



UNIVERSITÄT
LEIPZIG

Charged scalar field in (near-extremal) Reissner-Nordström spacetime

19 Sept 2023

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Physik-Combo MPI MiS

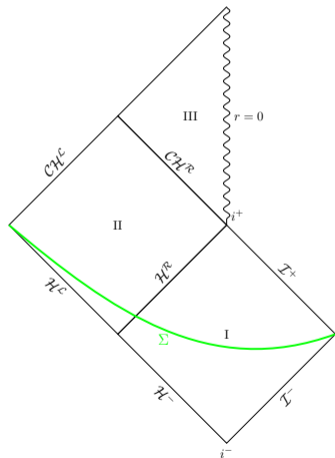
Outline

1. Overview and motivation
 2. Setup
 - 2.1 Background geometry
 - 2.2 Quantum theory
 3. Scattering problem
 - 3.1 Interior region
 - 3.2 Exterior region
 4. Results
 5. Conclusions and outlook
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Overview and motivation

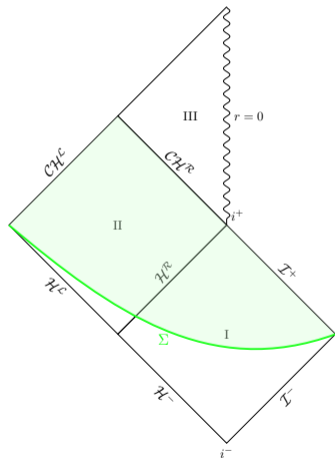
Strong cosmic censorship

- BH interiors \rightarrow determinism?



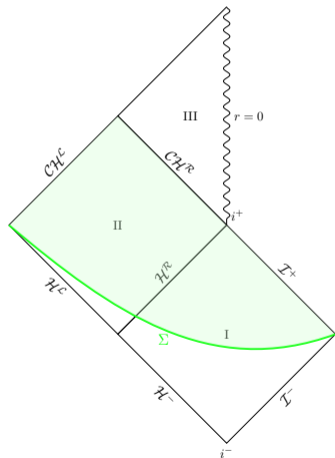
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 - Cauchy hypersurface Σ
 - Domain of dependence $D^+(\Sigma)$
- Strong cosmic censorship
 - Original formulation [Penrose 1965]
 - Weaker formulation [Christodolou 1991]



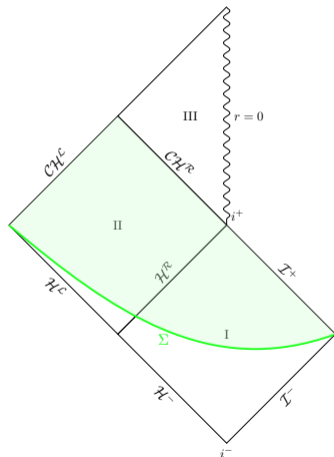
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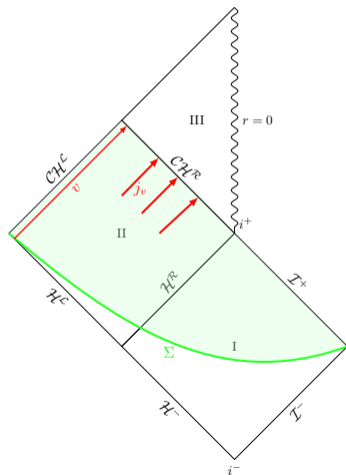
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 - Classically: violated in near-extremal RNdS [Cardoso et al. 2018], [Dias, Reall, and Santos 2019]
 - Quantum effects restore sCC [Klein, Zahn, and Hollands 2021], [Hollands, Wald, and Zahn 2020]



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[Cardoso et al. 2018], [Dias, Reall, and Santos 2019]
 - Quantum effects restore sCC [Klein, Zahn, and Hollands 2021], [Hollands, Wald, and Zahn 2020]
- (weak) backreaction: $\partial_v Q(u, v) = -4\pi r^2 \langle j_v \rangle_U$
 \implies (dis)charge BH interior RNdS



QFTCS: an overview

- Quantum fields on curved background (classical) spacetimes
- Semiclassical Einstein-Maxwell equations

$$G_{\mu\nu} = 8\pi(\langle T_{\mu\nu} \rangle_U + T_{\mu\nu}^{EM})$$

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1. Fix the background geometry \implies Reissner-Nordström ST
2. Compute EVs of observables of a certain field in a (Hadamard) state of choice
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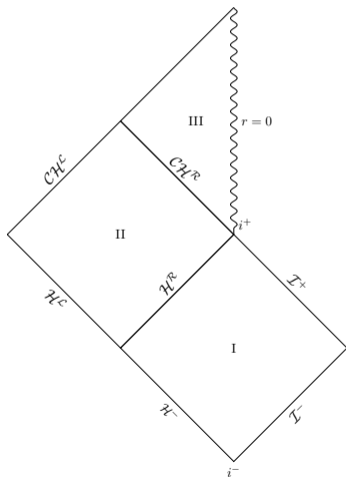
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- Analytical analysis of real scalar in (near-extremal) RN [Zilberman and Ori 2021]
 - My MSc. project: **can this be generalized to the charged scalar?**

Setup

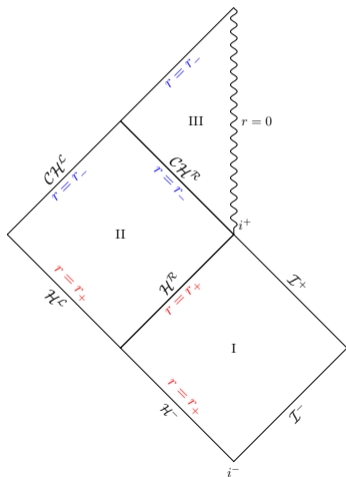
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– Reissner-Nordström ST

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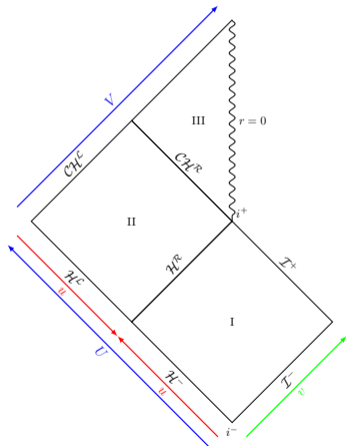
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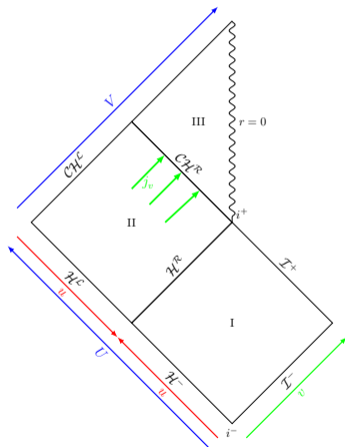
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- Lightlike coordinates $u = t - r_*$, $v = t + r_*$
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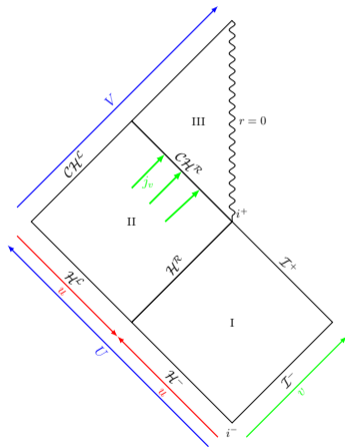
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- Near-extremality $\Delta = \sqrt{1 - Q^2/M^2} \ll 1$

Setup: Quantum theory

– Charged scalar field: $D_\mu D^\mu \Phi = 0$

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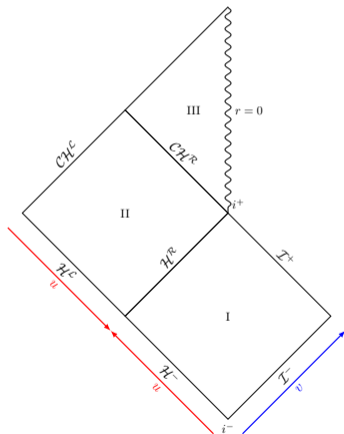
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- Fix

$$\chi(t, r_*, \theta, \phi) = \frac{qQ}{r_0} t, \quad r_0 \in \{r_+, r_-\}$$

Setup: Quantum theory

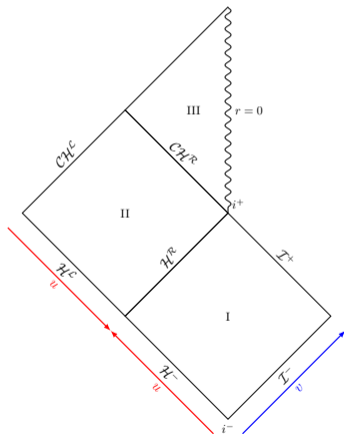


- Quantization: expansion in modes $\Phi_{\omega l}$

$$\Phi(x) = \sum_{\lambda, l, m} \int_0^{\infty} d\omega \left(\Phi_{\omega l m}^{\lambda}(x) a_{\omega l m}^{\lambda} + \Phi_{-\omega l m}^{\lambda}(x) b_{\omega l m}^{\lambda \dagger} \right)$$

- $\Phi_{\omega l m}$: positive frequency Unruh mode solutions

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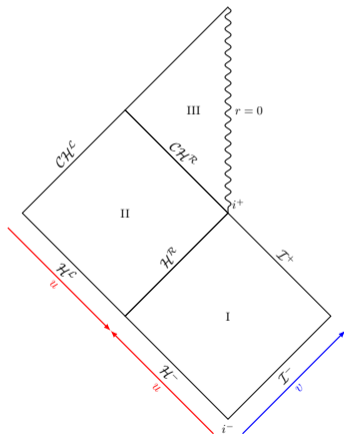


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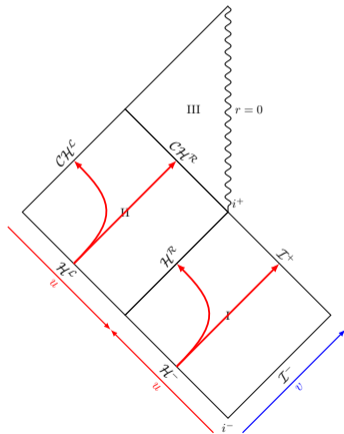


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$$\begin{array}{ll}
 h_{\omega l}^{\text{in}, I} \sim e^{-i\omega v} & \text{on } \mathcal{I}^- \\
 h_{\omega l}^{(+)\text{in}, II} \sim e^{-i\omega v} & \text{on } \mathcal{H}^R \\
 h_{\omega l}^{(+)\text{up}, I} \sim e^{-i\omega u} & \text{on } \mathcal{H}^- \\
 h_{\omega l}^{(+)\text{up}, I} \sim e^{-i\omega u} & \text{on } \mathcal{H}^L
 \end{array}$$

Setup: Quantum theory

- Wronskian relation in region I:

$$|R_{\omega l}^{\text{up},I}|^2 + \frac{\omega - \omega_I}{\omega} |T_{\omega l}^{\text{up},I}|^2 = 1, \quad \omega_I = \frac{qQ}{r_+}.$$

Setup: Quantum theory

- Wronskian relation in region I:

$$|R_{\omega l}^{\text{up},l}|^2 + \underbrace{\frac{\omega - \omega_l}{\omega}}_{<0} |T_{\omega l}^{\text{up},l}|^2 = 1, \quad \omega_l = \frac{qQ}{r_+}$$

\implies superradiance $\implies |R_{\omega l}^{\text{up},l}|^2 > 1 \implies \langle j_\nu \rangle_U^{\mathcal{H}^R, \text{ren}} \neq 0$ even at extremality

Scattering problem



The current j_μ and the SET $T_{\mu\nu}$

– Classically
$$j_\nu = iq \left(\Phi (D_\nu \Phi)^* - \Phi^* D_\nu \Phi \right)$$

$$T_{\mu\nu} = \frac{1}{2} \left((D_\mu \Phi)^* D_\nu \Phi + D_\mu \Phi (D_\nu \Phi)^* \right) - \frac{1}{4} g_{\mu\nu} g^{\rho\lambda} \left((D_\rho \Phi)^* D_\lambda \Phi + D_\rho \Phi (D_\lambda \Phi)^* \right)$$

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– Literature:

– Hadamard point-split [Klein and Zahn 2021] $\rightarrow \langle j_\mu(x) \rangle_U \checkmark$

– Comparison state [Klein, Zahn, and Hollands 2021] $\rightarrow \langle T_{\nu\nu} \rangle_{U-C}^{IH} \checkmark$

– Particular components do not require renormalization [Balakumar, Bernar, and Winstanley 2022]

$\rightarrow \langle j_{r_*}(x) \rangle_\Psi, \langle T_{tr_*}(x) \rangle_\Psi \checkmark$

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– Essential ingredient: scattering coefficients of Boulware modes in regions I, II

$$\frac{d^2 h_{\omega\ell}}{dr_*^2} = \left\{ f(r) \left(\frac{l(l+1)}{r^2} + \frac{\partial_r f(r)}{r} \right) - \left(\omega - \frac{qQ}{r} \right)^2 \right\} h_{\omega\ell}. \quad (1)$$

Interior region $r \in (r_-, r_+) \iff r_* \in (+\infty, -\infty)$

– (Leading order) radial equation of motion

$$\frac{d^2 h_{\omega l}^{l(+)} }{d\tilde{r}_*^2} = \left\{ -l(l+1) \operatorname{sech}^2 \tilde{r}_* - \left(\tilde{\omega} - qQ(1 + \tanh \tilde{r}_*) \right)^2 + \mathcal{O}(\Delta) \right\} h_{\omega l}^{l(+)} . \quad (2)$$

– BCs:
$$h_{\omega l}^{in, l} = \begin{cases} e^{-i\omega r_*} & r_* \rightarrow -\infty \\ T_{\omega l}^{in, l} e^{-i(\omega - \omega_H) r_*} + R_{\omega l}^{in, l} e^{i(\omega - \omega_H) r_*} & r_* \rightarrow \infty \end{cases}$$

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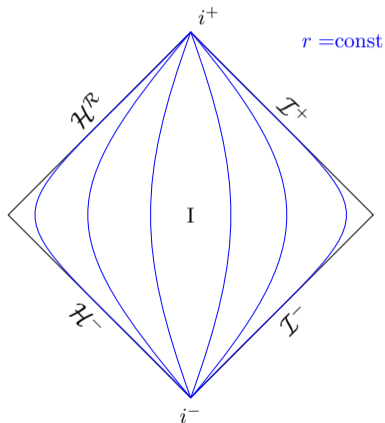
- Analytic solution to leading order in $\Delta \implies T_{\omega l}^{in,ll}, R_{\omega l}^{in,ll}$ ✓

Exterior region $r \in (r_+, \infty) \iff r_* \in (-\infty, \infty)$

$\omega M \ll 1, qQ \ll 1$

– Radial KG:

$$\frac{d^2 h_{\omega l}^{in,l}}{dr_*^2} = \left\{ f(r) \left(\frac{l(l+1)}{r^2} + \frac{\partial_r f(r)}{r} \right) - \left(\omega - \frac{qQ}{r} \right)^2 \right\} h_{\omega l}^{in,l}$$



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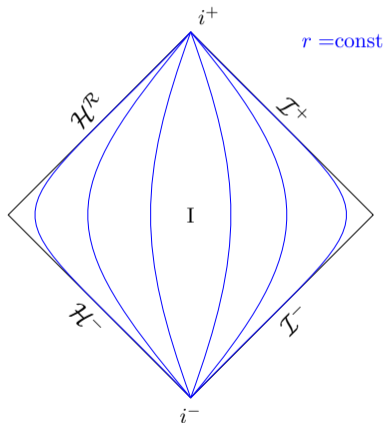
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– Real scalar [Zilberman and Ori 2021]



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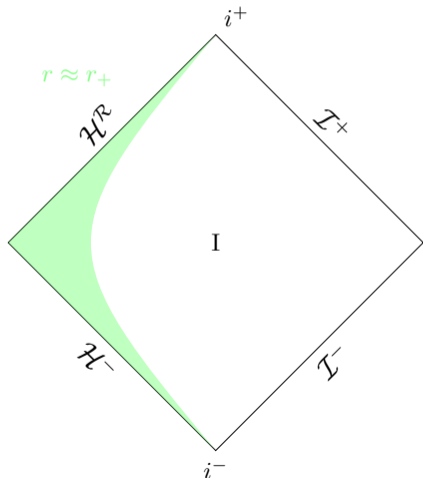
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– Subregions $\left\{ \begin{array}{l} \text{Vicinity of EH} \end{array} \right.$



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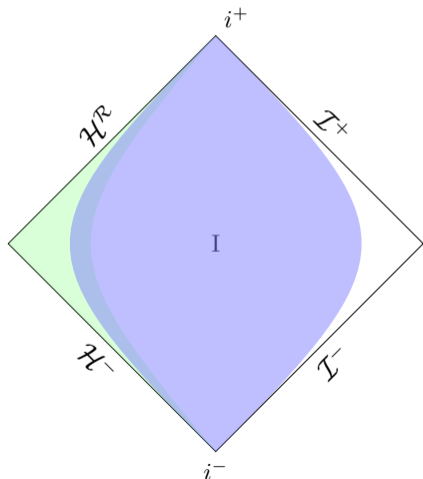
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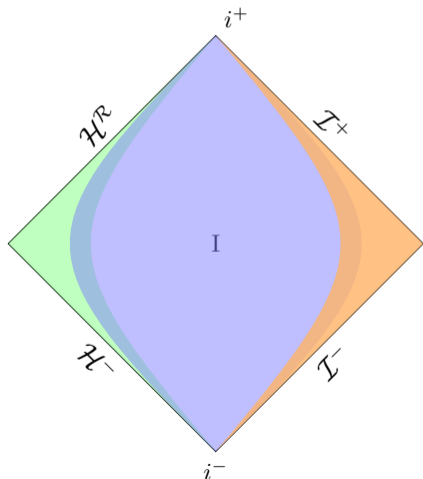
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– Real scalar [Zilberman and Ori 2021]

– Subregions $\begin{cases} \text{Vicinity of EH} \\ \text{Intermediate region} \\ \text{Asymptotically flat region} \end{cases}$



Exterior region $r \in (r_+, \infty) \iff r_* \in (-\infty, \infty)$

$\omega M \ll 1, qQ \ll 1$

– Radial KG:

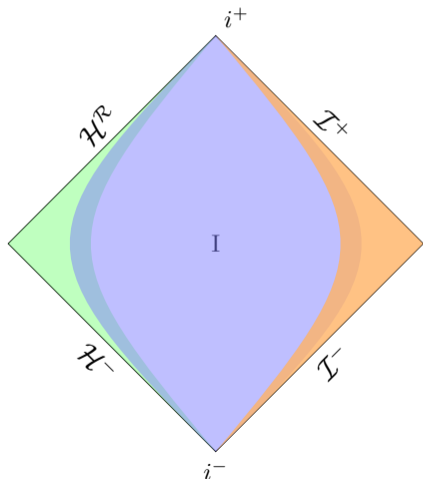
$$\frac{d^2 h_{\omega l}^{in,l}}{dr_*^2} = \left\{ f(r) \left(\frac{l(l+1)}{r^2} + \frac{\partial_r f(r)}{r} \right) - \left(\omega - \frac{qQ}{r} \right)^2 \right\} h_{\omega l}^{in,l}$$

– BCs: $h_{\omega l}^{in,l} = \begin{cases} e^{-i(\omega-\omega_l)r_*} & r_* \rightarrow -\infty \\ A_{\omega l} e^{-i\omega r_*} + B_{\omega l} e^{i\omega r_*} & r_* \rightarrow \infty \end{cases}$

– Real scalar [Zilberman and Ori 2021]

– Subregions $\begin{cases} \text{Vicinity of EH} \\ \text{Intermediate region} \\ \text{Asymptotically flat region} \end{cases}$

– Obtain $A_{\omega l}, B_{\omega l}$ ✓



Exterior region $r \in (r_+, \infty) \iff r_* \in (-\infty, \infty)$ $\omega M \ll 1, qQ \ll 1$

$$\begin{aligned} |T_{\omega, l=0}^{in, l}|^2 &= 4r_+^2 \omega^2, & |R_{\omega, l=0}^{in, l}|^2 &= 1 - 2\omega(\omega - \omega_l)r_+^2 \\ |T_{\omega, l>0}^{in, l}|^2 &\sim \Delta^{4l+2}, & |R_{\omega, l>0}^{in, l}|^2 &= 1 \end{aligned}$$

Results



At the event horizon (EH)

– Explicitly

$$\langle j_V \rangle_U^{EH} = \frac{1}{12\pi^2 r_+^2} \left(\frac{q^4 Q^3}{r_+} + \frac{q^2 Q M^2 \Delta^2}{2r_+^3} \right) + \mathcal{O}(q^4 \Delta, q^2 M^{-2} \Delta^3)$$

$$\langle T_{VV} \rangle_U^{EH} = \frac{-M^4 \Delta^4}{480\pi^2 r_+^8} + \frac{q^4 Q^4}{48\pi^2 r_+^4} + \mathcal{O}(\Delta^5 M^{-4}, q^4 \Delta, q^2 M^{-2} \Delta^3)$$

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- Explicitly

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- Conservation laws & semiclassical E-M equations [Balakumar, Bernar, and Winstanley 2022]

$$\langle j_{r_*} \rangle_\Psi = \frac{\mathcal{K}}{r^2} \tag{3}$$

$$\langle T_{tr_*} \rangle_\Psi = -\frac{\mathcal{L}}{r^2} + \frac{\mathcal{K}Q}{r^3} \tag{4}$$

At the event horizon (EH)

- Explicitly

$$\langle j_v \rangle_U^{EH} = \frac{1}{12\pi^2 r_+^2} \left(\frac{q^4 Q^3}{r_+} + \frac{q^2 Q M^2 \Delta^2}{2r_+^3} \right) + \mathcal{O}(q^4 \Delta, q^2 M^{-2} \Delta^3)$$

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- Conservation laws & semiclassical E-M equations [Balakumar, Bernar, and Winstanley 2022]

$$\langle j_v \rangle_\Psi - \langle j_u \rangle_\Psi = \langle j_{r_*} \rangle_\Psi = \frac{\mathcal{K}}{r^2} \quad (3)$$

$$\langle T_{vv} \rangle_\Psi - \langle T_{uu} \rangle_\Psi = \langle T_{tr_*} \rangle_\Psi = -\frac{\mathcal{L}}{r^2} + \frac{\mathcal{K}Q}{r^3} \quad (4)$$

At the event horizon (EH)

- Explicitly

$$\frac{\mathcal{K}}{r_+^2} = \langle j_v \rangle_U^{EH} = \frac{1}{12\pi^2 r_+^2} \left(\frac{q^4 Q^3}{r_+} + \frac{q^2 Q M^2 \Delta^2}{2r_+^3} \right) + \mathcal{O}(q^4 \Delta, q^2 M^{-2} \Delta^3)$$

$$-\frac{\mathcal{L}}{r_+^2} + \frac{\mathcal{K}Q}{r_+^3} = \langle T_{vv} \rangle_U^{EH} = \frac{-M^4 \Delta^4}{480\pi^2 r_+^8} + \frac{q^4 Q^4}{48\pi^2 r_+^4} + \mathcal{O}(\Delta^5 M^{-4}, q^4 \Delta, q^2 M^{-2} \Delta^3)$$

- Conservation laws & semiclassical E-M equations [Balakumar, Bernar, and Winstanley 2022]

$$\langle j_v \rangle_\Psi - \langle j_u \rangle_\Psi = \langle j_{r_*} \rangle_\Psi = \frac{\mathcal{K}}{r^2} \quad (3)$$

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At the event horizon (EH)

- Explicitly

$$\frac{\mathcal{K}}{r_+^2} = \langle j_v \rangle_U^{EH} = \frac{1}{12\pi^2 r_+^2} \left(\frac{q^4 Q^3}{r_+} + \frac{q^2 Q M^2 \Delta^2}{2r_+^3} \right) + \mathcal{O}(q^4 \Delta, q^2 M^{-2} \Delta^3)$$

$$-\frac{\mathcal{L}}{r_+^2} + \frac{\mathcal{K}Q}{r_+^3} = \langle T_{vv} \rangle_U^{EH} = \frac{-M^4 \Delta^4}{480\pi^2 r_+^8} + \frac{q^4 Q^4}{48\pi^2 r_+^4} + \mathcal{O}(\Delta^5 M^{-4}, q^4 \Delta, q^2 M^{-2} \Delta^3)$$

- Conservation laws & semiclassical E-M equations [Balakumar, Bernar, and Winstanley 2022]

$$\langle j_v \rangle_\Psi - \langle j_u \rangle_\Psi = \langle j_{r_*} \rangle_\Psi = \frac{\mathcal{K}}{r^2} \implies \mathcal{K}Q > 0 \quad (3)$$

$$\langle T_{vv} \rangle_\Psi - \langle T_{uu} \rangle_\Psi = \langle T_{tr_*} \rangle_\Psi = -\frac{\mathcal{L}}{r^2} + \frac{\mathcal{K}Q}{r^3} \implies \mathcal{L} > 0 \quad (4)$$

At the event horizon (EH)

- Explicitly

$$\frac{\mathcal{K}}{r_+^2} = \langle j_v \rangle_U^{EH} = \frac{1}{12\pi^2 r_+^2} \left(\frac{q^4 Q^3}{r_+} + \frac{q^2 Q M^2 \Delta^2}{2r_+^3} \right) + \mathcal{O}(q^4 \Delta, q^2 M^{-2} \Delta^3)$$

$$-\frac{\mathcal{L}}{r_+^2} + \frac{\mathcal{K}Q}{r_+^3} = \langle T_{vv} \rangle_U^{EH} = \frac{-M^4 \Delta^4}{480\pi^2 r_+^8} + \frac{q^4 Q^4}{48\pi^2 r_+^4} + \mathcal{O}(\Delta^5 M^{-4}, q^4 \Delta, q^2 M^{-2} \Delta^3)$$

- Conservation laws & semiclassical E-M equations [Balakumar, Bernar, and Winstanley 2022]

$$\langle j_v \rangle_\Psi - \langle j_u \rangle_\Psi = \langle j_{r_*} \rangle_\Psi = \frac{\mathcal{K}}{r^2} \implies \mathcal{K}Q > 0 \quad (3)$$

$$\langle T_{vv} \rangle_\Psi - \langle T_{uu} \rangle_\Psi = \langle T_{tr_*} \rangle_\Psi = -\frac{\mathcal{L}}{r^2} + \frac{\mathcal{K}Q}{r^3} \implies \mathcal{L} > 0 \quad (4)$$

- Weak backreaction: $\partial_v Q = -4\pi r^2 \langle j_v \rangle \implies$ BH discharge

At the inner horizon (IH)

– Explicitly

$$\langle j_v \rangle_U^{IH} = \frac{1}{12\pi^2 r_-^2} \left(\frac{q^4 Q^3}{r_+} + \frac{q^2 Q M^2 \Delta^2}{2r_+^3} \right) + \mathcal{O}(q^4 \Delta, q^2 M^{-2} \Delta^3)$$

$$\langle T_{vv} \rangle_U^{IH} = \frac{-M^4 \Delta^4}{480\pi^2 r_-^2 r_+^6} + \frac{q^4 Q^4}{48\pi^2 r_-^2 r_+^2} + \mathcal{O}(\Delta^5 M^{-4}, q^4 \Delta, q^2 M^{-2} \Delta^3)$$

At the inner horizon (IH)

- Explicitly

$$\langle j_v \rangle_U^{IH} = \frac{1}{12\pi^2 r_-^2} \left(\frac{q^4 Q^3}{r_+} + \frac{q^2 Q M^2 \Delta^2}{2r_+^3} \right) + \mathcal{O}(q^4 \Delta, q^2 M^{-2} \Delta^3)$$

$$\langle T_{vv} \rangle_U^{IH} = \frac{-M^4 \Delta^4}{480\pi^2 r_-^2 r_+^6} + \frac{q^4 Q^4}{48\pi^2 r_-^2 r_+^2} + \mathcal{O}(\Delta^5 M^{-4}, q^4 \Delta, q^2 M^{-2} \Delta^3)$$

- At the inner horizon,

$$\langle j_u \rangle_U^{IH} \ll \langle j_v \rangle_U^{IH}, \quad \langle T_{uu} \rangle_U^{IH} \ll \langle T_{vv} \rangle_U^{IH}$$

At the inner horizon (IH)

- Explicitly

$$\langle j_v \rangle_U^{IH} = \frac{1}{12\pi^2 r_-^2} \left(\frac{q^4 Q^3}{r_+} + \frac{q^2 Q M^2 \Delta^2}{2r_+^3} \right) + \mathcal{O}(q^4 \Delta, q^2 M^{-2} \Delta^3)$$

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- At the inner horizon,

$$\langle j_u \rangle_U^{IH} \ll \langle j_v \rangle_U^{IH}, \quad \langle T_{uu} \rangle_U^{IH} \ll \langle T_{vv} \rangle_U^{IH}$$

- Leading order contribution from $l = 0$ modes

At the inner horizon (IH)

- Explicitly

$$\langle j_v \rangle_U^{IH} = \frac{1}{12\pi^2 r_-^2} \left(\frac{q^4 Q^3}{r_+} + \frac{q^2 Q M^2 \Delta^2}{2r_+^3} \right) + \mathcal{O}(q^4 \Delta, q^2 M^{-2} \Delta^3)$$

$$\langle T_{vv} \rangle_U^{IH} = \frac{-M^4 \Delta^4}{480\pi^2 r_-^2 r_+^6} + \frac{q^4 Q^4}{48\pi^2 r_-^2 r_+^2} + \mathcal{O}(\Delta^5 M^{-4}, q^4 \Delta, q^2 M^{-2} \Delta^3)$$

- At the inner horizon,

$$\langle j_u \rangle_U^{IH} \ll \langle j_v \rangle_U^{IH}, \quad \langle T_{uu} \rangle_U^{IH} \ll \langle T_{vv} \rangle_U^{IH}$$

- Leading order contribution from $l = 0$ modes
- Weak backreaction $\partial_v Q = -4\pi r^2 \langle j_v \rangle \implies$ discharge of BH interior









Conclusions and outlook



Conclusions

- Studied charged scalar field in Unruh state on RN spacetime
- Near-extremal domain \implies analytical treatment
- Quantum effects: charge superradiance, Hawking radiation
- Behaviour of $\langle j_V \rangle$, $\langle T_{VV} \rangle$ at the inner horizons \rightarrow connection to sCC
- (Brief) backreaction analysis
- Internship with the RTG: numerical cross-check

Thank you!

-  Balakumar, Visakan, Rafael P. Bernar, and Elizabeth Winstanley (2022). “Quantization of a charged scalar field on a charged black hole background”. In: *Phys. Rev. D*.
-  Cardoso, Vitor et al. (2018). “Quasinormal modes and strong cosmic censorship”. In: *Physical review letters* 120.3, p. 031103.
-  Christodolou, D (1991). “The formation of black holes and singularities in spherically symmetric gravitational collapse Commun”. In: *Pure and Appl. Mathematics* 44, pp. 339–373.
-  Dias, Oscar JC, Harvey S Reall, and Jorge E Santos (2019). “Strong cosmic censorship for charged de Sitter black holes with a charged scalar field”. In: *Classical and Quantum Gravity*.
-  Hollands, Stefan, Robert M Wald, and Jochen Zahn (2020). “Quantum instability of the Cauchy horizon in Reissner–Nordström–deSitter spacetime”. In: *Classical and Quantum Gravity* 37.11, p. 115009.
-  Klein, Christiane and Jochen Zahn (July 2021). “Renormalized charged scalar current in the Reissner–Nordström–de Sitter spacetime”. In: *Phys. Rev. D* 104 (2), p. 025009. DOI: 10.1103/PhysRevD.104.025009.
-  Klein, Christiane, Jochen Zahn, and Stefan Hollands (Dec. 2021). “Quantum (Dis)Charge of Black Hole Interiors-supplementary material”. In: *Phys. Rev. Lett.* 127 (23), p. 231301.
-  Penrose, Roger (Jan. 1965). “Gravitational Collapse and Space-Time Singularities”. In: *Phys. Rev. Lett.* 14 (3), pp. 57–59. DOI: 10.1103/PhysRevLett.14.57.
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