This has to be encoded in the dynamics on asymptotically AdS spacetimes

$$ds^{2} = \frac{dr^{2}}{r^{2}} + r^{2}g_{\mu\nu}(r,x)dx^{\mu} dx^{\nu}$$

Generic behavior in AdS_{d+1} :

First identification: leading behavior = source

$$\Phi(r,x) = r^{d-\Delta} J(x) \quad (1+\cdots) + r^{-\Delta} \langle \mathcal{O}(x) \rangle_{\rho} \quad (1+\cdots)$$

Second identification: sub-leading behavior = Exp. Val.

- ► The leading behavior is universal → sources are fixed external inputs.
- ► The sub-leading behavior depends on the interior geometry → correlations are dynamical responses (holographic reconstruction).

A special case: the thermal state

- ► Observation: In QFT, Z(β) = Tre^{-βH} can be obtained as a path integral over ℝ^{d-1} × S¹_β.
- ► GKPW: Sum over geometries with boundary condition ℝ^{d-1} × S¹_β.

Dominant saddle: Euclidean cigar ("black hole"):

$$ds^{2} = r^{2} f(r) dt_{E}^{2} + \frac{dr^{2}}{r^{2} f(r)} + r^{2} d\vec{x}_{d-1}^{2}, \quad f(r) = 1 - \frac{r_{+}^{d}}{r^{d}}$$



The Lorentzian GKPW dictionary

GKPW was only formulated for Euclidean geometries.

Long history of Lorentzian generalizations [Herzog, Son, Starinets, Skenderis, van Rees...].

 Main obstacle: Additional data due to time-ordering (causality).

Solution: Lesson from before, the Schwinger-Keldysh contour.

The LHS: what do we want to compute?

From our previous study, the real-time generating functional is:

$$Z[J_R, J_L] = \left\langle P e^{i \int_{\mathbb{R}^{d-1} \times \mathcal{C}_{SK}} (J_R \mathcal{O}_R - J_L \mathcal{O}_L)} \right\rangle_{\beta} =$$


The gravitational SK (grSK) geometry

Generalized GKPW prescription:

$$\left\langle P e^{i \int_{\mathbb{R}^{d-1} \times \mathcal{C}_{SK}} (J_R \mathcal{O}_R - J_L \mathcal{O}_L)} \right\rangle_{\beta} = \int_{g|_{\mathsf{bdy}, \sim} \mathbb{R}^{d-1} \times \mathcal{C}_{SK}} Dg \, e^{iS}$$
$$= e^{iS[g_*]}$$

where g_* is a solution of Einstein's equations with bdy. condition $\mathbb{R}^{d-1} \times \mathcal{C}_{SK}$.

Constructing the solution

- Euclidean insight: dominant saddle in the imaginary-time section is $ds_E^2 = r^2 f(r) dt_E^2 + \frac{dr^2}{r^2 f(r)} + r^2 d\vec{x}_{d-1}^2$.
- Dominant solution in Lorentzian segments: $ds_{\pm}^{2} = -r^{2}f(r)dt_{\pm}^{2} + \frac{dr^{2}}{r^{2}f(r)} + r^{2}d\vec{x}_{d-1}^{2}.$
- Building new solutions by copy-paste (poor's man solution):

$$\left[g^{(i)}_{\mu\nu} - g^{(j)}_{\mu\nu}\right]_{\Sigma} = 0\,, \quad \left[K^{(i)}_{\mu\nu} - K^{(j)}_{\mu\nu}\right]_{\Sigma} = 0\,.$$

- Easy part: pasting to the Euclidean segment.
 Simply take Σ to be the surface t_E = 0 ~ 0 + β, and t₊ ~ 0, t₋ = 0 + iβ.
- Tricky part: pasting of Lorentzian segments across their horizon.

The CGL prescription

To paste the Lorentzian segments, make r complex_{[Glorioso, Crossley,}

Liu]

$$ds^2 = -r^2 f dv^2 + 2dv dr + r^2 d\vec{x}^2$$



We then restrict $r \in \mathbb{C}$ to live on the Hankel contour above.



Credit: Heroic tikz work by Tom Angrick

All in one patch: Mock tortoise coordinate

It is convenient to introduce the mock tortoise coordinate [Jana,

Loganayagam, Rangamani]

$$ds^{2} = -r^{2}fdv^{2} + i\beta r^{2}fdvd\zeta + r^{2}d\vec{x}^{2},$$
$$\frac{d\zeta}{dr} = \frac{2}{i\beta} \frac{1}{r^{2}f(r)},$$
Convinience prefactor Good ol' tortoise coordinate

Main property: $\zeta(r) \propto \log(r - r_+)$, logarithmic branch cut!

The prefactors have been chosen such that



In particular

$$\lim_{r \to \infty + i\epsilon} \zeta(r) = 0, \quad \lim_{r \to \infty - i\epsilon} \zeta(r) = 1.$$

Scalar dynamics in grSK

To test the prescription, consider

$$S = -\frac{1}{2} \int d^{d+1}x \sqrt{-g} g^{AB} \partial_A \Phi \partial_B \Phi$$

For $r \to \infty \pm i \epsilon$, nothing really changes

$$\Phi(r,k) \sim J_{R/L}(k) (1+\cdots) + \frac{\phi_{R/L}^{(1)}(k)}{r^d} (1+\cdots) , \quad k = (\omega, \vec{k})$$



Across the horizon $r - r_+ \rightarrow (r - r_+)e^{2\pi i}$.

Continuity of the full solution then requires:

$$A_R(k) = A_L(k)$$
, $B_R(k) = B_L(k)e^{-\beta\omega}$

~

It is convenient to write linear combinations, a fully regular solution:

$$\Phi_{\rm in}(r,k) \equiv \Phi_R(r,k) - e^{-\beta\omega} \Phi_L(r,k)$$

$$\sim \begin{cases} A(k)(1+\cdots) & r \to r_+ \\ -(1+n_\omega)J_R(k) + n_\omega J_L(1+\cdots) & r \to \infty + i\epsilon \end{cases}$$

and a fully irregular one:

$$\Phi_{\text{out}}(r,k) \equiv \Phi_R(r,k) - \Phi_L(r,k)$$

$$\sim \begin{cases} B(k) \left(r - r_+\right)^{\frac{i\beta\omega}{2\pi}} \left(1 + \cdots\right) & r \to r_+ \\ -n_\omega \left(J_R(k) - J_L(k)\right) \left(1 + \cdots\right) & r \to \infty + i\epsilon \end{cases}$$

where $n_{\omega} = \frac{1}{e^{\beta\omega}-1}$ is the familiar Boltzmann factor. Notice a

different linear combination produces:

$$J_a = \frac{1}{2} (J_R + J_L) , \qquad J_d = J_R - J_L .$$

Solving the EoM

With the boundary conditions sorted out, we can solve

$$-\frac{1}{\sqrt{-g}}\partial_A\left(\sqrt{-g}g^{AB}\partial_B\Phi\right) = 0\,,$$

or in Fourier domain

-

$$\frac{1}{r^{d-1}}\mathbb{D}_+\left(r^{d-1}\mathbb{D}_+\Phi(r,k)\right) + \left(\omega^2 - k^2f\right)\Phi(r,k) = 0.$$

where $\mathbb{D}_{+} = r^{2} f \frac{d}{dr} - i\omega$ is motivated by time-reversal.

We define the ingoing propagator $G^{\text{in}}(r,k)$ as a solution satisfying:

$$\lim_{r \to \infty \pm i\epsilon} G^{\text{in}}(r,k) = 1 , \quad \lim_{r \to r_+} G^{\text{in}}(r,k) = \text{ regular}$$

$$\int \text{Due to regularity, it's the same in both branches}$$

Using time-reversal symmetry, $v \rightarrow i\beta\zeta - v$, the *outgoing* propagator is

$$G^{\text{out}}(r,\omega,\vec{k}) = e^{-\beta\omega\zeta} G^{\text{in}}(r,-\omega,\vec{k})$$

Fully irregular, $\zeta \to 0,1$

The general solution to the EoM in grSK geometry is

$$\Phi(\zeta, \omega, k) = \left(J_a + \left(n_\omega + \frac{1}{2}\right)J_d\right)G^{\text{in}}(r, \omega, \vec{k}) - n_\omega J_d e^{\beta\omega(1-\zeta)}G^{\text{in}}(r, -\omega, k)$$

Notice, we *only* need to compute $G^{in}(r, \omega, k)!$

A quick check:

$$\lim_{\zeta \to 0} \Phi(\zeta, \omega, k) = \left(J_a + \left(n_\omega + \frac{1}{2}\right)J_d\right) - n_\omega e^{\beta\omega}J_d = J_L$$
$$\lim_{\zeta \to 1} \Phi(\zeta, \omega, k) = \left(J_a + \left(n_\omega + \frac{1}{2}\right)J_d\right) - n_\omega J_d = J_R$$

Evaluating the on-shell action

With the solution on hand, we evaluate:

where

$$\Pi(r,k) \equiv -\sqrt{-g}g^{rB}\partial_B \Phi = -r^{d-1}\mathbb{D}_+\Phi$$

Defining

$$K^{\mathsf{in}}(\omega,k) = -\lim_{r \to \infty} \left(r^{d-1} \mathbb{D}_+ G^{\mathsf{in}}(r,\omega,k) + \mathsf{c.t.} \right)$$

We find

$$S_{\text{on-shell}} = -\frac{1}{2} \int_{k} J_{d}^{\dagger}(k) \, K^{\text{in}}(\omega,k) \left(J_{a}(k) + \left(n_{\omega} + \frac{1}{2} \right) J_{d}(k) \right) + \text{c.c.}$$

which, by GKPW, gives us the generating functional of connected correlators:

$$iW[J_a, J_d] = \log Z[J_a, J_d] \sim iS_{\text{on-shell}}[J_a, J_d]$$

A reassuring answer

We can compare:

$$S_{\text{on-shell}} = -\frac{1}{2} \int_{k} J_{d}^{\dagger}(k) \, K^{\text{in}}(\omega, k) \left(J_{a}(k) + \left(n_{\omega} + \frac{1}{2} \right) J_{d}(k) \right) + \text{c.c.}$$

with the harmonic oscillator answer

$$\log Z[x_a, x_d] = -i \int x_d \,\alpha_I \,x_a - \int x_d \,\alpha_R \,x_d$$

It is the same structure!

From the H.O. analysis, the J[†]_dJ_d coefficient gives the average force:

$$\langle \mathcal{O}(-k)\mathcal{O}(k) \rangle^{\text{Kel.}} = -\frac{1}{2} \operatorname{coth}\left(\frac{\beta\omega}{2}\right) \operatorname{Im} K^{\operatorname{in}}(\omega, \vec{k})$$

Same factor we found in the H.O. analysis!

• While the $J_d^{\dagger} J_a$ coefficient gives the retarded response:

$$\langle \mathcal{O}(-k)\mathcal{O}(k)\rangle^{\mathsf{Ret.}} = iK^{\mathsf{in}}(\omega,\vec{k})$$

The two results are related:

$$\langle \mathcal{O}(-k)\mathcal{O}(k)\rangle^{\mathsf{Kel.}} = \frac{1}{2} \operatorname{coth}\left(\frac{\beta\omega}{2}\right) \operatorname{Re}\left\langle \mathcal{O}(-k)\mathcal{O}(k)\right\rangle^{\mathsf{Ret.}}$$

which is the fluctuation-dissipation theorem.

- ▶ Notice there is no $J_a^{\dagger}J_a$ term, as expected from unitarity.
- Our gravitational calculation agrees completely with the QM expectation.

All together: Holographic open quantum systems

Observer/measuring system

$$S = S_A + S_B + \int d^d x J(x) \mathcal{O}(x)$$

Holographic system, for example $\mathcal{N}=4$ SYM

$$\begin{split} \text{Influence functional:} & \underset{\text{Fancy-pants harmonic oscillator}}{\text{Fancy-pants harmonic oscillator}} \\ \left\langle e^{i\int(J_R\mathcal{O}_R-J_L\mathcal{O}_L)} \right\rangle_{\beta}^{\mathcal{N}=4\,\text{SYM}} = \int Dg D\Phi\, e^{iS[g,\Phi]} \sim e^{iS_{\text{on-shell}}[J_R,J_L]} \big|_{\text{grSK}} \end{split}$$

[Jana,Loganayagam,Rangamani][Loganayagam,Rangamani,JV]

The real deal: metric fluctuations

Going beyond toy scalars:

$$\begin{split} \left\langle e^{i\int (J_R\mathcal{O}_R - J_L\mathcal{O}_L)} \right\rangle_{\beta}^{\mathcal{N} = 4\,\text{SYM}} &\sim e^{iS_{\text{on-shell}}} \left(\text{metric fluctuations}\right) \\ &= e^{iS_{\text{on-shell}}} \int Dh \, e^{iS^{(2)}[h] + iS^{(3)}[h] + \cdots} \end{split}$$

where $g = g_{grSK} + h$, with $|h| \ll 1$, and $S^{(n)}[h]$ come from expanding the Einstein-Hilbert action.

Humble beginnings: gaussian fluctuations

Start with only

Gaussian integration

$$Z_2[h] = \int Dh \, e^{iS^{(2)}[h]} \stackrel{\downarrow}{=} e^{iS^{(2)}_{\text{on-shell}}[h|_{\text{bdy.}}]}$$

Generic behavior: $h_{AB}(r,x)dx^{A}dx^{B} \sim \frac{dr^{2}}{r^{2}} + r^{2} \left(\begin{array}{c} \gamma_{\mu\nu}(x) + \frac{\langle T_{\mu\nu}(x) \rangle}{r^{d}} \end{array} \right) dx^{\mu}dx^{\nu} + \cdots$ Source

Metric fluctuations as open quantum systems

Comparing with the scalar results:

 $\left\langle e^{i \int \left(\gamma_{\mu\nu}^{R} T_{R}^{\mu\nu} - \gamma_{\mu\nu}^{L} T_{L}^{\mu\nu}\right)} \right\rangle_{\beta}^{\mathcal{N}=4 \, \text{SYM}} = e^{i S_{\text{on-shell}}^{(2)}[h|_{\text{bdy.}} \sim \gamma]} + \cdots$

Perturbation around grSK geometry

Connection with Hydrodynamics: $\langle T_{\mu\nu}T_{\rho\sigma}\cdots\rangle$ in the limit

 $\omega, |\vec{k}| \ll T.$

Handling metric fluctuations: designer scalars

 $g = g_{\text{grSK}} + h$ SO(d-2) harmonic descomposition Diff. invariance Constraint equations

$$S_{\mathcal{M}} = -\frac{1}{2} \int d^{d+1}x \sqrt{-g} e^{\chi(r)} \nabla_A \Phi_{\mathcal{M}} \nabla^A \Phi_{\mathcal{M}}, \quad \sqrt{-g} e^{\chi(r)} \sim r^{\mathcal{M}}$$

 $[{\sf Ghosh}, {\sf Loganayagam}, {\sf Prabhu}, {\sf Rangamani}, {\sf Sivakumar}, {\sf Vishal}] [{\sf He}, {\sf Loganayagam}, {\sf Rangamani}, {\sf Sivakumar}, {\sf JV}]$

The dilaton makes all the difference

Tensors:
$$e^{\chi(r)} = 1$$
.
Vectors: $e^{\chi(r)} = \frac{1}{r^{2(d-1)}}$,
Scalars: $e^{\chi(r)} = \frac{1}{r^{2(d-2)}\Lambda_k^2}$, $\Lambda_k = k^2 + \frac{1}{2}(d-1)r^3f'$.
 e^{χ}
 e^{χ}
 $non-Markovian$
 $Markovian$
 r_+

Dynamic of designer scalars in SK

Let us study the system (scalars are hard):

$$S_{\text{designer}} = -\frac{1}{2} \int d^{d+1}x \sqrt{-g} \, r^{\mathcal{M}-d+1} \, \nabla_A \Phi_{\mathcal{M}} \nabla^A \Phi_{\mathcal{M}} + S_{\text{bdy.}}$$

The equation of motion is

$$r^{-\mathcal{M}}\mathbb{D}_+\left(r^{\mathcal{M}}\mathbb{D}_+\Phi_{\mathcal{M}}\right) + (\omega^2 - k^2 f)\Phi_{\mathcal{M}} = 0.$$

For d > 2, this equation can be solved in a gradient expansion:

$$\frac{\omega}{r_+}, \frac{k}{r_+} \ll 1$$



$$\Phi_{\mathcal{M}} \sim c_1 + \frac{c_2}{r^{\mathcal{M}+1}}$$

For $\mathcal{M} > -1$, we impose Dirichlet conditions ($S_{bdy.} = 0$).

For $\mathcal{M} < -1$, compute the conjugate momentum:

$$\pi_{\mathcal{M}} = -r^{\mathcal{M}} \mathbb{D}_{+} \Phi_{\mathcal{M}} \sim \tilde{c}_{1} + \tilde{c}_{2} r^{\mathcal{M}}$$

• Quantization is done using Neumann conditions $(S_{bdy.} = \int d^d x \Phi_{\mathcal{M}} \pi_{\mathcal{M}}).$

Markovian fields (tensors)

Same old scalar:

$$\begin{split} \Phi_{\mathcal{M}}(\zeta, w, \vec{k}) &= J_a G_{\mathcal{M}}^{\text{in}} + \left[\left(n_{\beta} + \frac{1}{2} \right) G_{\mathcal{M}}^{\text{in}} - n_{\beta} e^{\beta(\zeta-1)} G_{\mathcal{M}}^{\text{rev}} \right] J_d \\ S_{\text{on-shell}}[J_a, J_d] &= -\int_k J_d^{\dagger} \mathcal{K}_{\mathcal{M}}^{\text{in}} \left[J_a + \left(n_{\beta} + \frac{1}{2} \right) J_d \right] \\ K_{\mathcal{M}}^{\text{in}} &= -iw + \frac{k^2}{1 - \mathcal{M}} + \Delta_{\mathcal{M}}^{2,0}(r_+)w^2 + \dots \end{split}$$

Note: Correlations are analytic in ω and $|\vec{k}| \Rightarrow$ there are no conserved tensor-like currents!

Non-Markovian fields (vectors and scalars^{*})

Here things are trickier

▶ We could compute log Z[J_a, J_d] = iS_{on-shell}[J_a, J_d], using Newmann conditions.

Alternatively, we could still use Dirichlet

$$\Phi_{\mathcal{M}}(\zeta, w, \vec{k}) = \check{\Phi}_a G_{\mathcal{M}}^{\mathsf{in}} + \left[\left(n_\beta + \frac{1}{2} \right) G_{\mathcal{M}}^{\mathsf{in}} - n_\beta e^{\beta(\zeta-1)} G_{\mathcal{M}}^{\mathsf{rev}} \right] \check{\Phi}_d$$

fixing the one-point function (very non-kosher).

Old friend: Quantum Effective Action

What does it mean to fix the one-point function?

Gravity side: changing from Newmann to Dirichlet:

Original action: Newmann conditions

$$\tilde{S} = -\frac{1}{2} \int d^{d+1}x \sqrt{-g} r^{\mathcal{M}-d+1} \nabla_A \Phi_{\mathcal{M}} \nabla^A \Phi_{\mathcal{M}} + \int d^d x \Phi_{\mathcal{M}} \pi_{\mathcal{M}} - \int d^d x \Phi_{\mathcal{M}} \pi_{\mathcal{M}}$$

Legendre transform to Dirichlet conditions



$$i\Gamma[\check{\Phi}_{a},\check{\Phi}_{d}] = \log Z[J_{a},J_{d}] - i \int d^{d}x \left(J_{R}\check{\Phi}_{R} - J_{L}\check{\Phi}_{L}\right)$$

∠_1

 The Legendre transform of the generating functional is the Quantum Effective Action.

$$\begin{split} \Gamma[\check{\Phi}_a,\check{\Phi}_d] &= S_{\text{on-shell}}^{\text{Dirichlet}}[\check{\Phi}_a,\check{\Phi}_d] \\ &= -\int_k \check{\Phi}_d^{\dagger} \mathcal{K}_{-\mathcal{M}}^{\text{in}} \left[\check{\Phi}_a + \left(n_{\beta} + \frac{1}{2}\right)\check{\Phi}_d\right] \end{split}$$

Interpreation: Wilsonian Influence Functional

Before, influence functional:

$$e^{iF[x_R,x_L]} = \int_{\rho} DQ_R DQ_L \, e^{iS_R - iS_L}$$

Now, Wilsonian Influence Functional

$$e^{iF[x_R,x_L]} = \int_{\rho} \left(\begin{array}{c} DQ_R^{SL} DQ_L^{SL} \end{array} \right) \left(\begin{array}{c} DQ_R^{LL} DQ_L^{LL} \end{array} \right) e^{iS_R - iS_L} \\ = \int \left(DQ_R^{SL} DQ_L^{SL} \right) e^{iS_{\text{WIF}}} \\ \uparrow \end{array}$$

LT wrt Long-Lived modes \Rightarrow QEA \Rightarrow Fluid EFT



$$\begin{split} \mathcal{S}_{\text{WIF}} \propto & -\sum_{\alpha=1}^{N_{V}} \int_{k} \left(\check{\mathcal{P}}_{d}^{\alpha} \right)^{\dagger} K_{\text{Mom.}}^{\text{in}} \left[\check{\mathcal{P}}_{a}^{\alpha} + \left(n_{\beta} + \frac{1}{2} \right) \check{\mathcal{P}}_{d}^{\alpha} \right] \\ & - \int_{k} \left(\check{\mathcal{Z}}_{d} \right)^{\dagger} K_{\text{Sound}}^{\text{in}} \left[\check{\mathcal{Z}}_{a} + \left(n_{\beta} + \frac{1}{2} \right) \check{\mathcal{Z}}_{d} \right] \end{split}$$

$$-i\omega + \frac{1}{\frac{1}{d}}k^2 + \cdots$$

$$\begin{split} K_{\text{Sound}}^{\text{in}} = & -\omega^2 + \frac{k^2}{d-1} + \nu_s k^2 \Gamma_s(\omega,k) \\ & \\ & \text{Hawking sound} \end{split}$$

 $K_{\mathsf{Mom}}^{\mathsf{in}} =$

Beyond the Gaussian level:

Witten diagrams in SK geometry

What about $S^{(n)}[h]$ for n > 1



The ingredients:

- ▶ Ingoing Bulk-to-Bdy Prop.: $G_{in}(\zeta, k)$.
- Outgoing Bulk-to-Bdy Prop.: $G_{out}(\zeta, k) = e^{-\beta\omega\zeta}G_{in}(\zeta, \bar{k})$
- Bulk-to-Bulk Prop:

 $G_{\rm bb}(\zeta,\zeta';k) = \mathcal{N}(k)e^{\beta\omega\zeta'}G_{\rm L}(\zeta_>,k)G_{\rm R}(\zeta'_<,k).$

Important: $G_{in}(\zeta + 1, k) = G_{in}(\zeta, k).$

Contact diagram:

$$= \oint d\zeta \mathfrak{L}(\zeta) = \int_{r_{\mathsf{H}}}^{r_c} dr \left(\mathfrak{L}(\zeta(r) + 1) - \mathfrak{L}(\zeta(r)) \right)$$

• Exchange diagram $= \oint d\zeta \oint d\zeta' \left[F_1(\zeta,\zeta')\Theta(\zeta-\zeta') + F_2(\zeta,\zeta')\Theta(\zeta'-\zeta)\right]$ $= \int_{r_{\rm H}}^{r_c} dr \int_{r_{\rm H}}^{r_c} dr' \left[\mathfrak{F}_1\theta(r-r') + \mathfrak{F}_2\theta(r'-r)\right]$

where

$$\mathfrak{F}_1 = F_1(\zeta, \zeta') - F_1(\zeta + 1, \zeta') + F_2(\zeta + 1, \zeta' + 1) - F_2(\zeta, \zeta' + 1),$$

$$\mathfrak{F}_2 = F_1(\zeta + 1, \zeta' + 1) - F_1(\zeta + 1, \zeta') + F_2(\zeta, \zeta') - F_2(\zeta, \zeta' + 1).$$
Taking stock, part II



Work in progress

- Higher order SK correlators for d > 2 [Ammon, Rangamani, Specht, JV].
- ► Connections to non-linear fluid actions [Rangamani, JV].
- Beyond the SK contour: OTOC [Ammon, Germerodt, Sieling, JV].

Final musings

 More complicated thermal systems: Kerr black holes, superfluids,...

(For charged black holes see [He, Loganayagam, Rangamani, JV])

- Transition amplitudes: beyond saddle point approximation (Golden dream: topology changes)
- Non local-probes: entanglement entropy, reflected entropy, negativity,... [Colin-Ellerin, Dong, Marolf, Rangamani, Wang][Pelliconi,



 Gravitational SK contours for more general spacetimes (see recent work on dS correlators[Di Pietro, Gorbenko,

Komatsu][Loganayagam, Shetye]).



Thank You!