Relative entropy in de Sitter

Markus B. Fröb

Institut für Theoretische Physik, Universität Leipzig

Quantum Effects in Gravitational Fields Conference 30. August 2023, Universität Leipzig

Based on arXiv:2308.14797 and work in progress with Albert Much (U Leipzig) and Kyriakos Papadopoulos (U Kuwait)

1 Entanglement entropy in QM and QFT

2 Tomita–Takesaki theory

3 De Sitter spacetime and relative entropy



Entanglement entropy in QM and QFT

Entanglement entropy in QM and QFT

Entanglement entropy in QM

- Entropy quantifies the amount of information an observer has access to
- For a system in a pure state $|\Psi\rangle$, everything is known \Leftrightarrow entropy is zero
- For a generic state described by density matrix ρ , define von Neumann entropy $S_{\rm vN}(\rho) = -\operatorname{tr}(\rho \ln \rho) \ge 0$ (since $0 \le \rho \le 1$)
- For a system in a pure state $|\Psi\rangle$, but with the observer having access only to degrees of freedom within some region A, define density matrix $\rho_A = \text{tr}_{A^{\perp}} |\Psi\rangle\langle\Psi|$ (trace over degrees of freedom of the complement region A^{\perp})
- Entanglement entropy: $S(A) = S_{vN}(\rho_A) = -\operatorname{tr}(\rho_A \ln \rho_A)$
- Other entropy measures: Tsallis entropy $S_q^{\mathsf{T}}(\rho) = \frac{1}{q-1}(1 \mathrm{tr}\rho^q)$, Rényi entropy $S_{\alpha}^{\mathsf{R}}(\rho) = \frac{1}{1-\alpha} \ln \mathrm{tr}\rho^{\alpha}$, and $\lim_{q \to 1} S_q^{\mathsf{T}}(\rho) = S_{\mathsf{vN}}(\rho) = \lim_{\alpha \to 1} S_{\alpha}^{\mathsf{R}}(\rho)$
- Thermal density matrix $\rho = \frac{1}{Z} \exp(-\beta H)$ with inverse temperature β , $Z = \exp(-\beta F)$ with free energy F gives $S_{vN} = \beta(\langle H \rangle F) = S$ (thermodynamic entropy)

Entanglement entropy in QFT

- In QFT in d + 1 dimensions, density matrices only exist formally and the trace is infinite, both due to the infinite number of degrees of freedom
- Compute regularised entanglement entropies with UV cutoff ϵ : $S(A) = g_{d-1}[\partial A]\epsilon^{-(d-1)} + \cdots + g_1[\partial A]\epsilon^{-1} + g_0[\partial A]\ln\epsilon + S_0(A) + \mathcal{O}(\epsilon)$, where the g_i are homogeneous functions depending on the boundary ∂A (area law in 4D)
- Source of divergences: high-energy vacuum fluctuations of the fields
- $\blacksquare \Rightarrow$ differences in entropies between different states are finite
- Relative entropy (Kullback–Leibler divergence): $S(\rho \| \sigma) = tr(\rho \ln \rho \rho \ln \sigma)$
- Relative Rényi entropy: $S^{\mathsf{R}}_{\alpha}(\rho \| \sigma) = \frac{1}{\alpha 1} \ln \operatorname{tr}(\rho^{\alpha} \sigma^{1 \alpha})$
- α -z-Rényi entropy (generalized quantum Rényi div.): $S_{\alpha,z}^{\text{RG}}(\rho \| \sigma) = \frac{z}{\alpha-1} \ln \operatorname{tr}\left(\rho^{\frac{\alpha}{z}} \sigma^{\frac{1-\alpha}{z}}\right)$ and $\lim_{\alpha \to 1} S_{\alpha}^{\text{R}}(\rho \| \sigma) = S(\rho \| \sigma) = \lim_{\alpha, z \to 1} S_{\alpha,z}^{\text{RG}}(\rho \| \sigma)$

Tomita–Takesaki theory

Tomita–Takesaki theory

- Mathematically, difference between QM and QFT is the type of (factors of) von Neumann algebra of operators (I, II, III)
- Tomita–Takesaki theory gives information on structure of vN algebra A ⊂ B(H) acting on Hilbert space H, given a cyclic and separating vector Ω ∈ H
- Tomita operator S is the closure of the map $S_0\colon a\Omega o a^\dagger\Omega$ for $a\in\mathfrak{A}$
- Polar decomposition $S = J\Delta^{\frac{1}{2}}$ gives positive modular operator $\Delta = S^{\dagger}S \ge 0$ and antilinear modular conjugation J
- Modular flow $\sigma_s(a) = \Delta^{\mathsf{i}s} a \, \Delta^{-\mathsf{i}s} \in \mathfrak{A}$ for $a \in \mathfrak{A}$
- State ω defined by Ω is a thermal (KMS) state: $\omega(\sigma_s(a)b) = (\Omega, \sigma_s(a)b\Omega)$ satisfies $\omega(\sigma_{s-i}(a)b) = \omega(b \sigma_s(a))$, with inverse temperature normalised to $\beta = 1$
- Both J and $\Delta^{\frac{1}{2}}$ map \mathfrak{A} to commutant \mathfrak{A}'

Tomita–Takesaki theory

- Relative Tomita operator S_{Φ|Ψ} is closure of map aΦ → a[†]Ψ for a ∈ 𝔄 and cyclic and separating vectors Φ, Ψ ∈ ℋ, relative modular operator Δ_{Φ|Ψ} and relative modular conjugation J_{Φ|Ψ} defined by polar decomposition S_{Φ|Ψ} = J_{Φ|Ψ} Δ^{1/2}_{Φ|Ψ}
- Araki formula relates relative modular Hamiltonian $\ln \Delta_{\Phi|\Psi}$ to relative entropy: $S(\Phi||\Psi) = -(\Phi, \ln \Delta_{\Phi|\Psi}\Phi)$ (well-defined and finite)
- Important case: $\Phi = uu'\Omega$ and $\Psi = vv'\Omega$ for unitary operators $u, v \in \mathfrak{A}$ and $u', v' \in \mathfrak{A}'$ commuting with u and v
- $\bullet \Rightarrow \Delta_{\Phi|\Psi} = u' v \Delta_{\Omega} v^{\dagger}(u')^{\dagger} \text{ and } S(\Phi||\Psi) = -\left(v^{\dagger} u \Omega, \ln \Delta_{\Omega} v^{\dagger} u \Omega\right)$
- Relative entropy between two "excited" states relative to a "vacuum" state Ω can be computed using only the modular Hamiltonian ln Δ_Ω of the "vacuum" state, e.g., for coherent state with u = u' = v' = 1 and v = exp[iφ(f)]

De Sitter spacetime and relative entropy

De Sitter spacetime

- *d*-dimensional de Sitter is embedded in (d + 1)-dimensional Minkowski space as hyperboloid $\eta_{AB}X^AX^B = H^{-2}$
- Expanding half of dS (Poincaré patch) with metric $ds^2 = \eta_{AB} dX^A dX^B = -dt^2 + e^{2Ht} dx^2$ describes primordial inflation and current accelerated expansion of our universe
- Maximally symmetric solution of Einstein's equations with cosmological constant $\Lambda = (d-1)H^2$
- Generator of boosts is tangent to hyperboloid: $M_{0j} = X_0 \partial_{X^j} X_j \partial_{X^0} = -\frac{1}{2H} \Big(H^2 \mathbf{x}^2 e^{-2Ht} + 1 \Big) \partial_j \mathbf{x}_j \partial_t + H \mathbf{x}_j \mathbf{x}^i \partial_i$
- Modular Hamiltonian is known for dS vacuum state Ω and algebra \mathfrak{A} generated by fields restricted to intersection of hyperboloid and wedge $W_1 = \{X^A : X^1 \ge |X^0|\}$: $\ln \Delta_{\Omega} = iM_{01}$ (Borchers/Buchholz, Global properties of vacuum states in de Sitter space 1999)



Relative entropy in de Sitter spacetime

- Compute relative entropy S(Ω||Ψ_f) for coherent state Ψ_f = e^{iφ(f)}Ω via Araki formula (= relative entanglement entropy, with entanglement region A the intersection of hyperboloid and wedge A = {(t, x): 2Hx¹ ≥ |1 e^{-2Ht} + H²x²|}, and supp f ⊂ A)
- $S(\Omega \| \Psi_f) = i\pi \langle \Delta(M_{01}f), \Delta f \rangle$ with commutator function $\Delta(x, y) = i[\phi(x), \phi(y)]$ and symplectic product $\langle f, g \rangle = i \int [f^*(t, \mathbf{x})\dot{g}(t, \mathbf{x}) - g(t, \mathbf{x})\dot{f}^*(t, \mathbf{x})]_{t=0} d^{d-1}\mathbf{x}$
- Further manipulations: $S(\Omega || \Psi_f) = 2\pi \int \left[\mathbf{x}^1 \mathcal{H}(\hat{f}, x) + \frac{1}{2} H \mathbf{x}^2 \partial_1 \hat{f} \partial_t \hat{f} - H \mathbf{x}^1 \mathbf{x}^i \partial_i \hat{f} \partial_t \hat{f} \right]_{t=0} d^{d-1} \mathbf{x} \text{ with}$ $\hat{f}(x) = \int \Delta(x, y) f(y) \sqrt{-g} d^d y \text{ and } \mathcal{H}(g, x) = \frac{1}{2} \left(\dot{g}^2 + e^{-2Ht} \partial_i g \partial^i g + m^2 g^2 \right) \text{ the}$ Hamiltonian density \Rightarrow not manifestly positive!
- However, correct flat-space limit $H \rightarrow 0$

De Sitter spacetime and relative entropy

Relative entropy in de Sitter spacetime

- Solution: evaluate symplectic product on different Cauchy surface $\Sigma = \{(t, \mathbf{x}): 2Ht + \ln(1 + H^2\mathbf{x}^2) = 0\} \text{ instead of just } t = 0$ • $S(\Omega \| \Psi_f) = 2\pi \int \mathcal{Q}(\hat{f}) (1 + H^2\mathbf{x}^2)^{-\frac{d}{2}} d^{d-1}\mathbf{x}$ with $\mathcal{Q}(h) = \frac{\mathbf{x}^1}{2\sqrt{1+H^2\mathbf{x}^2}} [\partial_n h \partial_n h + [(1 + H^2\mathbf{x}^2)\delta^{kl} - H^2\mathbf{x}^k\mathbf{x}^l] \hat{\partial}_k h \hat{\partial}_l h + m^2 h^2]_{\Sigma} \ge 0$ • $\mathcal{Q}(h) = -n_\mu \xi_{\nu}^{(1)} T^{\mu\nu}(h)$ with n_μ normal to Σ , $\xi_{\nu}^{(1)}$ Killing vector associated to boosts: $M_{01} = \xi_{(1)}^{\mu} \partial_{\mu}$, and $T_{\mu\nu}$ canonical stress tensor
- Q is a Noether charge, compare Wald (Black hole entropy is the Noether charge 1993) and Floerchinger (Lectures on quantum fields and information theory)
- Relative entropy is also convex: $\lambda S(\Omega \| \Psi_f) + (1 \lambda)S(\Omega \| \Psi_g) \ge S(\Omega \| \Psi_{\lambda g + (1 \lambda)h})$ for supp $f, g \subset A$ and $\lambda \in [0, 1]$

${\sf Outlook}$

Outlook

- Modular Hamiltonian is also known in some other cases:
- De Sitter vacuum state Ω and algebra \mathfrak{A} generated by fields restricted to static patch: i ln $\Delta_{\Omega} = \frac{2\pi}{H} \partial_{T}$ (static patch appears thermal with inverse temperature $\beta = 2\pi/H$)
- Recent result for conformal theories in small diamonds of size ℓ in the static patch: i ln $\Delta_{\Omega} = \frac{2\pi}{H} \partial_{T} - \frac{\pi}{H^{2}\ell} \frac{e^{-HT}}{\sqrt{1-H^{2}R^{2}}} \Big[\partial_{T} - H(1-H^{2}R^{2}) \Big(R\partial_{R} + \frac{d-2}{2} \Big) \Big]$
- \blacksquare \Rightarrow compute relative entropy for small diamonds
- Other measures: mutual information $I(A, B) = S(A) + S(B) S(A \cup B)$ for $A \cap B = \emptyset$

- Outlook

Thank you for your attention

Questions?

This work has been funded by the Deutsche Forschungsgemeinschaft (DFG, German Research Foundation) — project no. 396692871 within the Emmy Noether grant CA1850/1-1 and project no. 406116891 within the Research Training Group RTG 2522/1.

References: arXiv:2308.14797, arXiv:2309.XXXXX