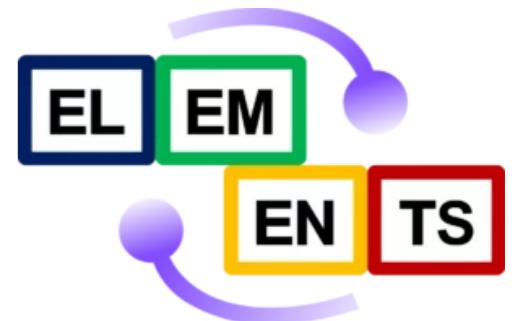


Towards a microscopic understanding of macroscopic properties of strong-interaction matter

Jens Braun

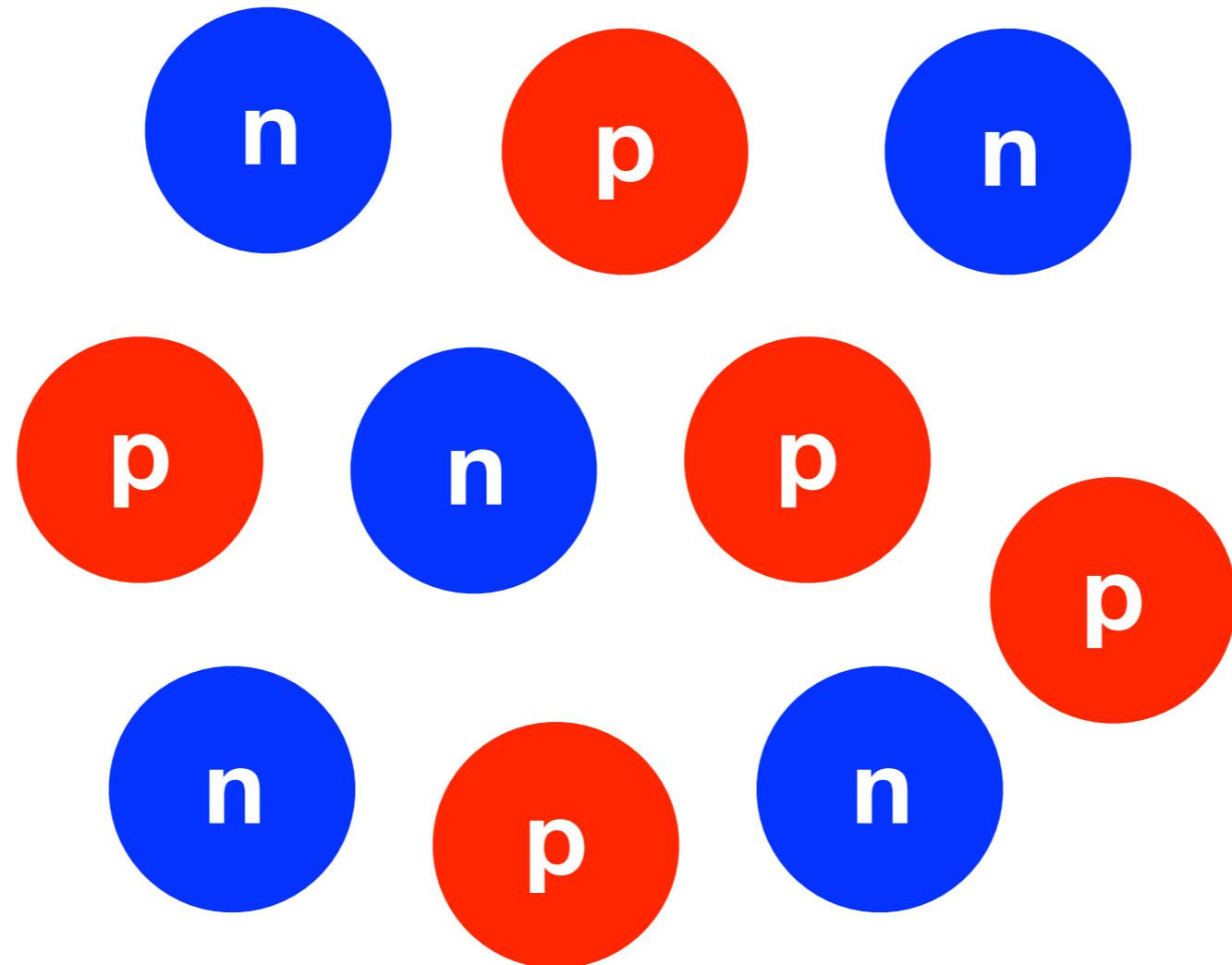
TU Darmstadt

26/07/2022

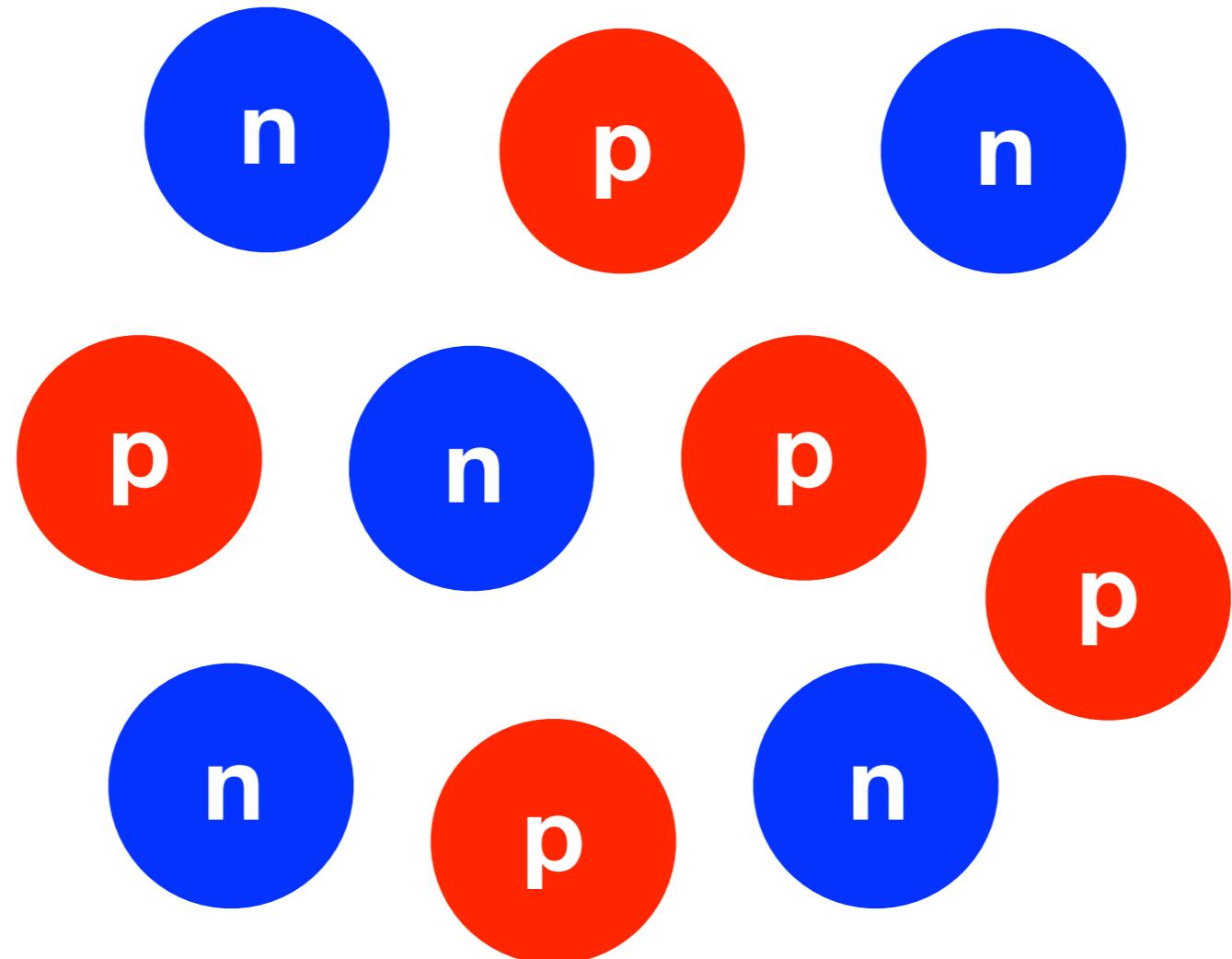


Strong-interaction matter: what is it?

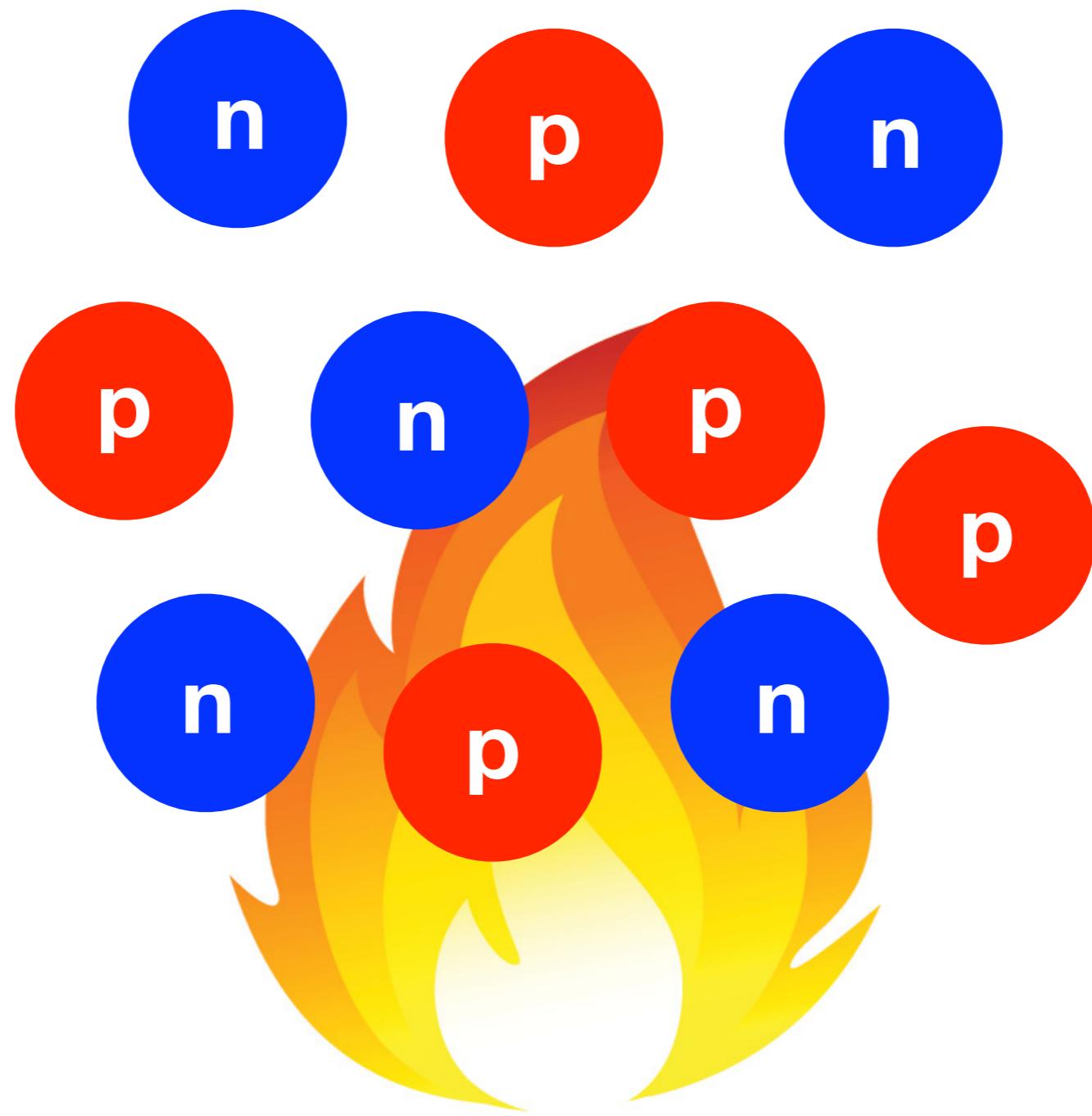
Strong-interaction matter: “one form of appearance”



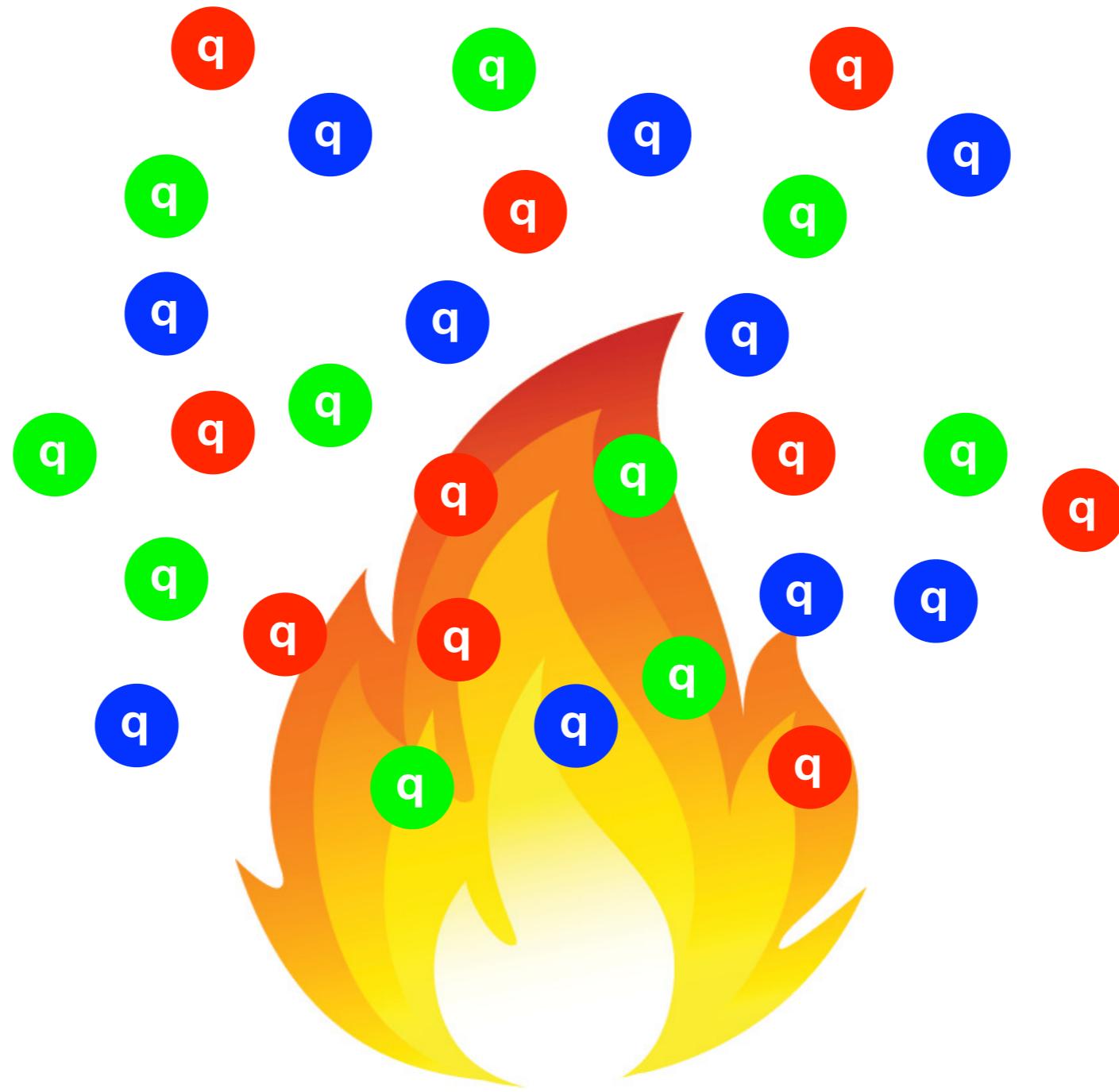
Strong-interaction matter: what can you do with it?



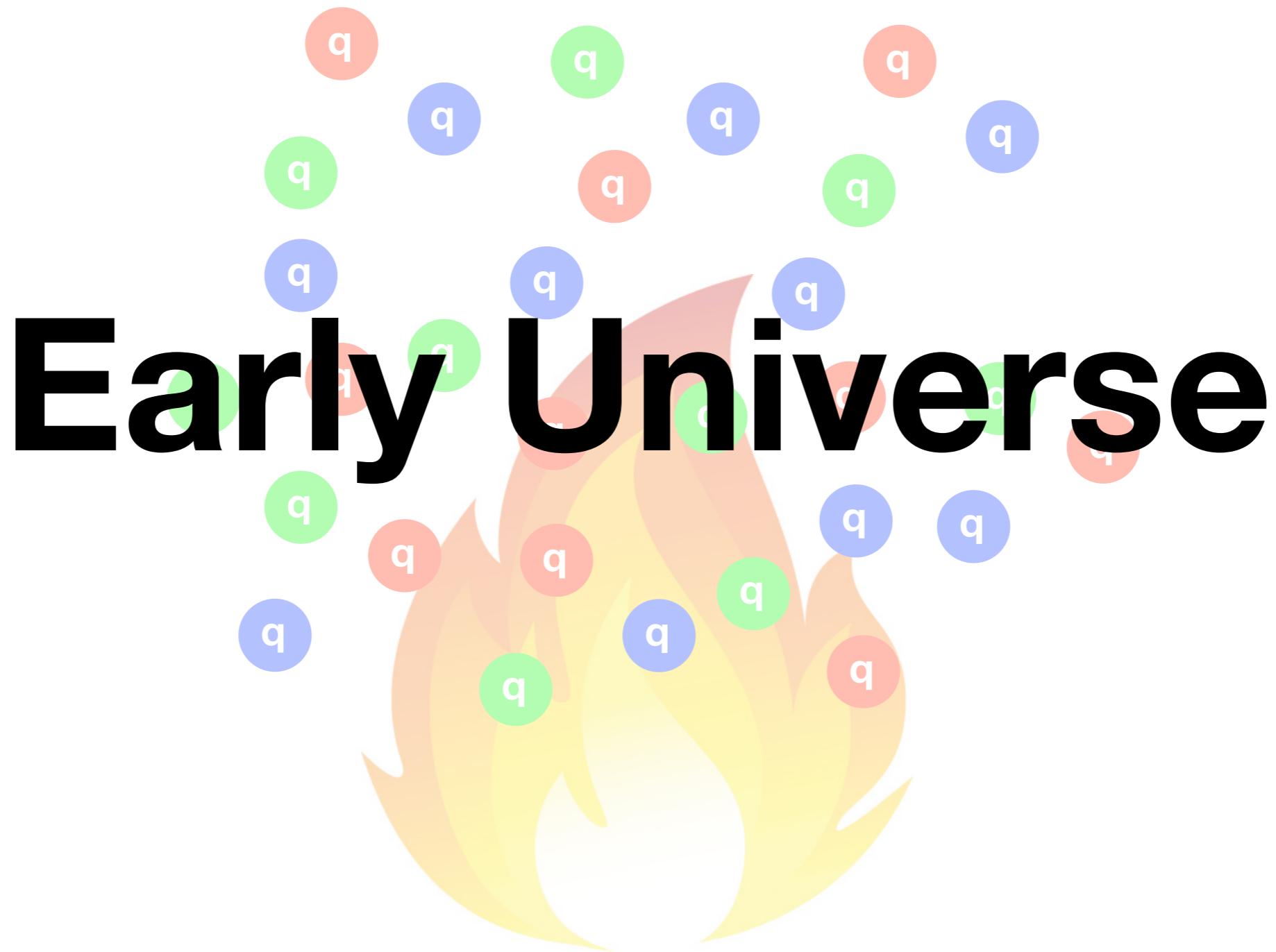
Strong-interaction matter: you can heat it up



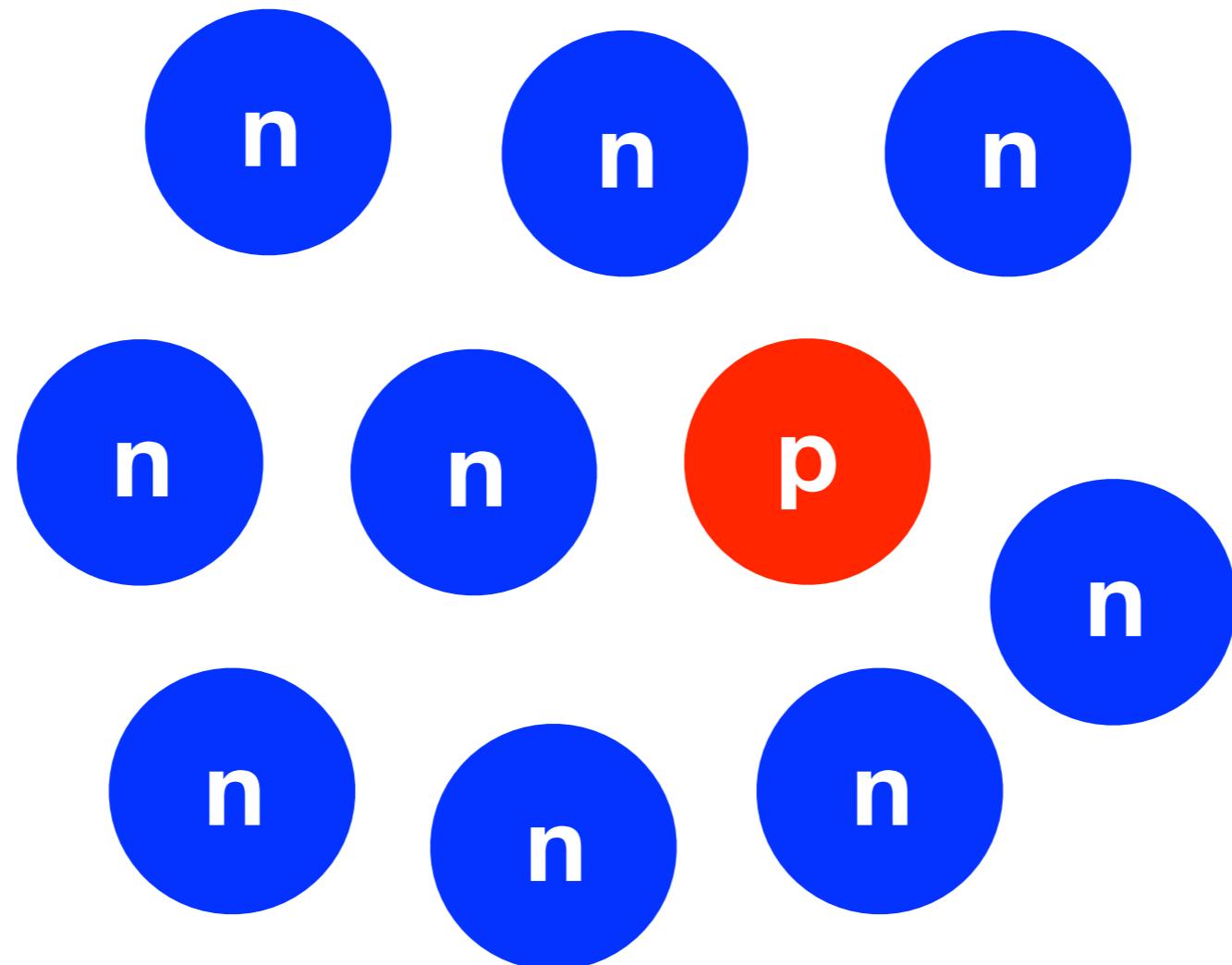
Strong-interaction matter: you can heat it up



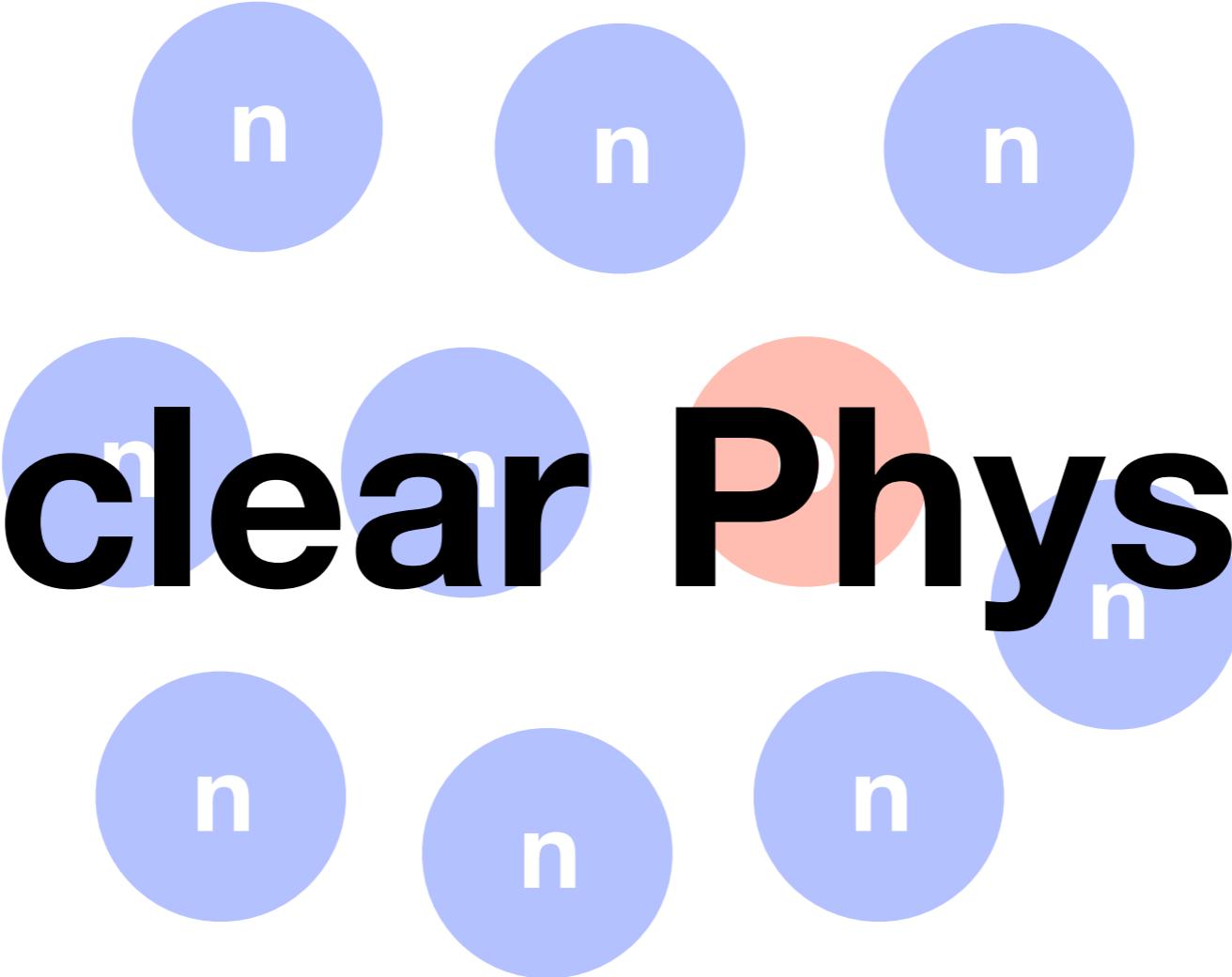
Strong-interaction matter: you can heat it up



Strong-interaction matter: you can “polarize” it

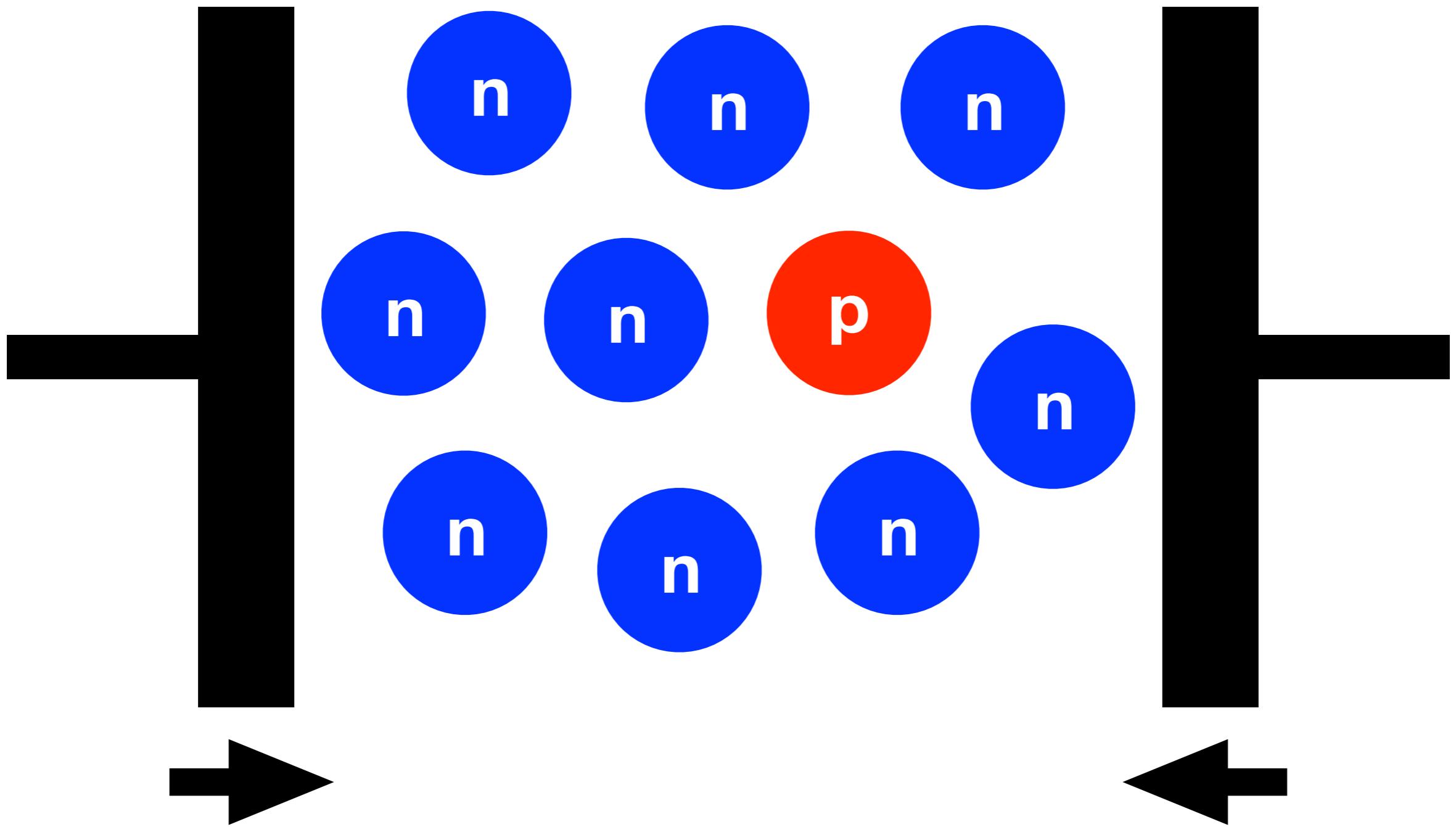


Strong-interaction matter: why “polarizing” it?

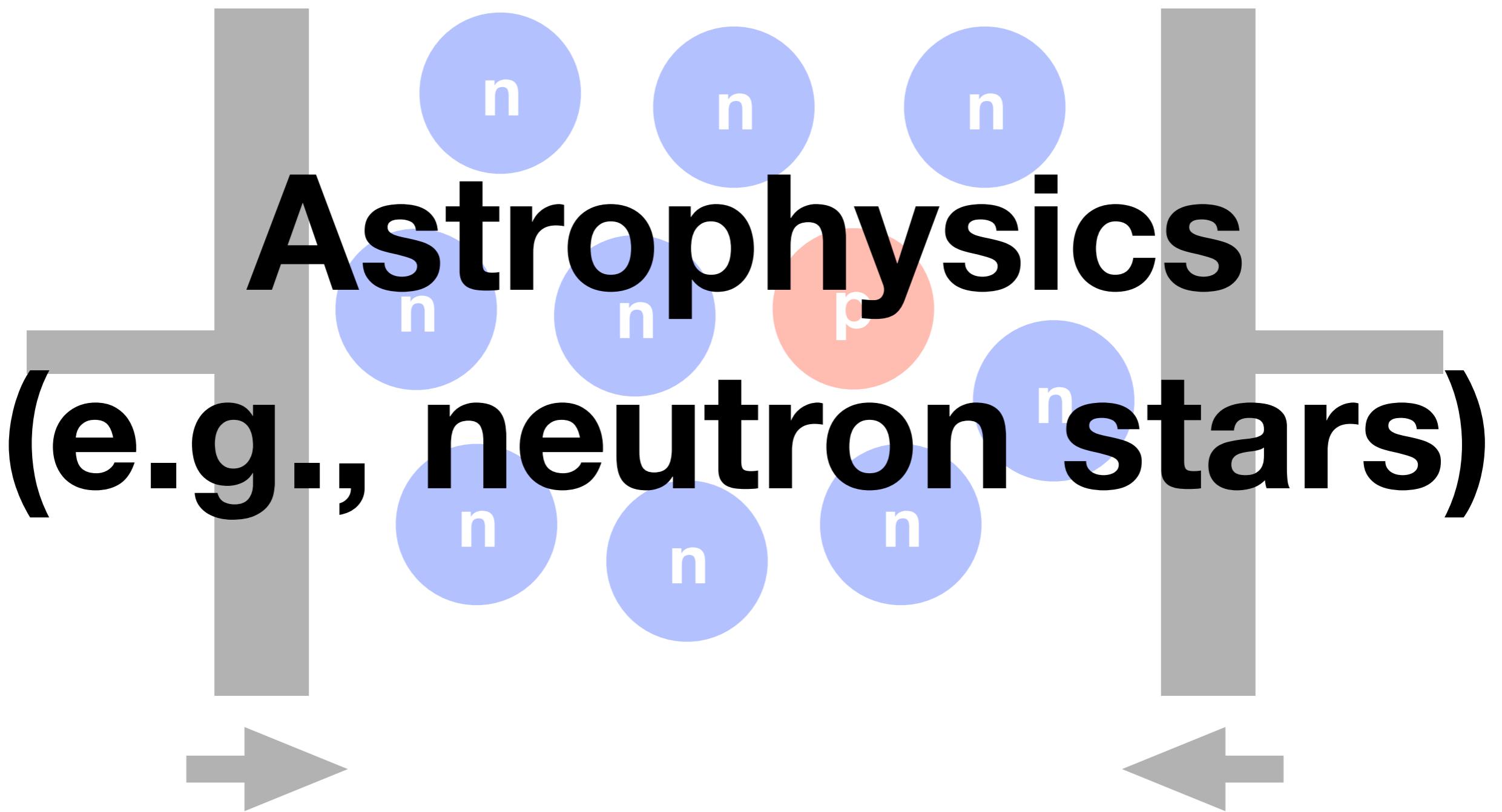


Nuclear Physics

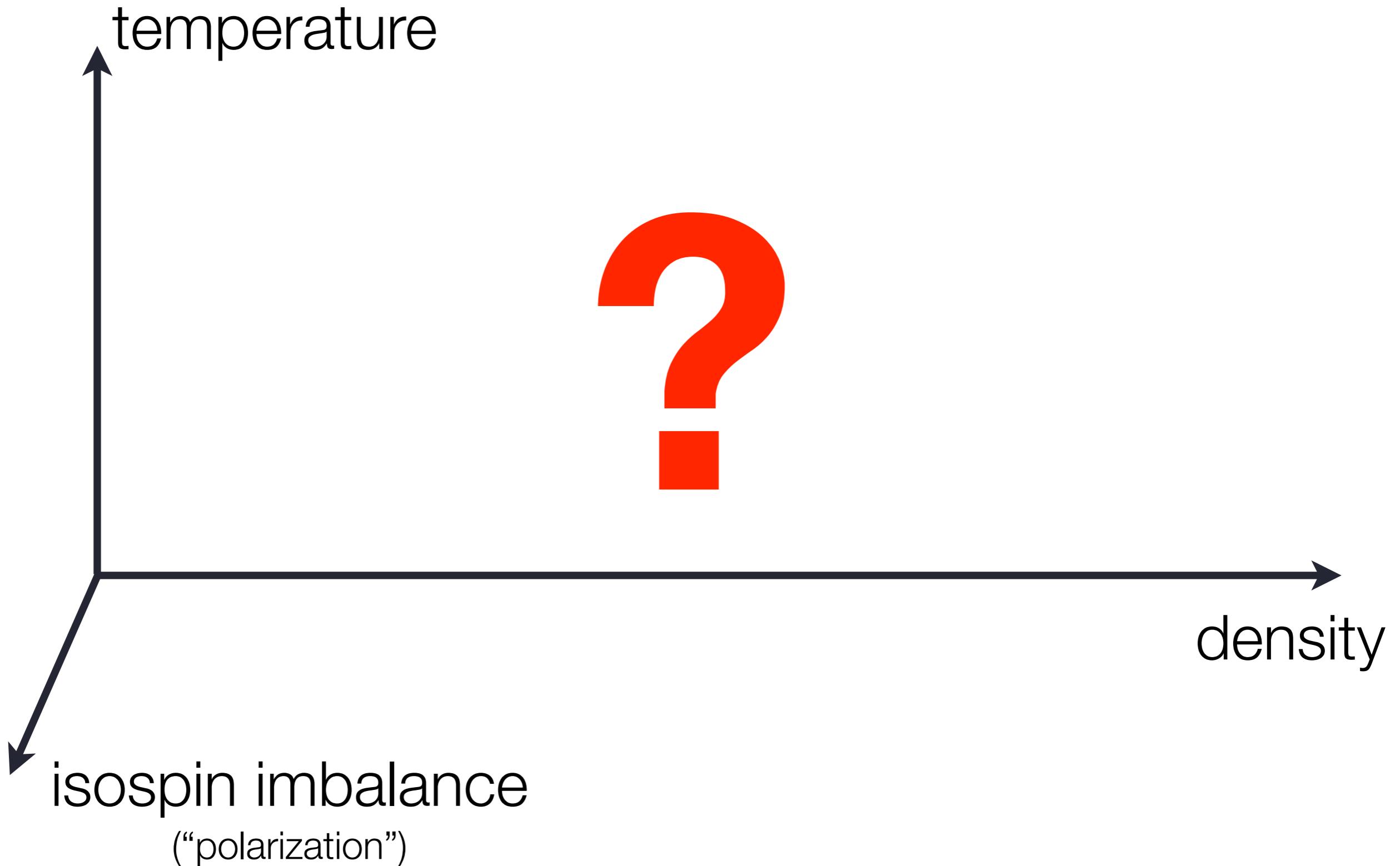
Strong-interaction matter: you can then compress it



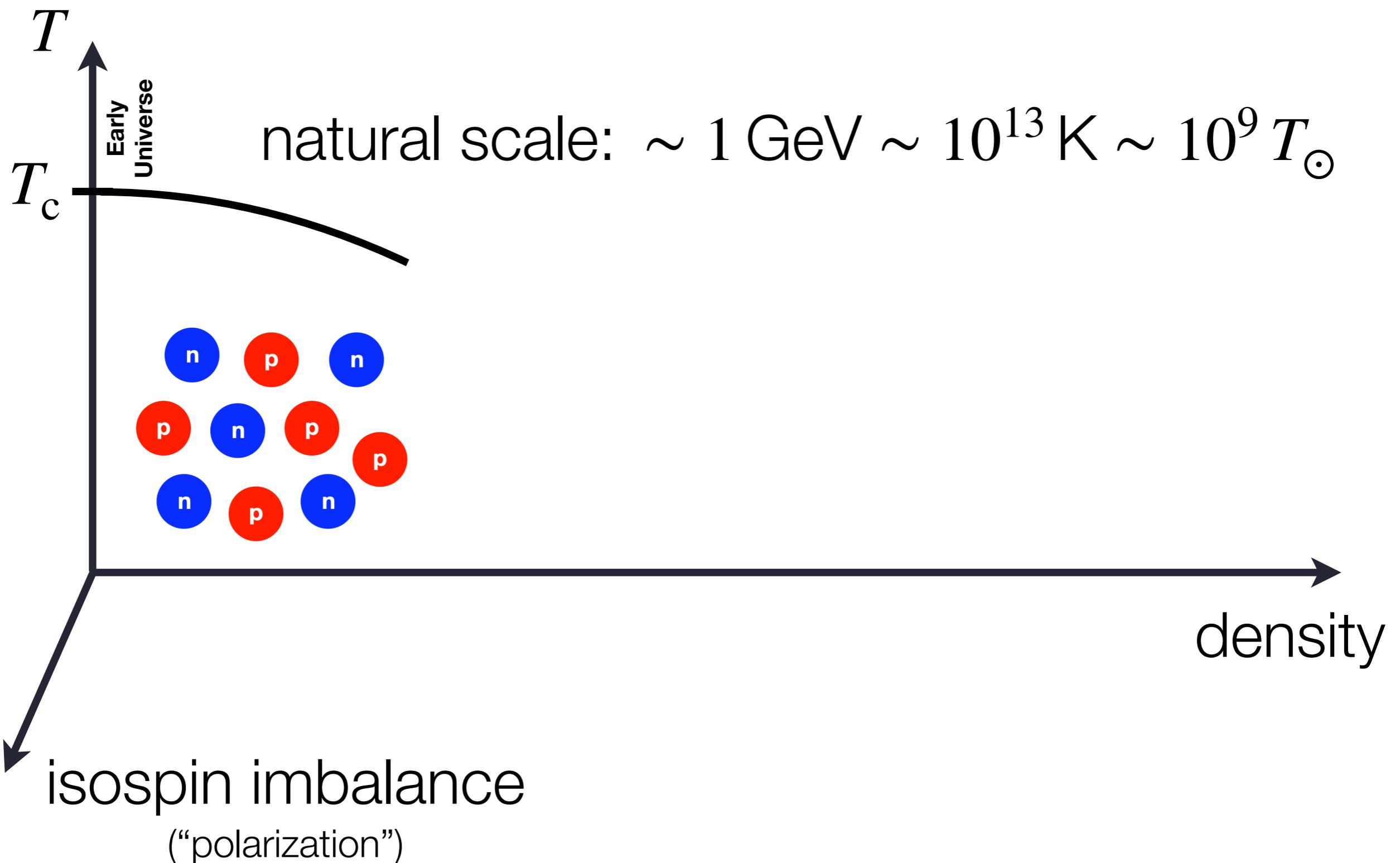
Strong-interaction matter: why compressing it?



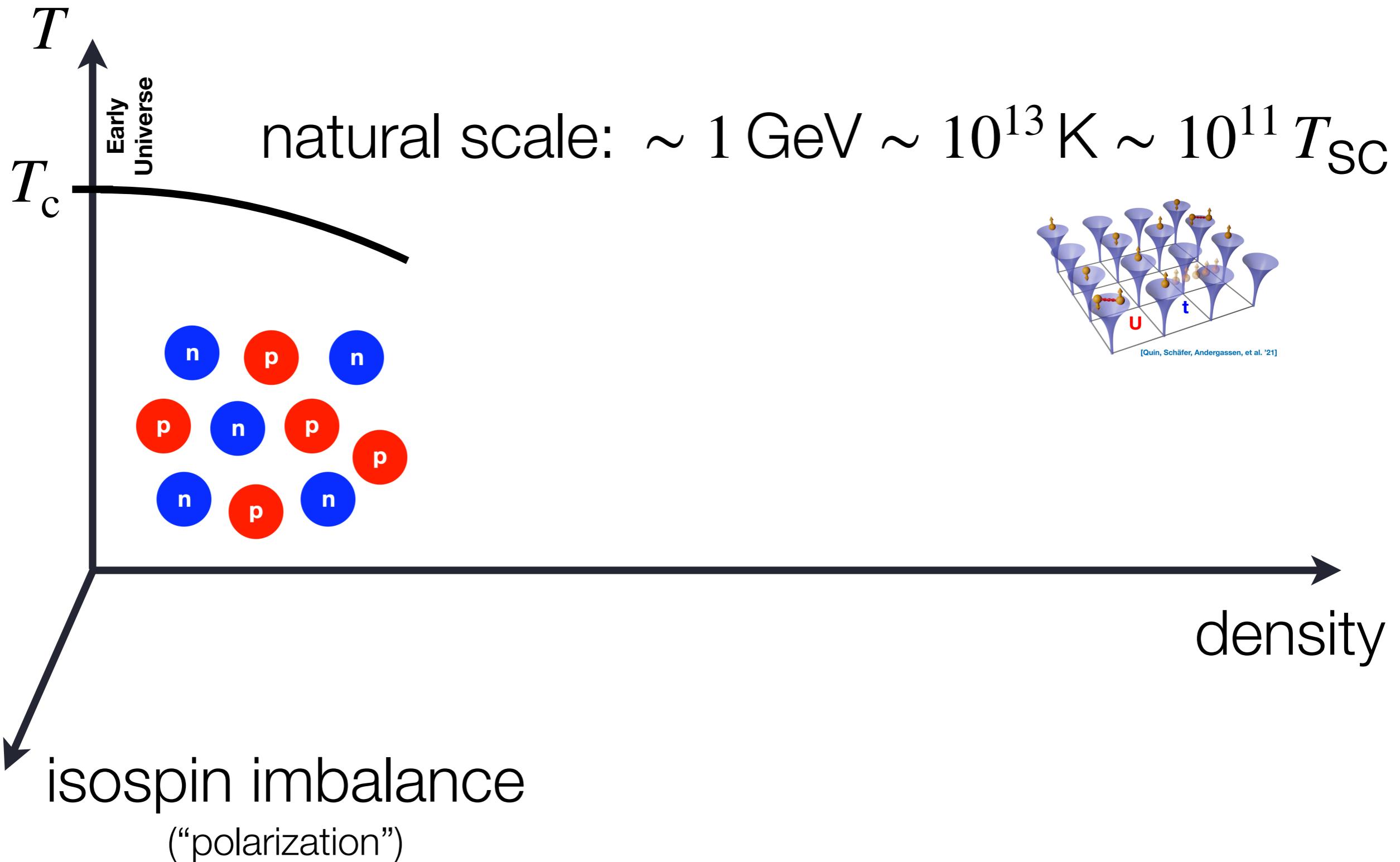
Phases of strong-interaction matter?



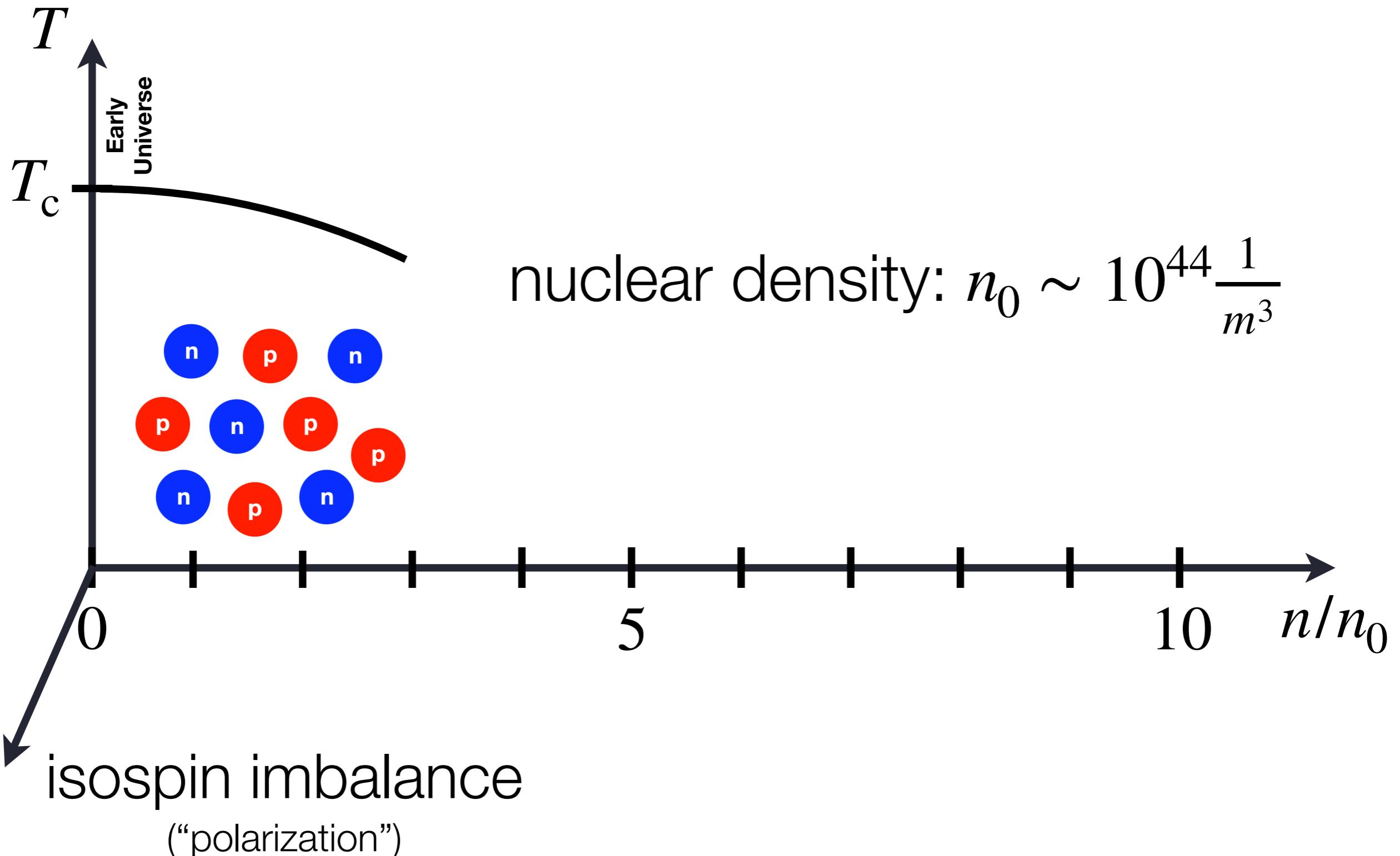
Phases of strong-interaction matter?



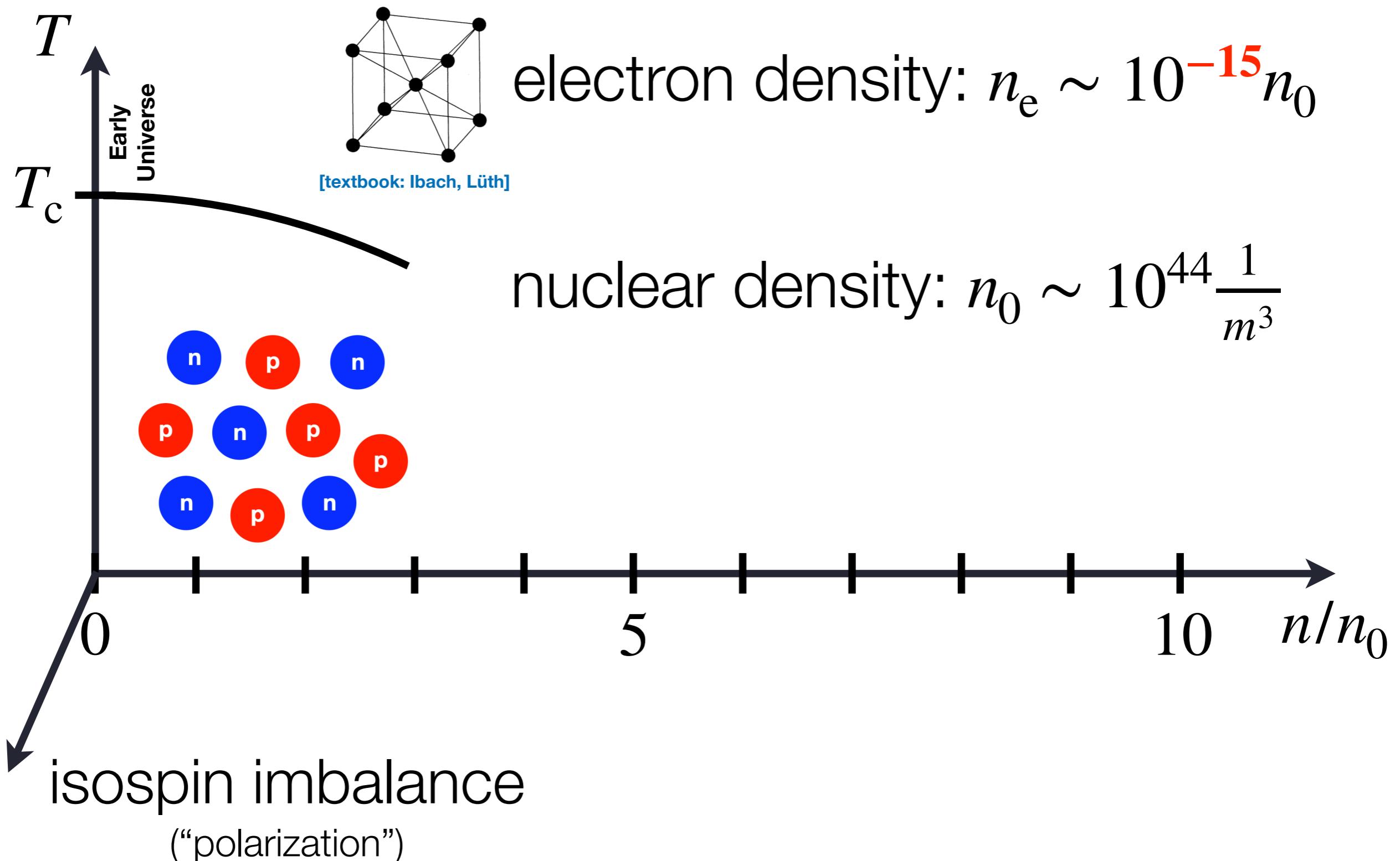
Phases of strong-interaction matter?



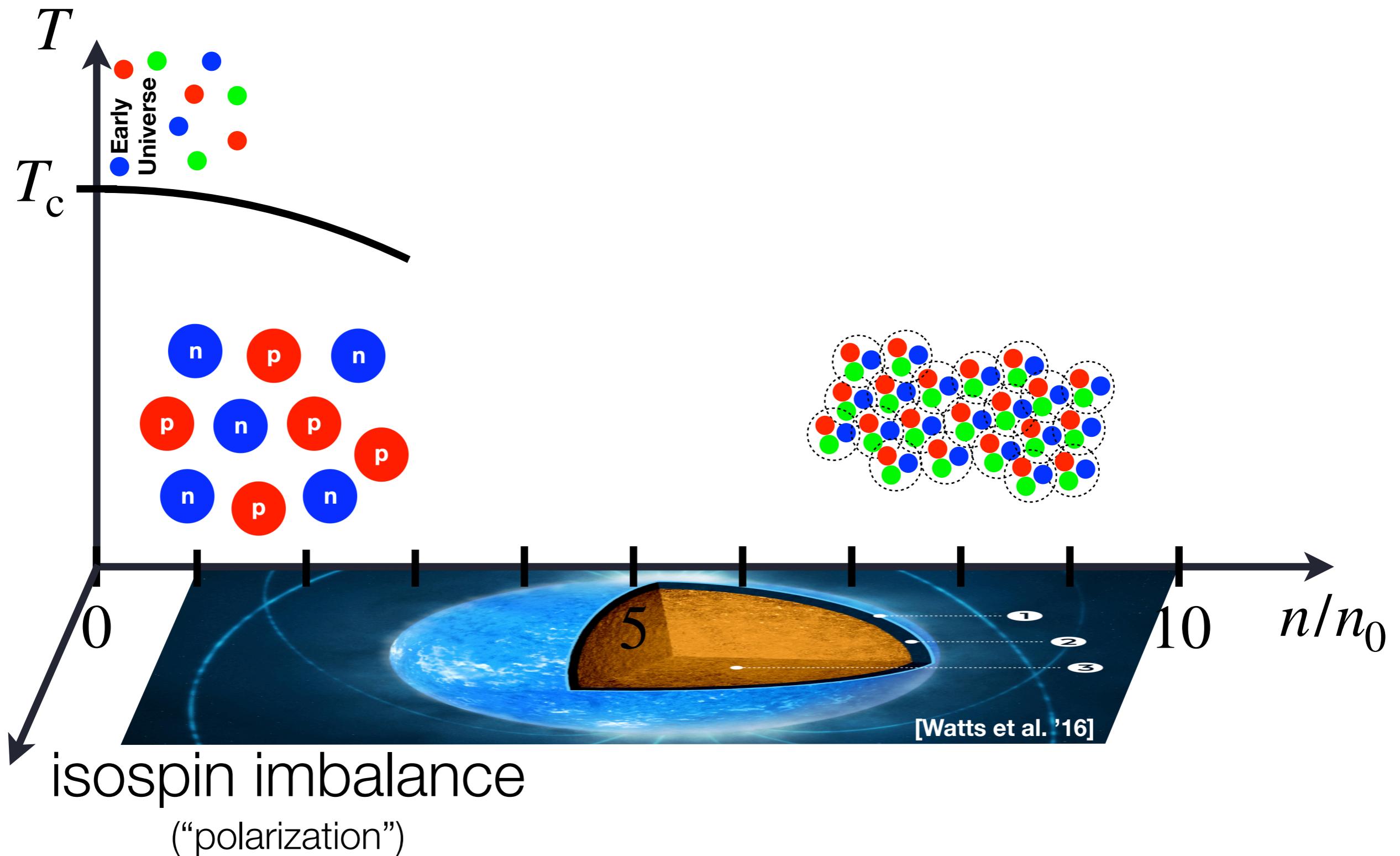
Phases of strong-interaction matter?



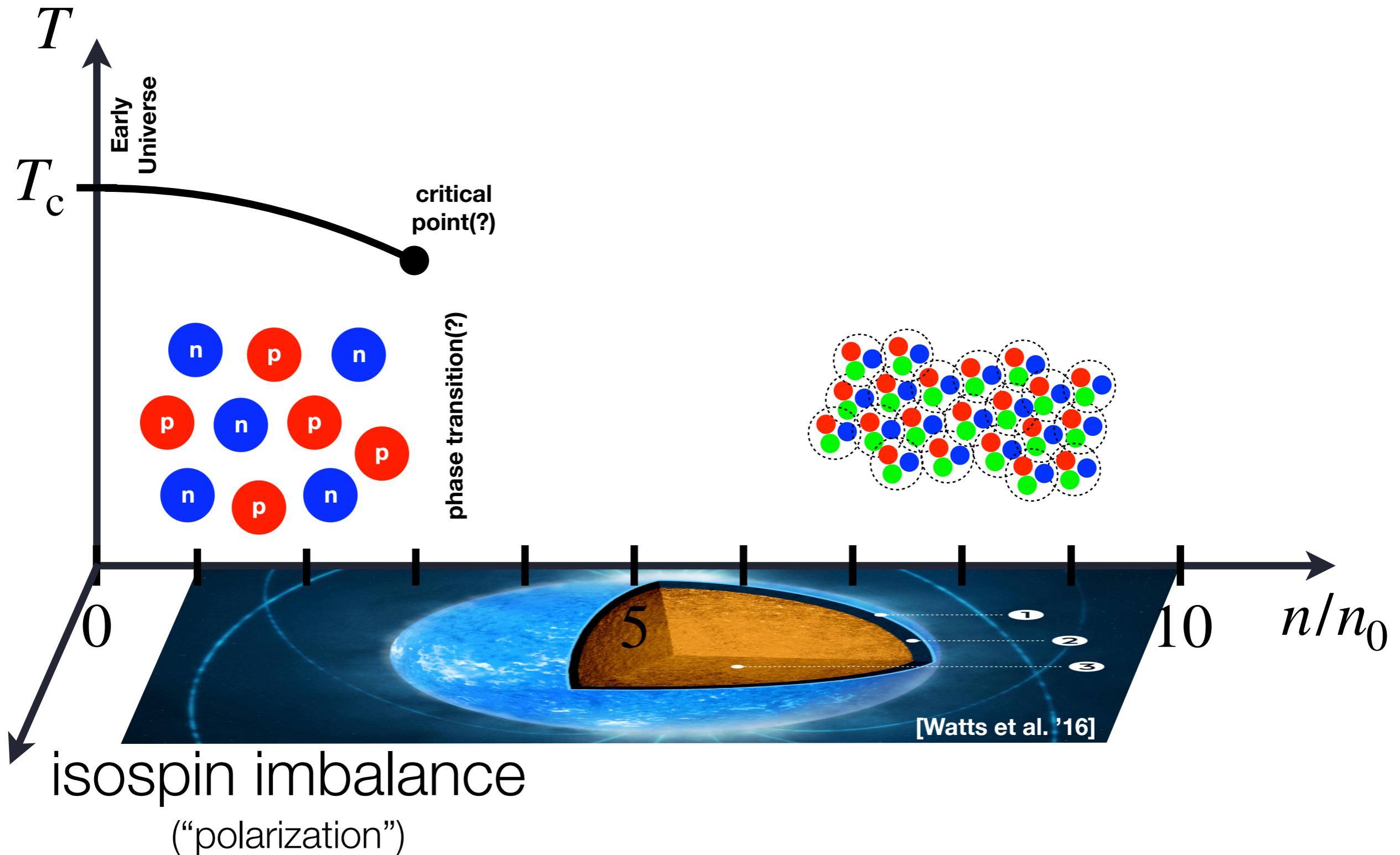
Phases of strong-interaction matter?



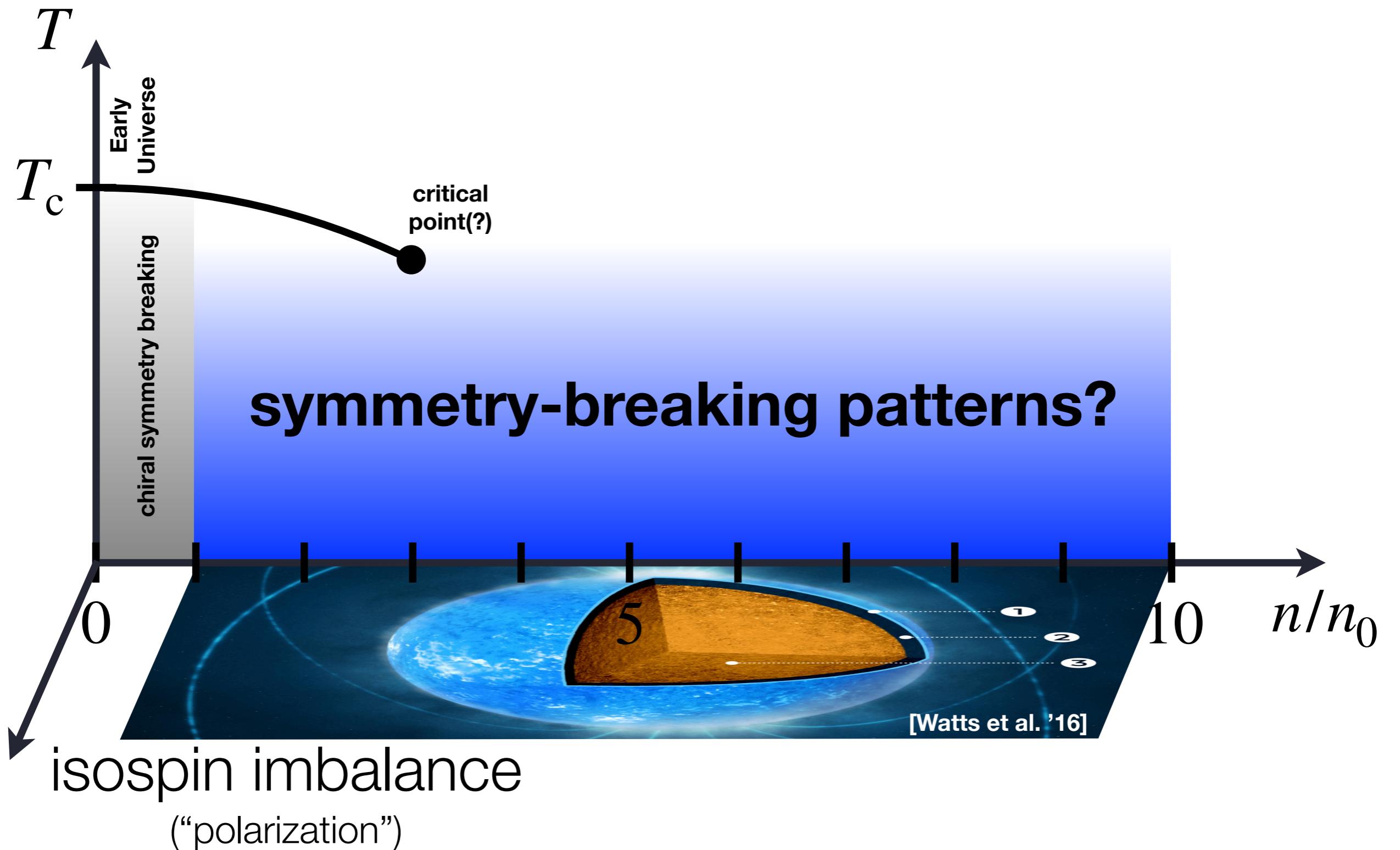
Phases of strong-interaction matter?



Phases of strong-interaction matter?



Phases of strong-interaction matter?



Theory of the strong interaction (QCD)

$$S = \int_0^{\beta} d\tau \int d^3x \left\{ \underbrace{\frac{1}{4} F_{\mu\nu}^a F_{\mu\nu}^a}_{\text{gluons}} + \bar{\psi} \left(i\partial + g_s A - i\mu \gamma_0 \right) \psi \right\}$$

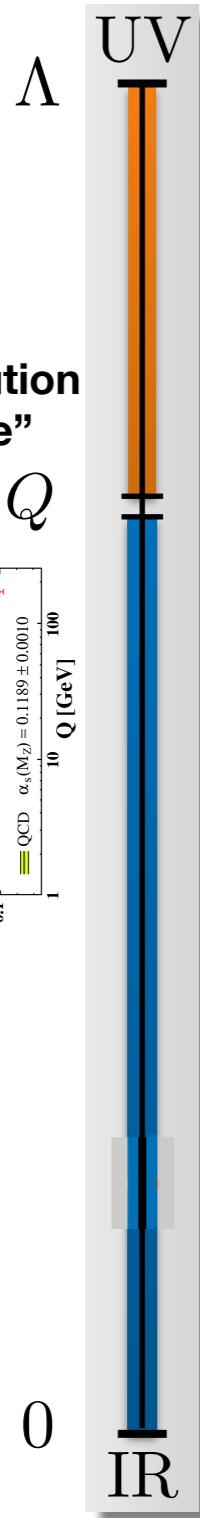
inverse temperature

quarks

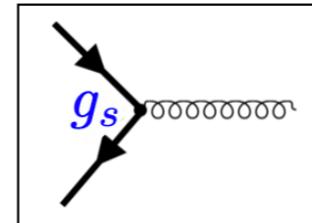
quark-gluon coupling chemical potential

Functional Renormalization Group

[Wetterich '92]

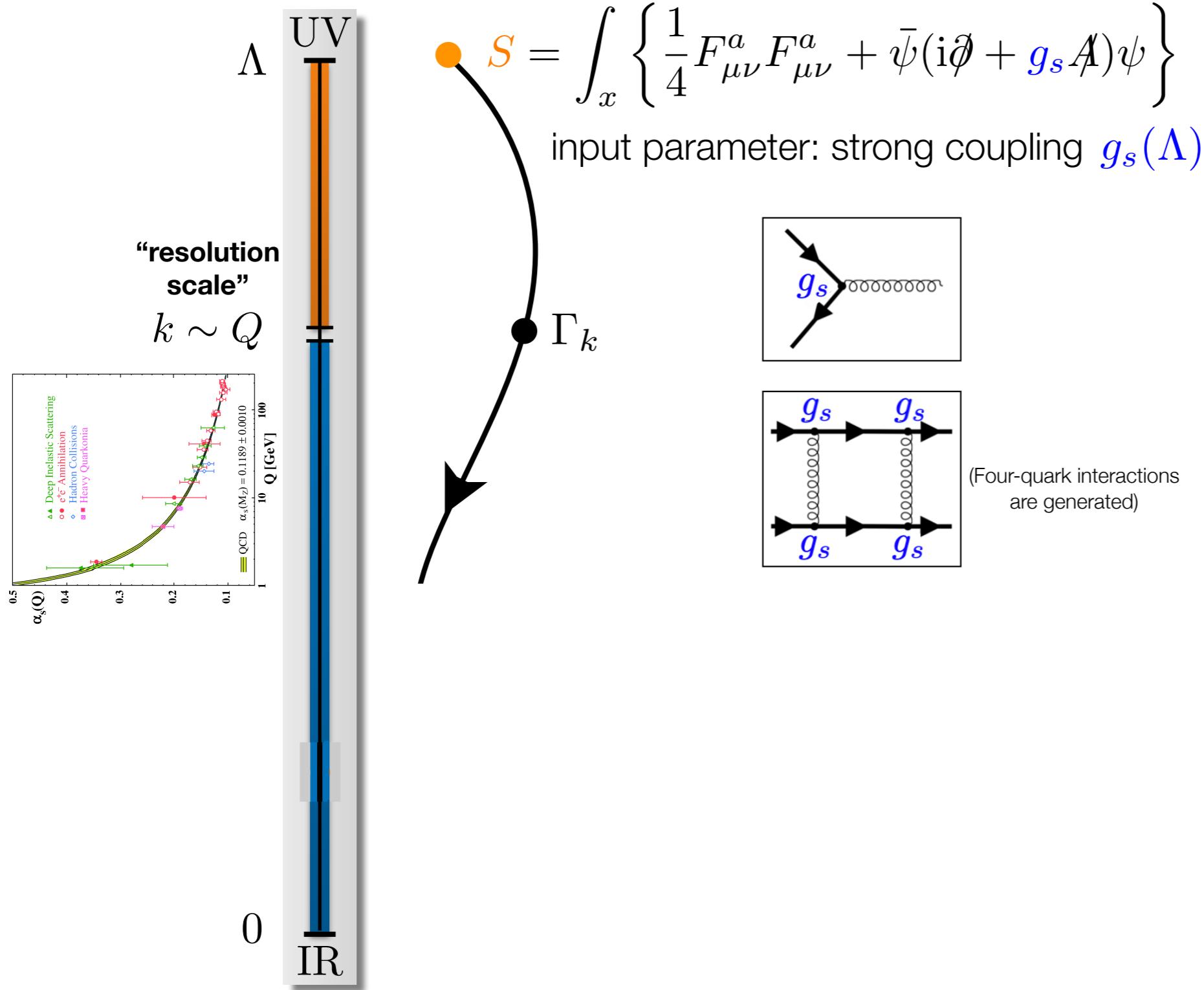


- $S = \int_x \left\{ \frac{1}{4} F_{\mu\nu}^a F_{\mu\nu}^a + \bar{\psi} (i\partial + g_s A) \psi \right\}$
input parameter: strong coupling $g_s(\Lambda)$



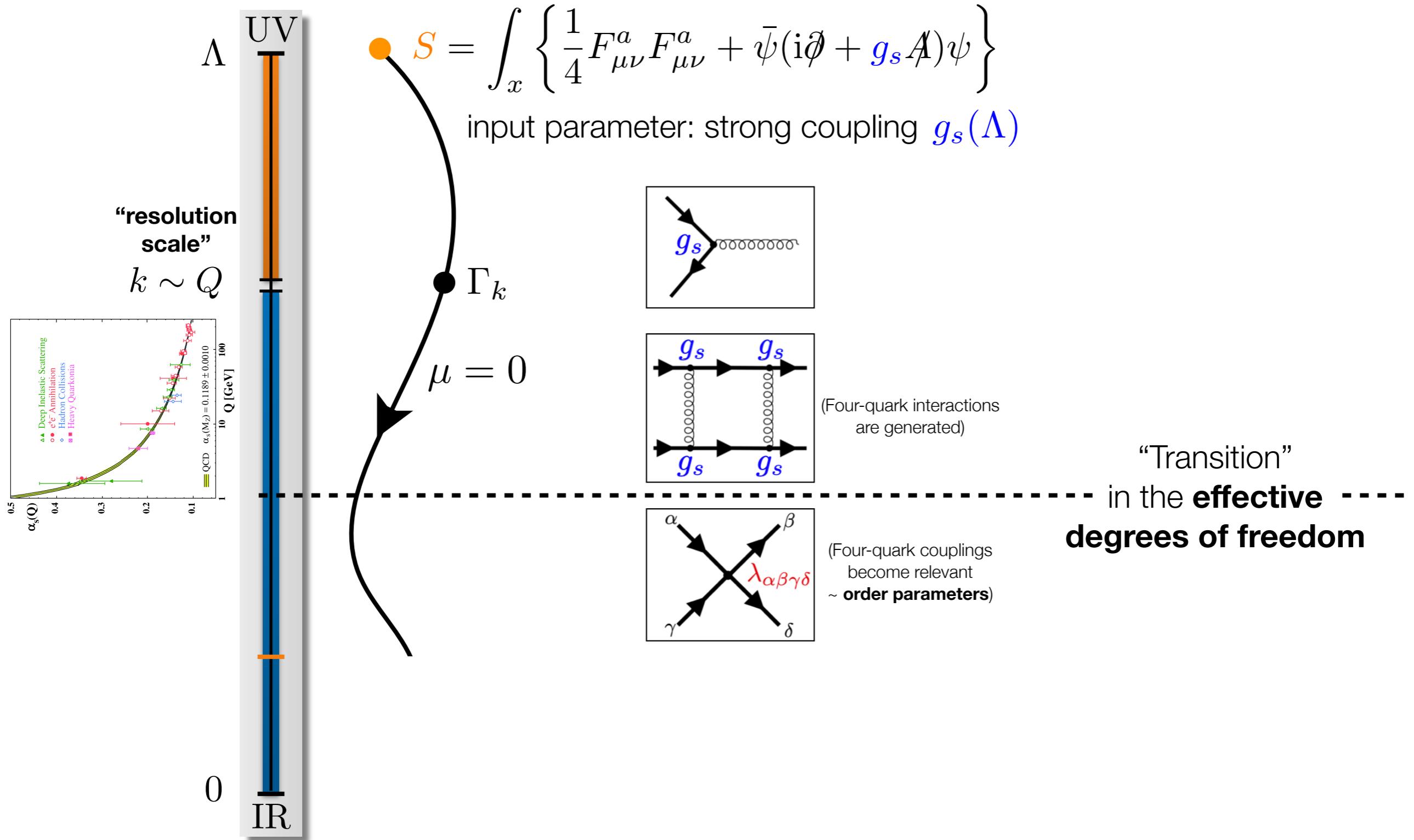
Functional Renormalization Group

[Wetterich '92]



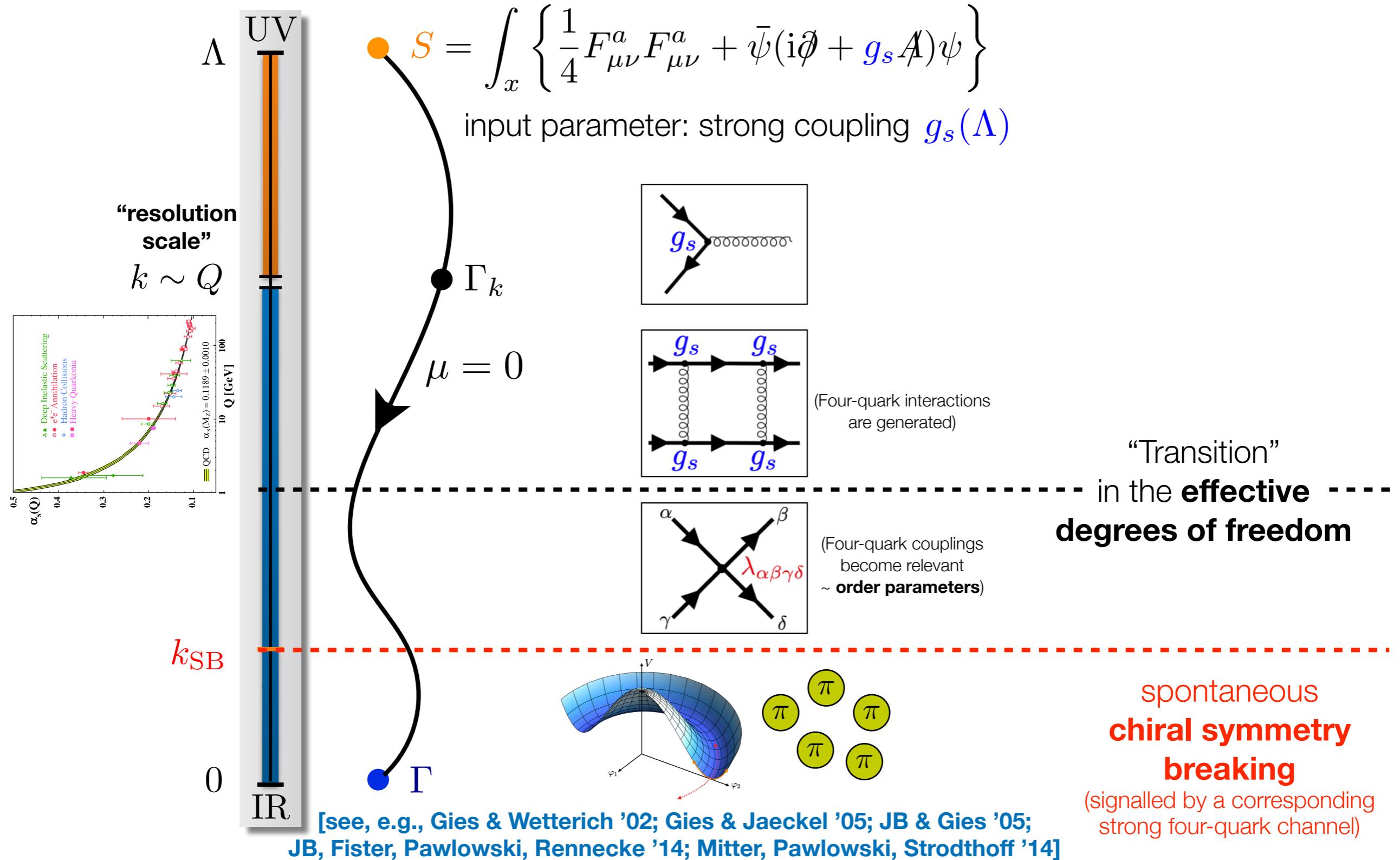
Functional Renormalization Group: Vacuum

[Wetterich '92]



Functional Renormalization Group: Vacuum

[Wetterich '92]



How many four-quark channels? Symmetries!

How many four-quark channels? Symmetries!

[Gies & Jaeckel '05; JB & Gies '05, '06; Mitter, Pawłowski, Strodthoff '14; JB, Leonhardt, Pospiech' 17]

symmetry	group
color	$SU(N_c)$
chiral	$SU_L(2) \otimes SU_R(2)$
vector	$U_V(1)$
axial	$U_A(1)$
Poincare	✓
parity	✓
charge conjugation	✓
# of channels (pointlike limit)	4
	Fierz-complete set

How many four-quark channels? Symmetries!

[JB, Leonhardt, Pospiech' 17]

symmetry	group
color	$SU(N_c)$
chiral	$SU_L(2) \otimes SU_R(2)$
vector	$U_V(1)$
axial	$U_A(1)$
Poincare	✗
parity	✓
charge conjugation	✗
# of channels (pointlike limit)	10

Fierz-complete
set

How many four-quark channels? Symmetries!

[JB, Leonhardt, Pospiech' 17; JB, Leonhardt, Pawłowski, Rosenblüh '20]

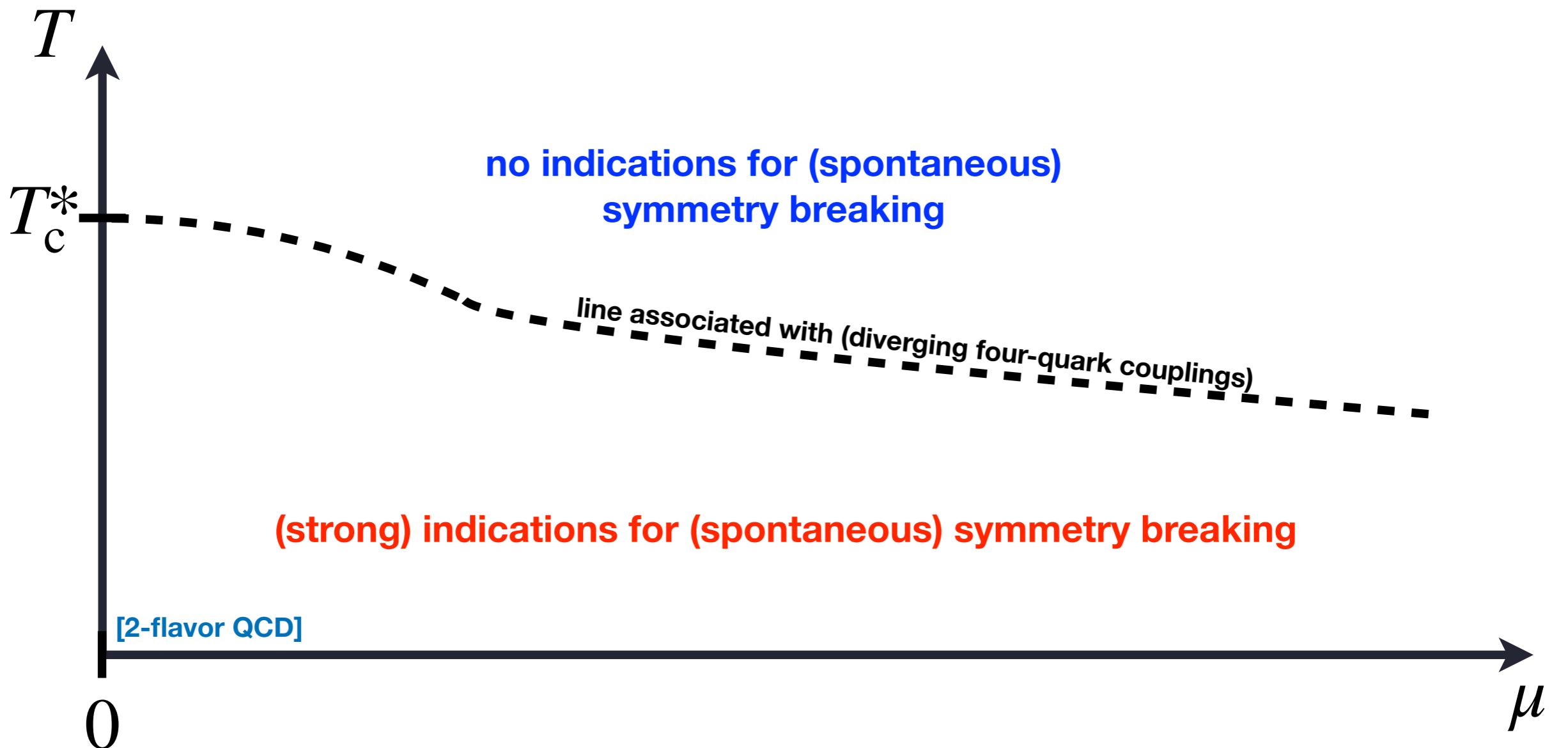
symmetry	group
color	$SU(N_c)$
chiral	$SU_L(2) \otimes SU_R(2)$
vector	$U_V(1)$
axial	$U_A(1)$
Poincare	✗
parity	✓
charge conjugation	✗
# of channels (pointlike limit)	32

Fierz-complete
set

2+1-flavor QCD

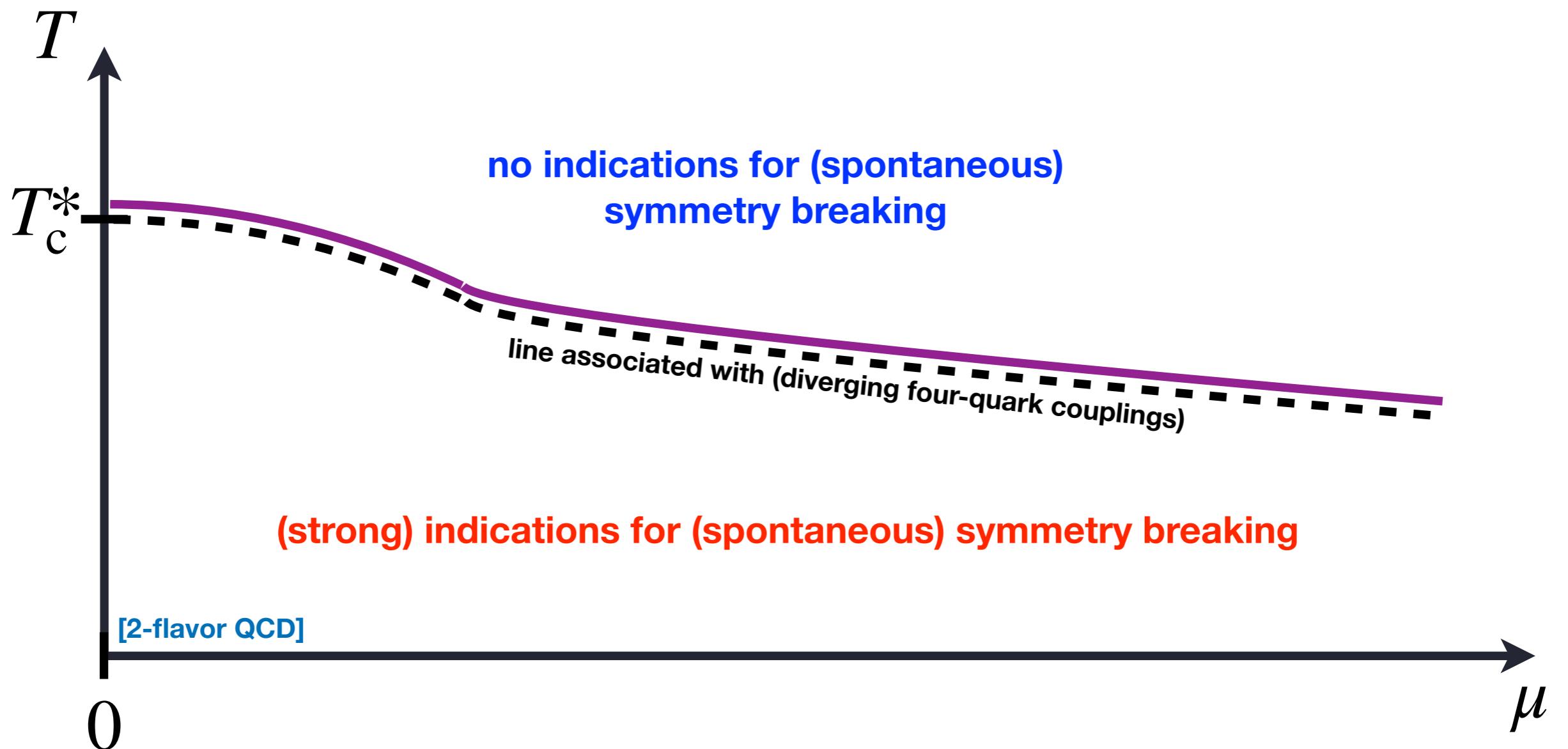
Symmetry-breaking patterns

[JB, Leonhardt, Pospiech '17 & '20]



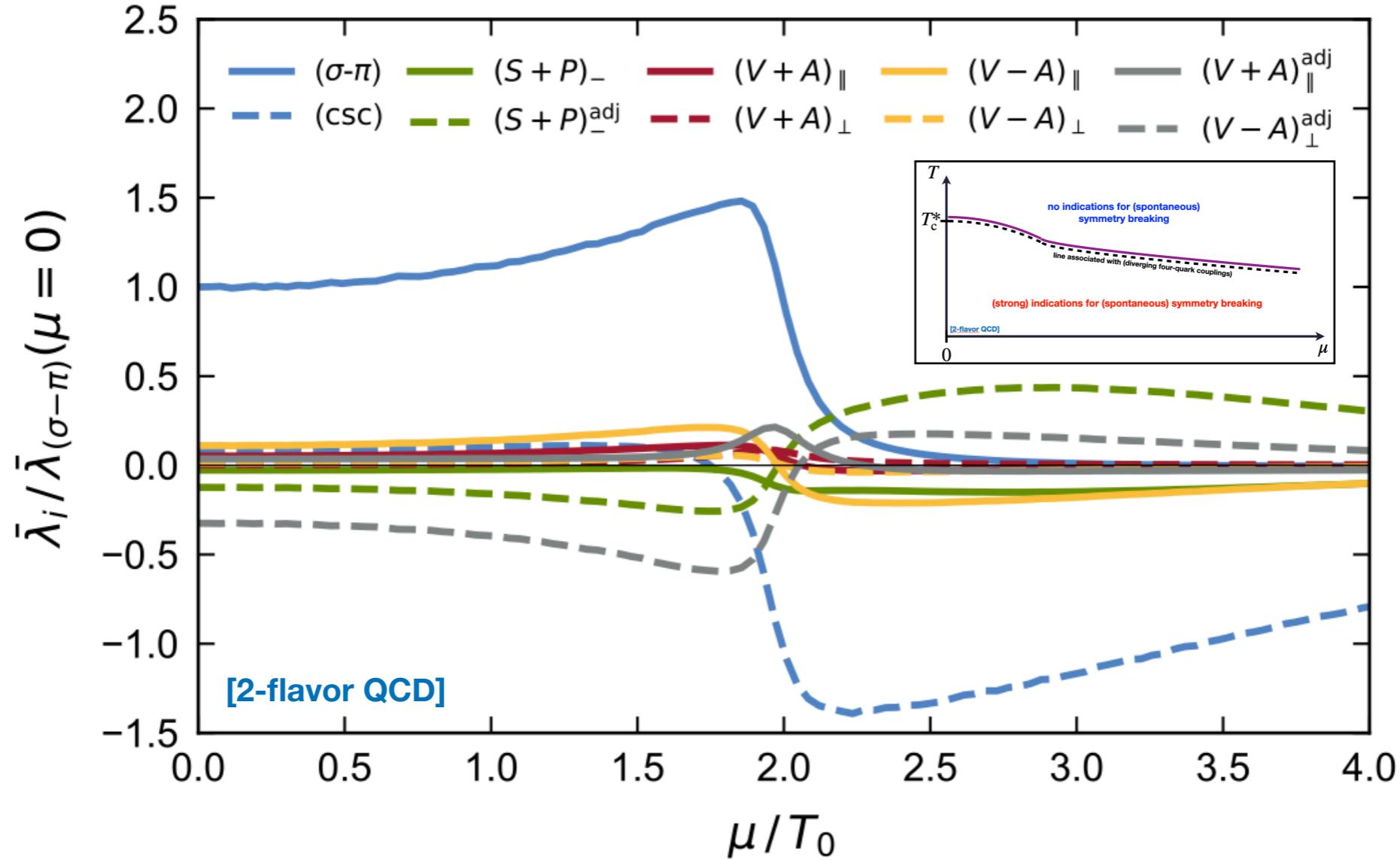
Symmetry-breaking patterns

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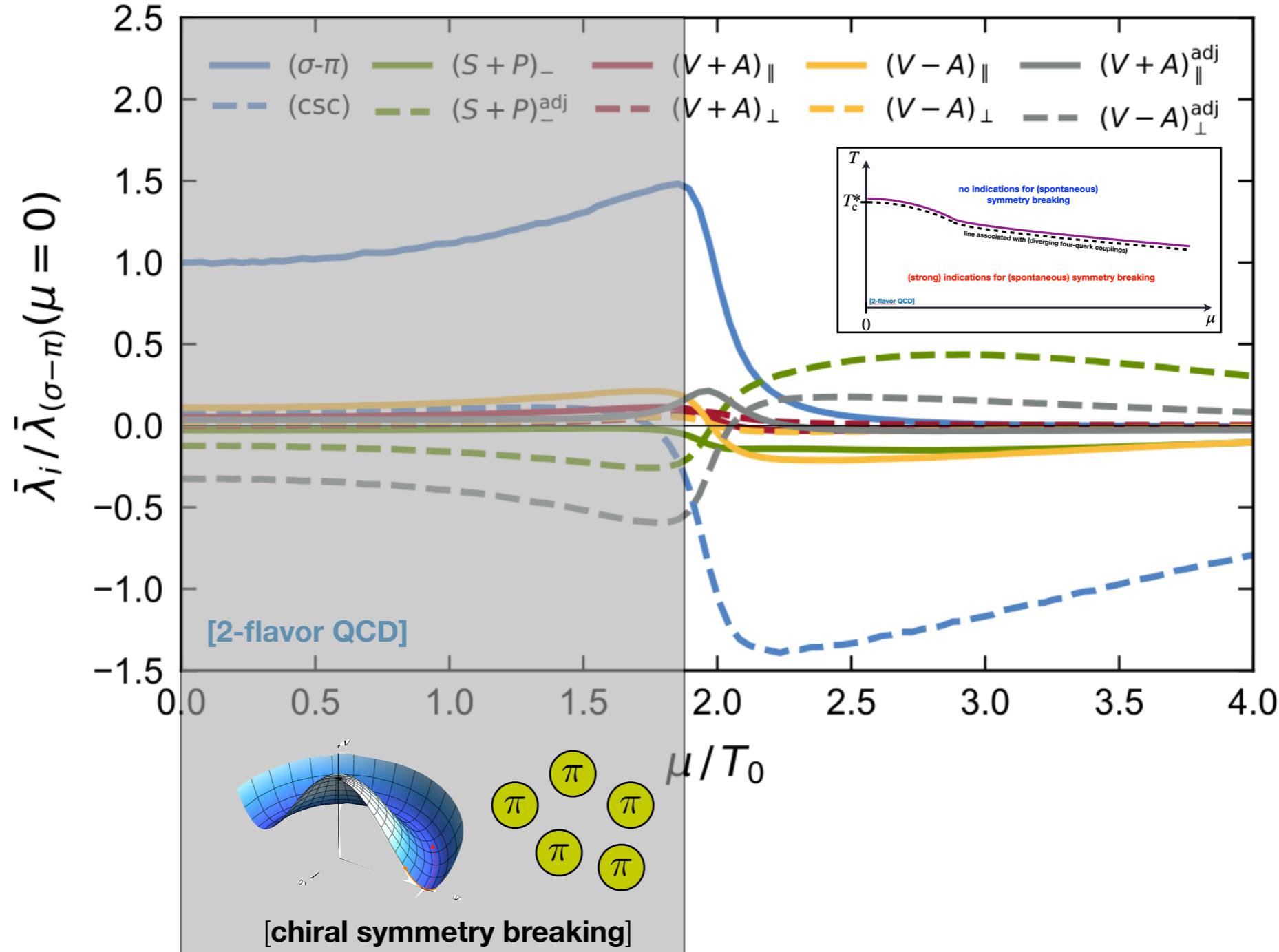
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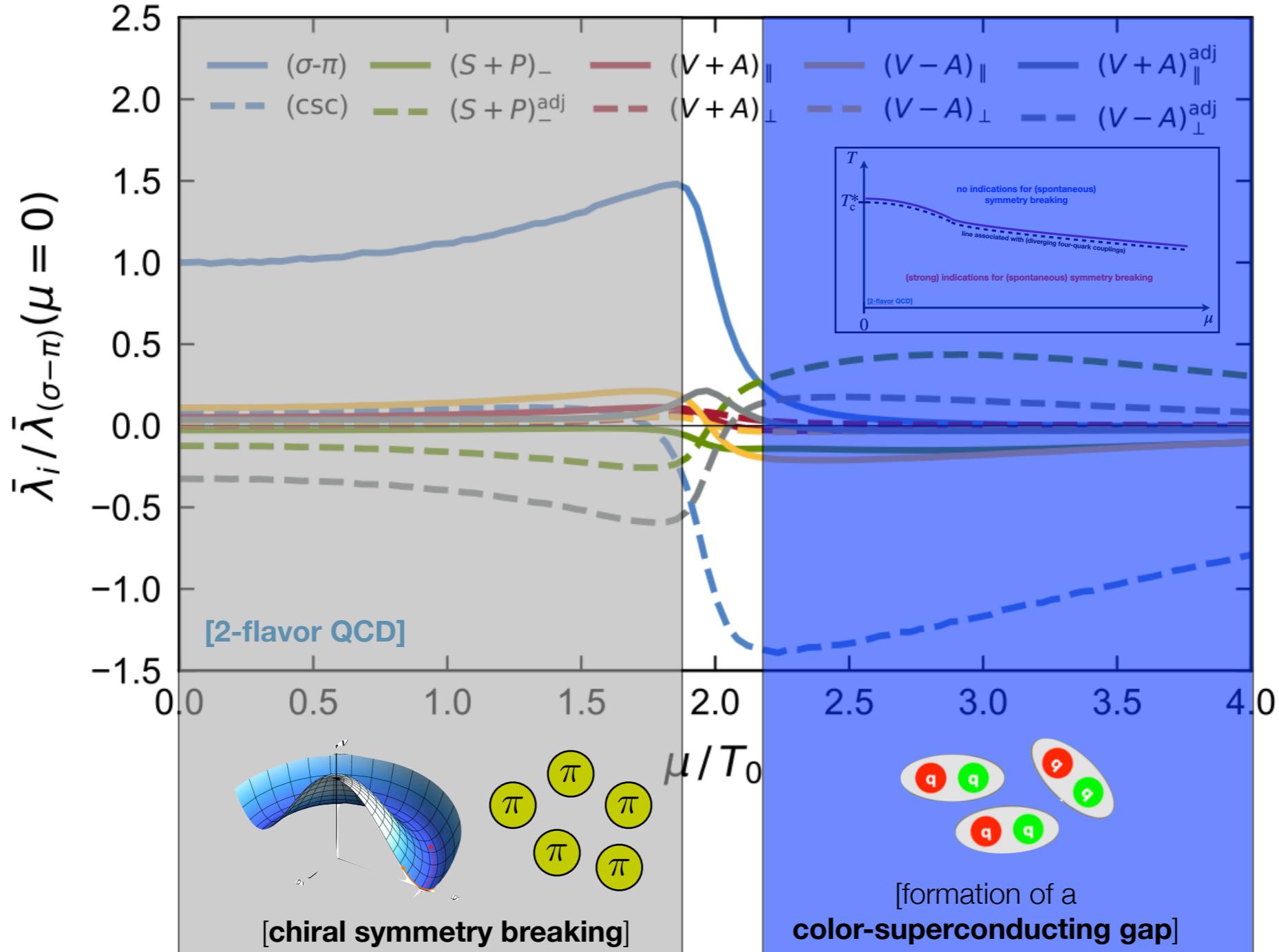
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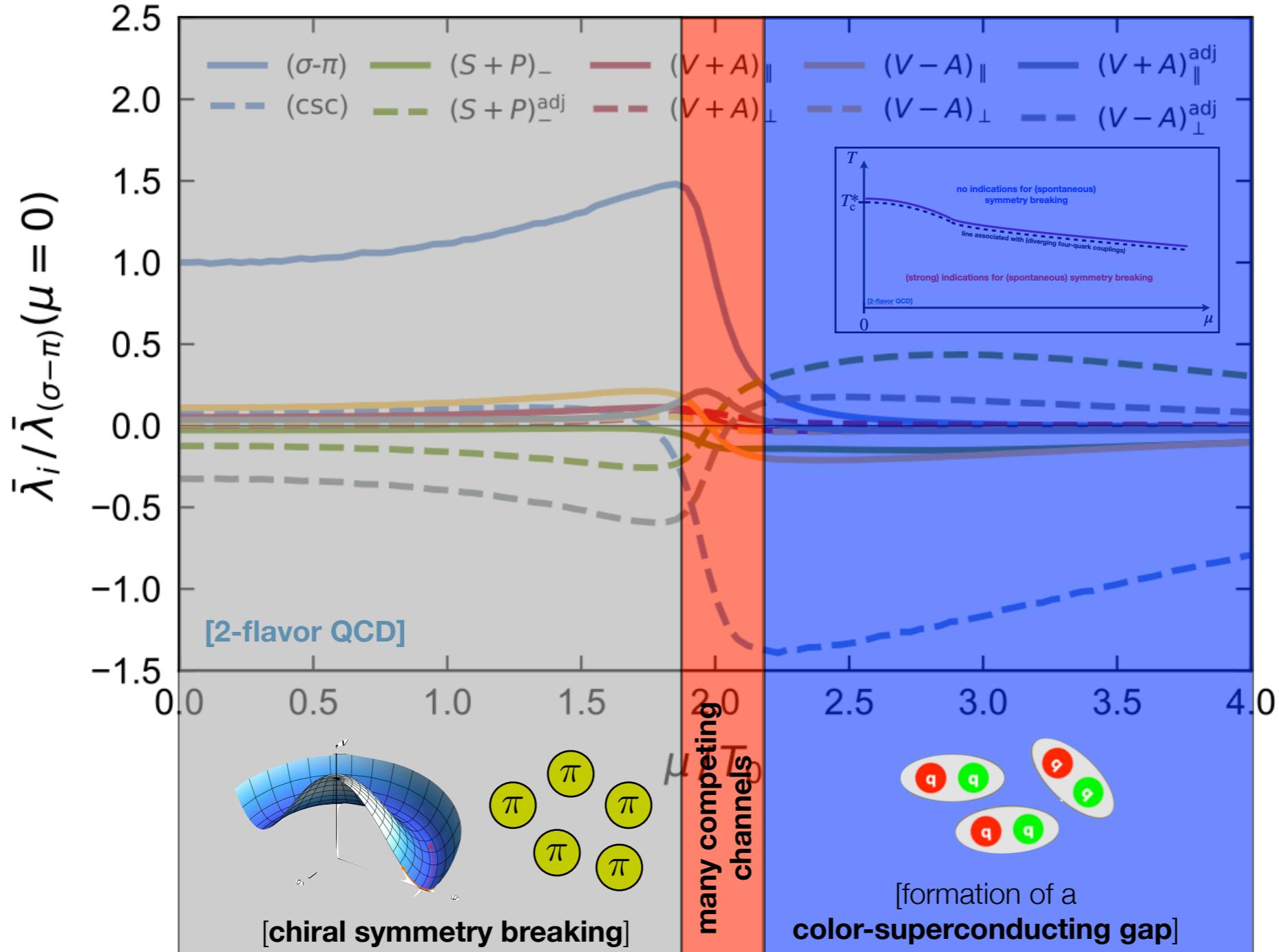
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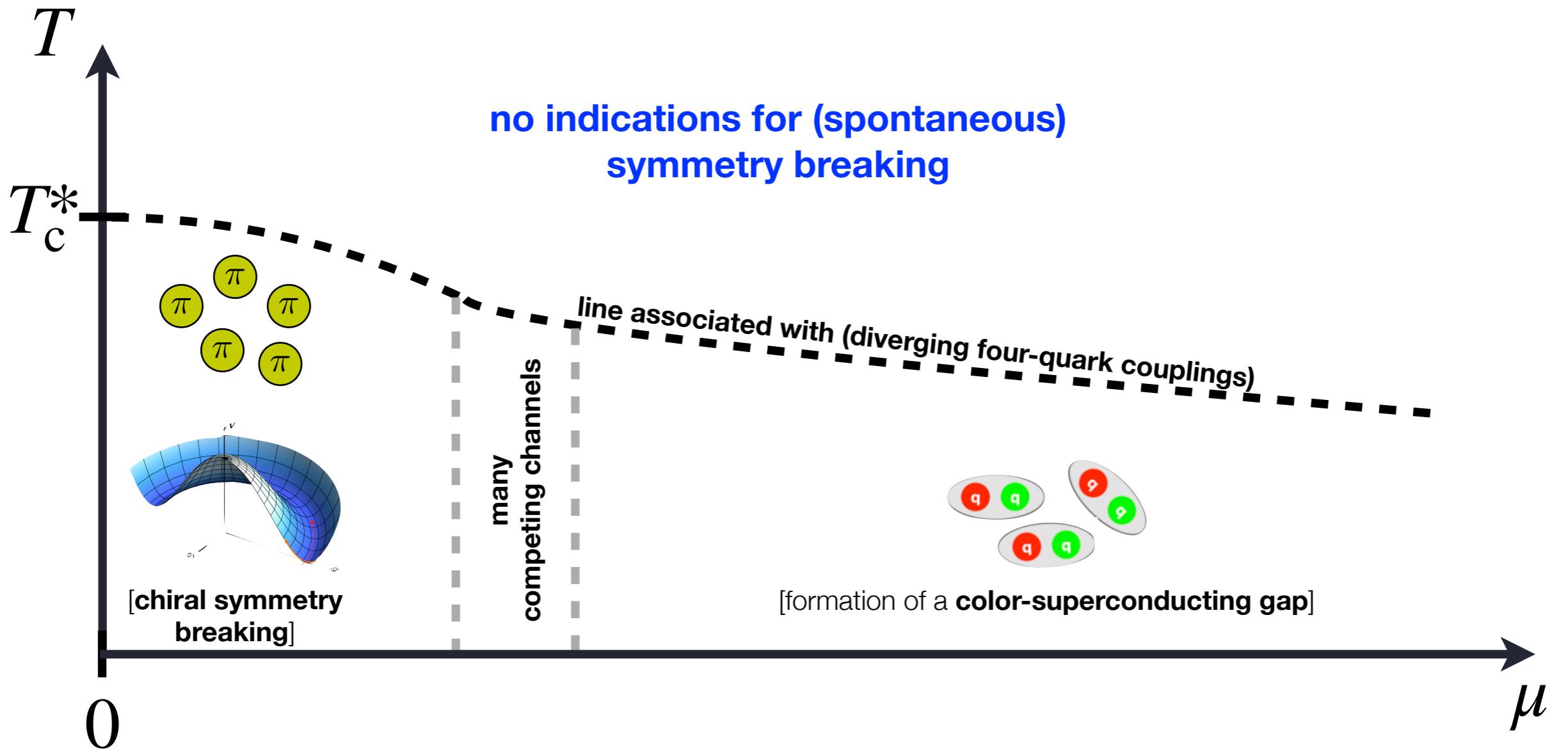
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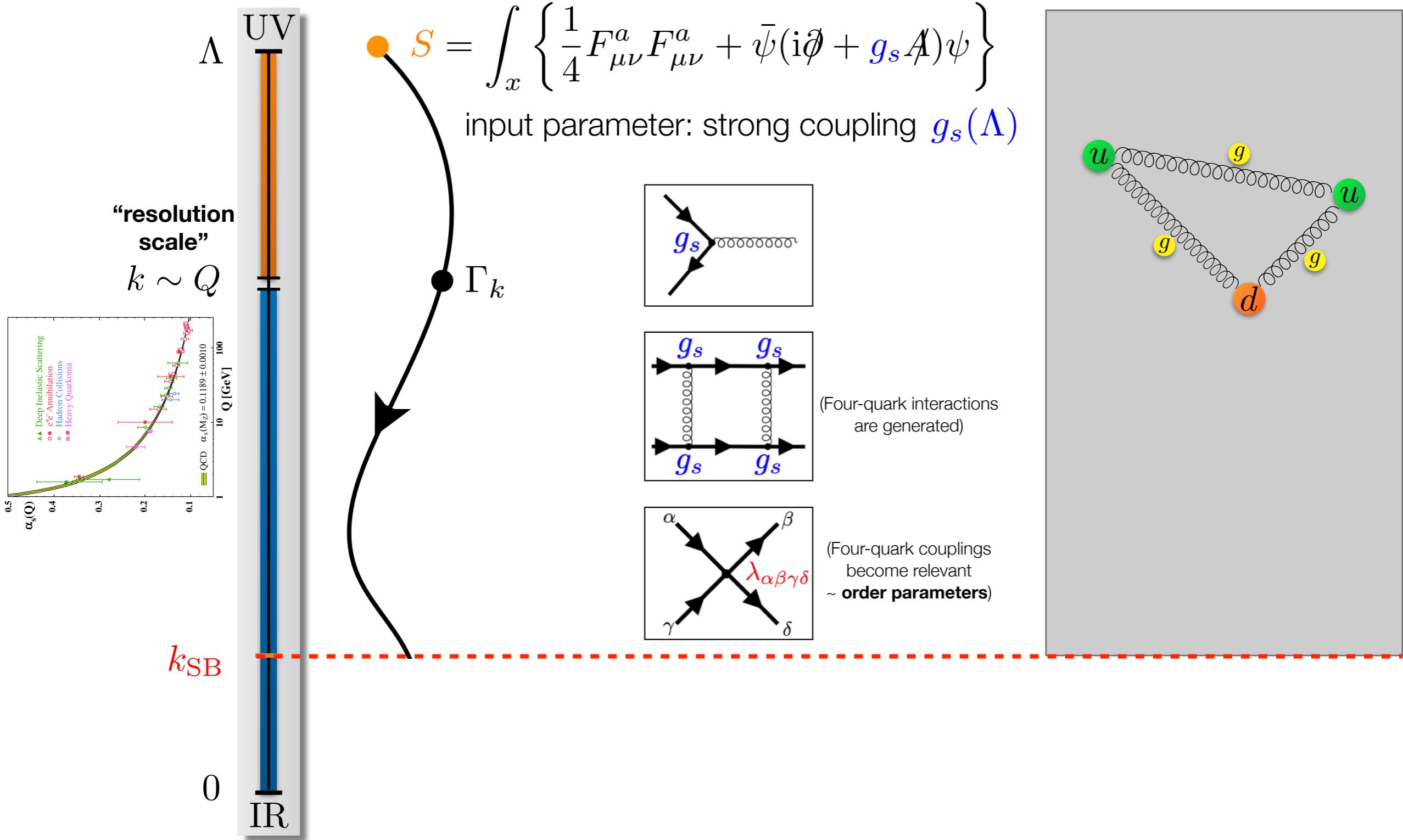
Symmetry-breaking patterns

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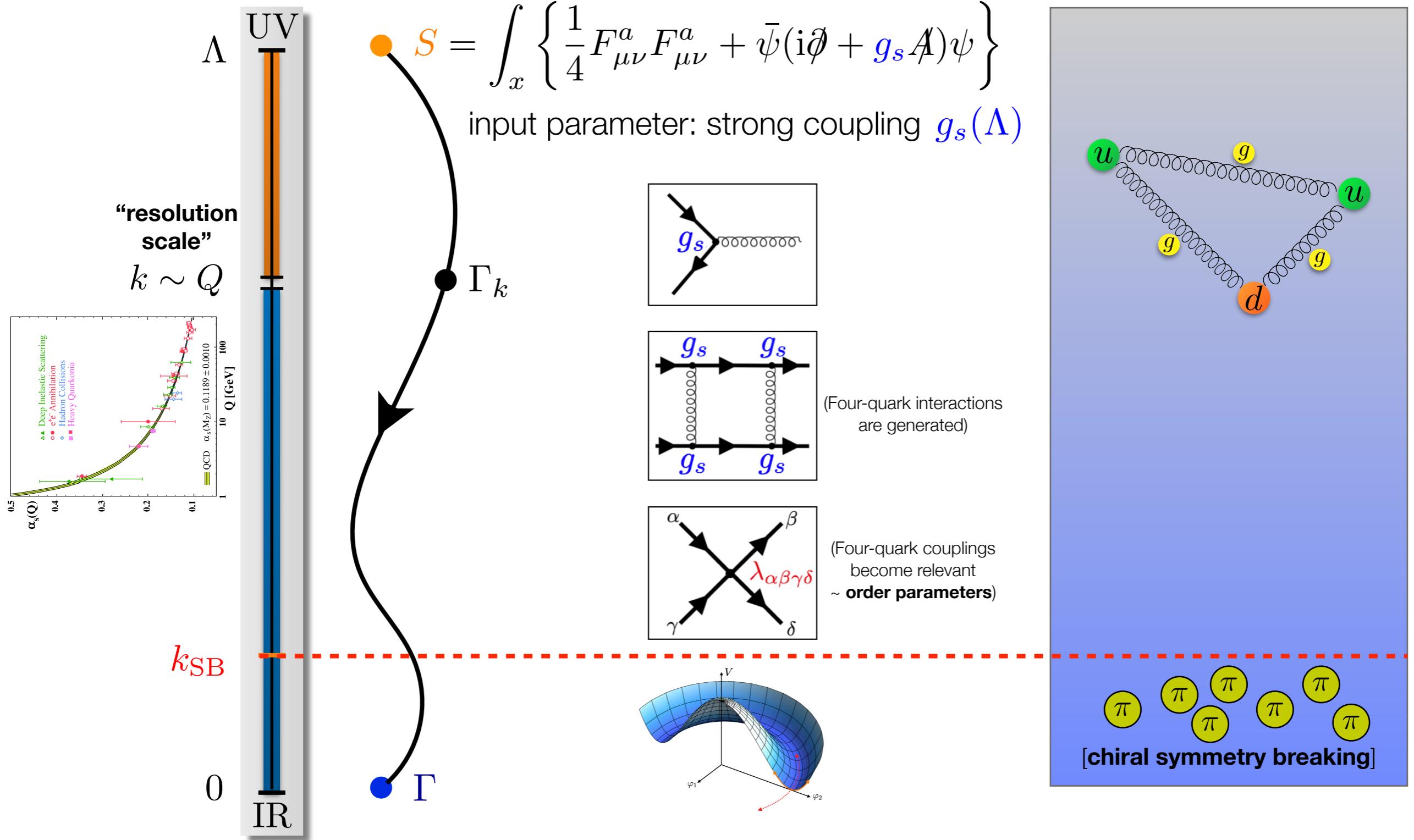
Functional Renormalization Group: Vacuum

[Wetterich '92]



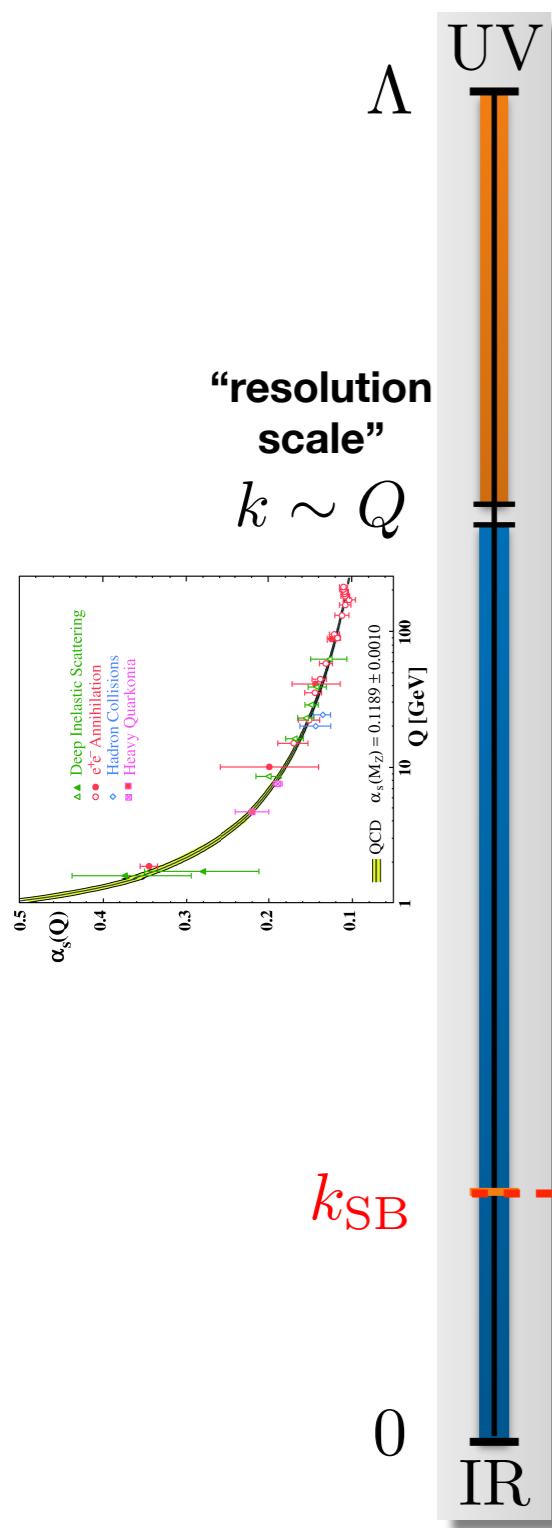
Functional Renormalization Group: Vacuum

[Wetterich '92]



Functional Renormalization Group: High density

[JB & Schallmo '21]



$$S = \int_x \left\{ \frac{1}{4} F_{\mu\nu}^a F_{\mu\nu}^a + \bar{\psi} (i\partial + g_s A) \psi \right\}$$

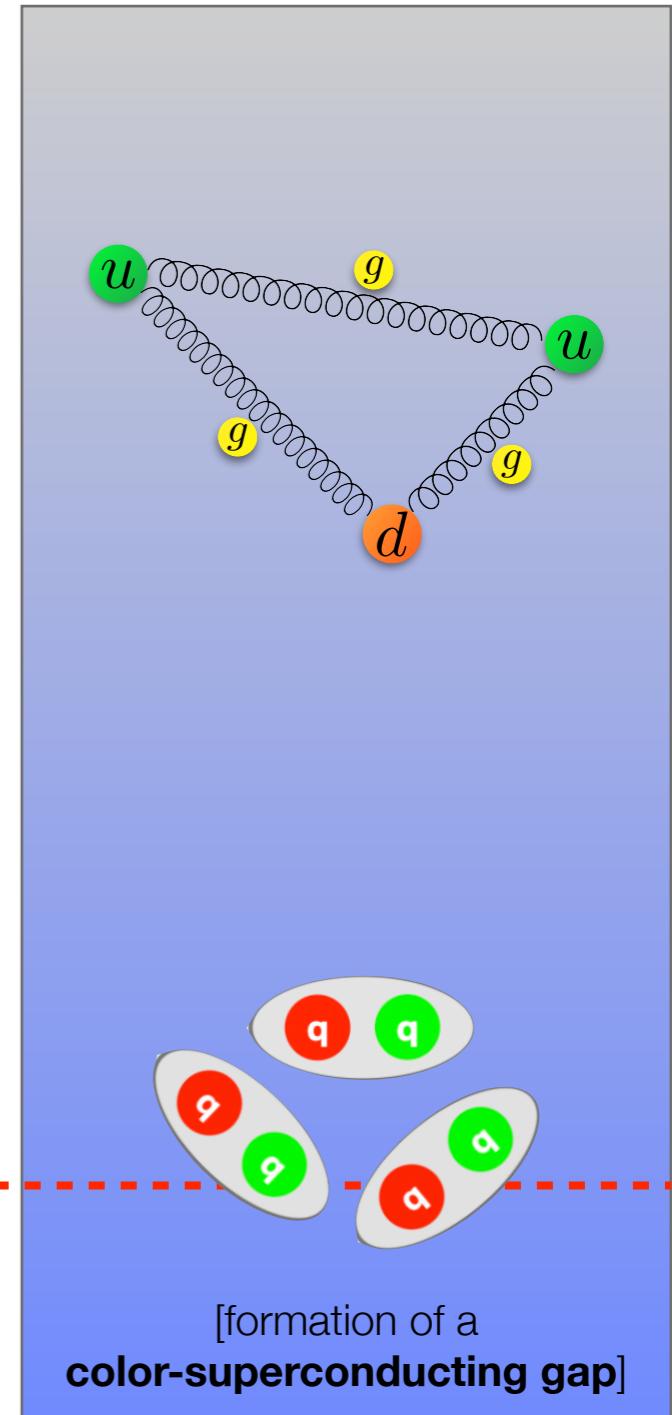
input parameter: strong coupling $g_s(\Lambda)$

$\mu > 0$

(Four-quark interactions are generated)

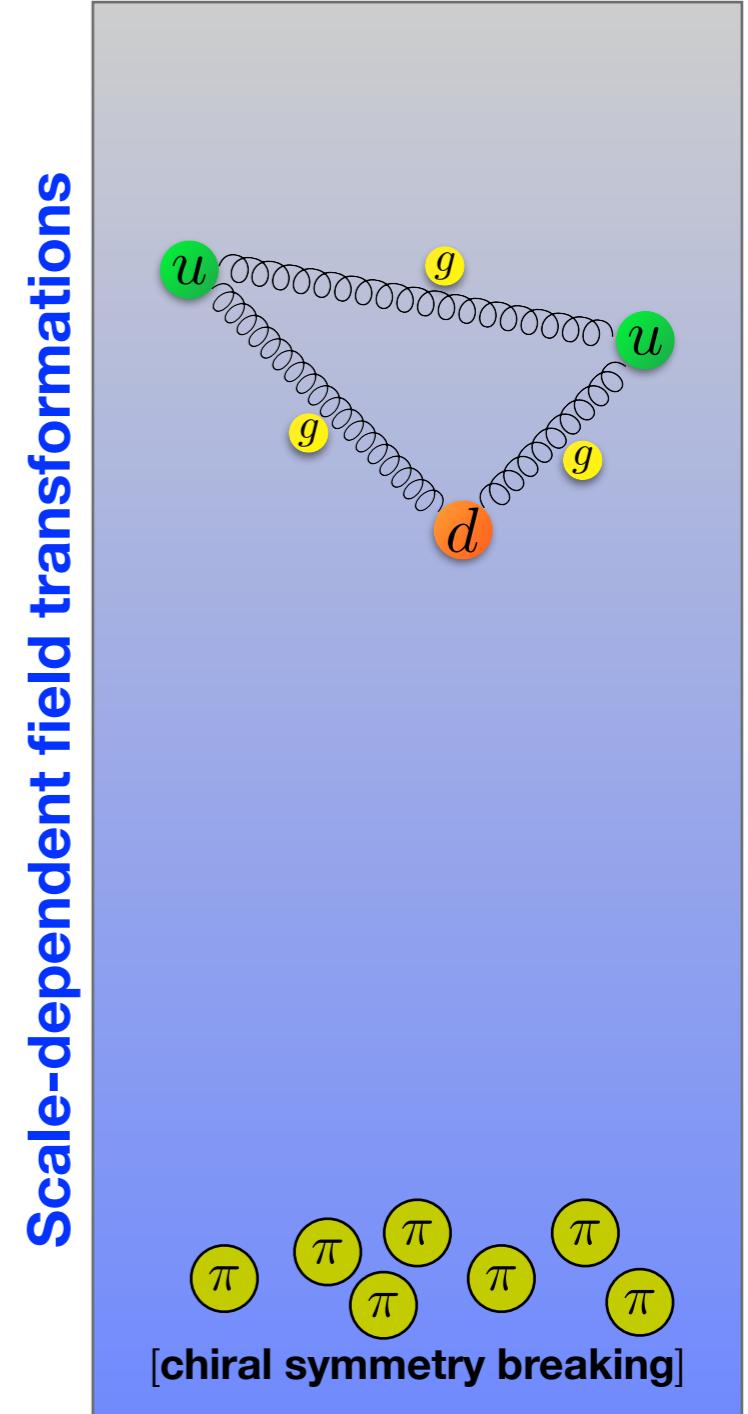
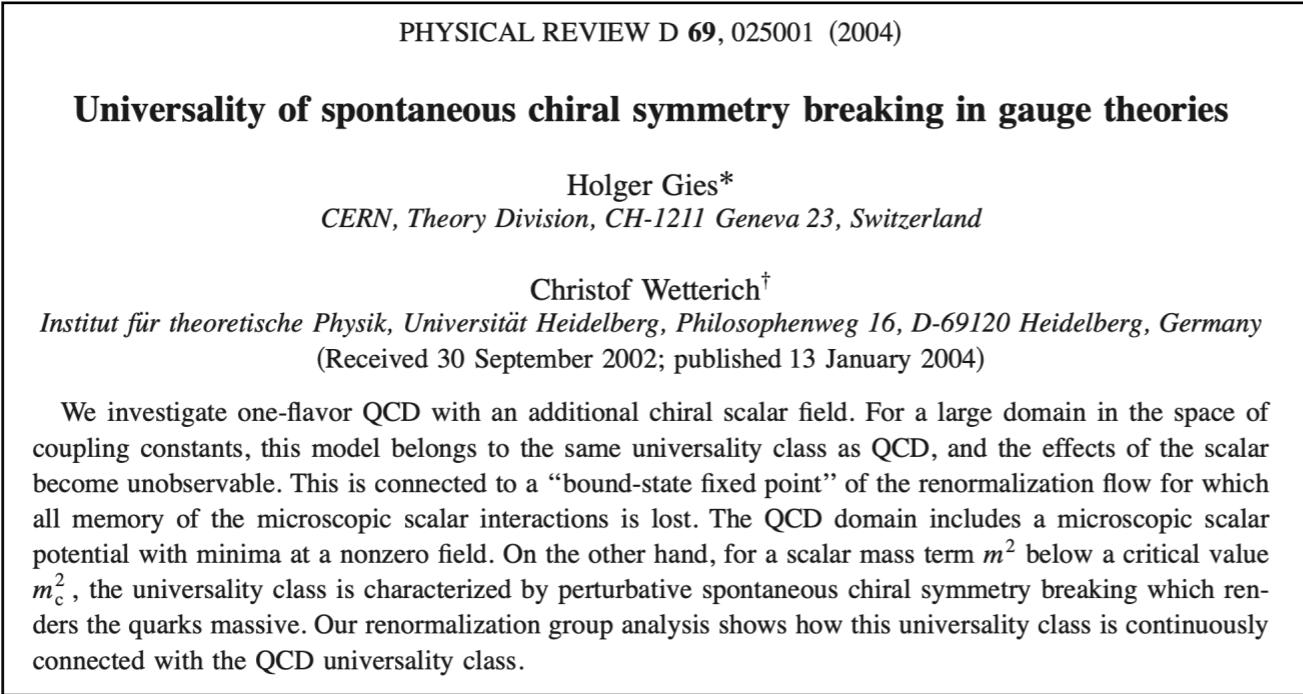
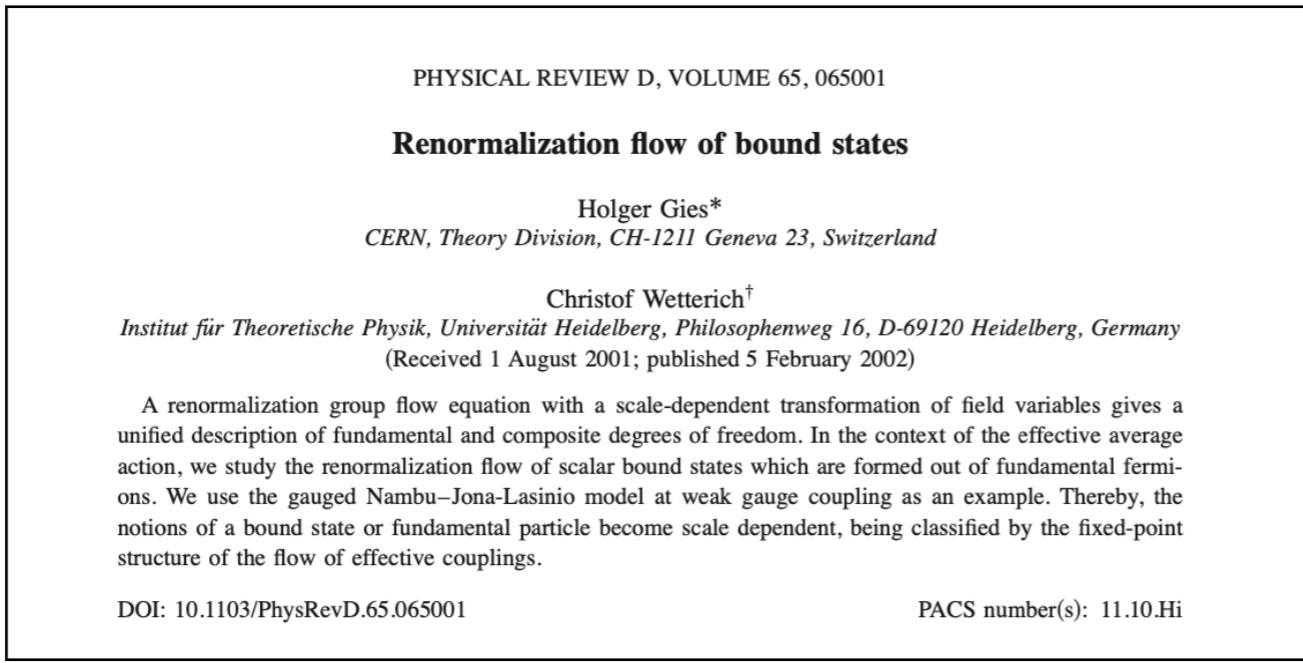
(Four-quark couplings become relevant ~ order parameters)

$\Gamma(\mu > 0)$



Continuous transition in the degrees of freedom

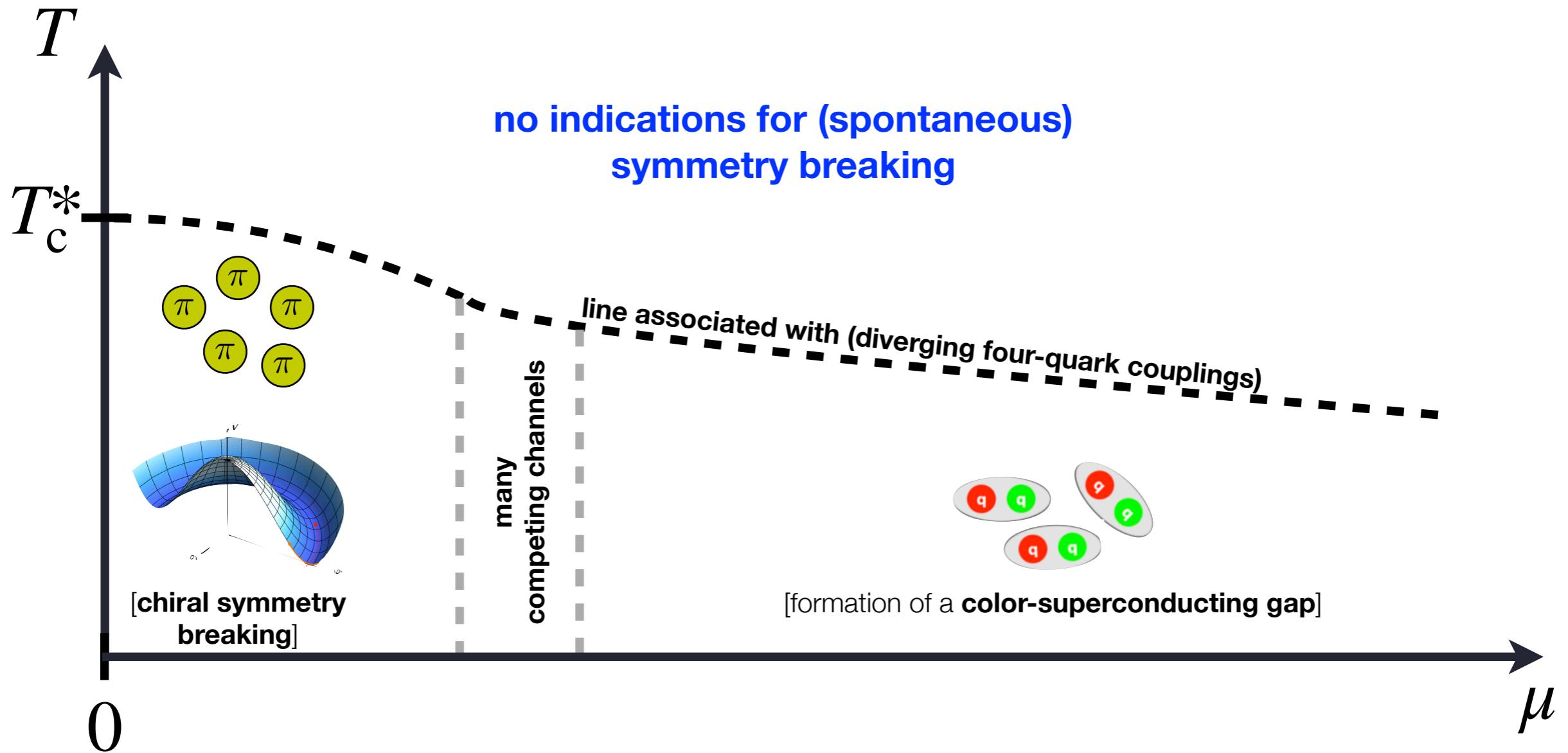
Continuous transition in the degrees of freedom



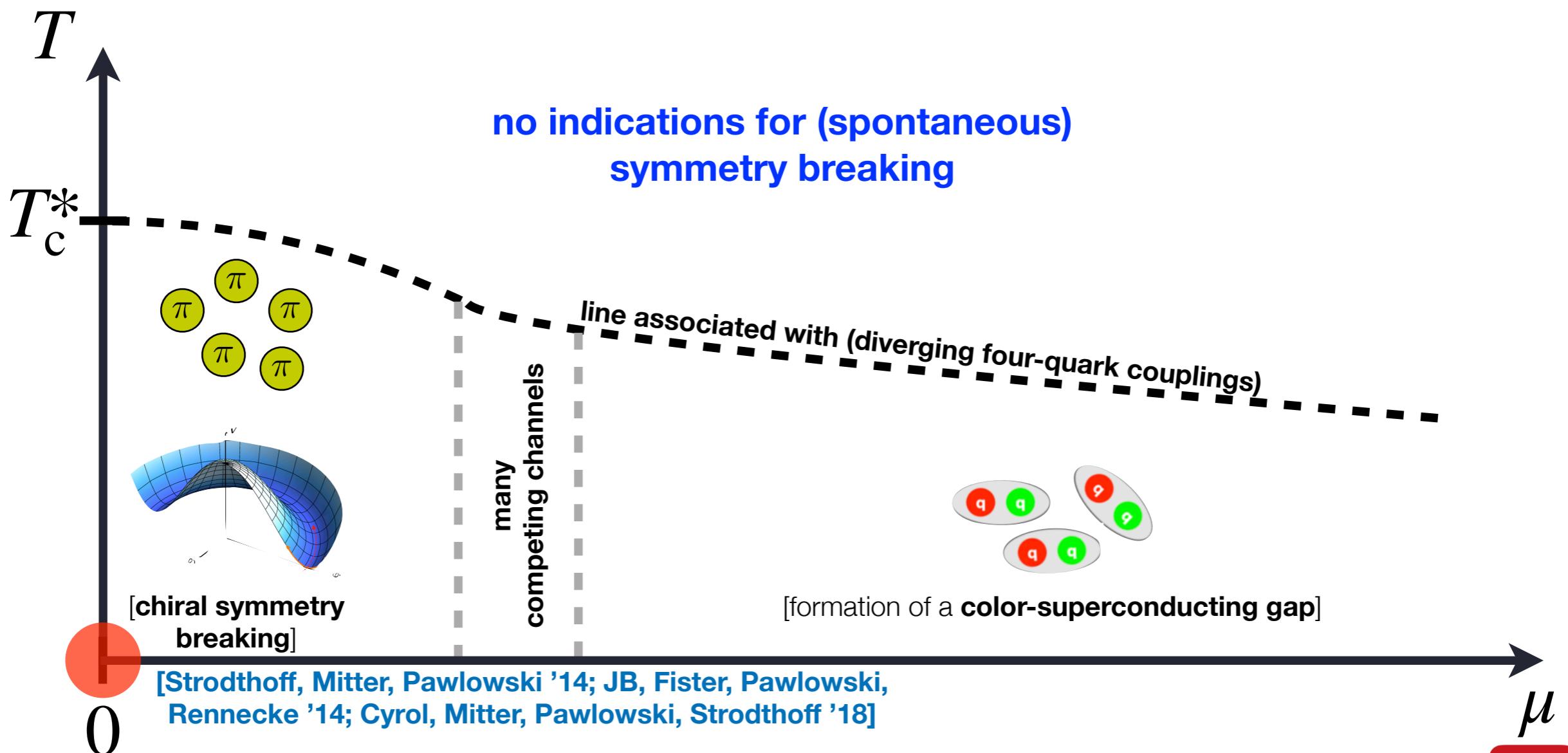
[see also Pawłowski '05; Floerchinger & Wetterich '09;
JB, Fister, Pawłowski, Rennecke '14; Strodthoff, Fukushima, Pawłowski '21]

Continuous transition in the degrees of freedom

Phase structure



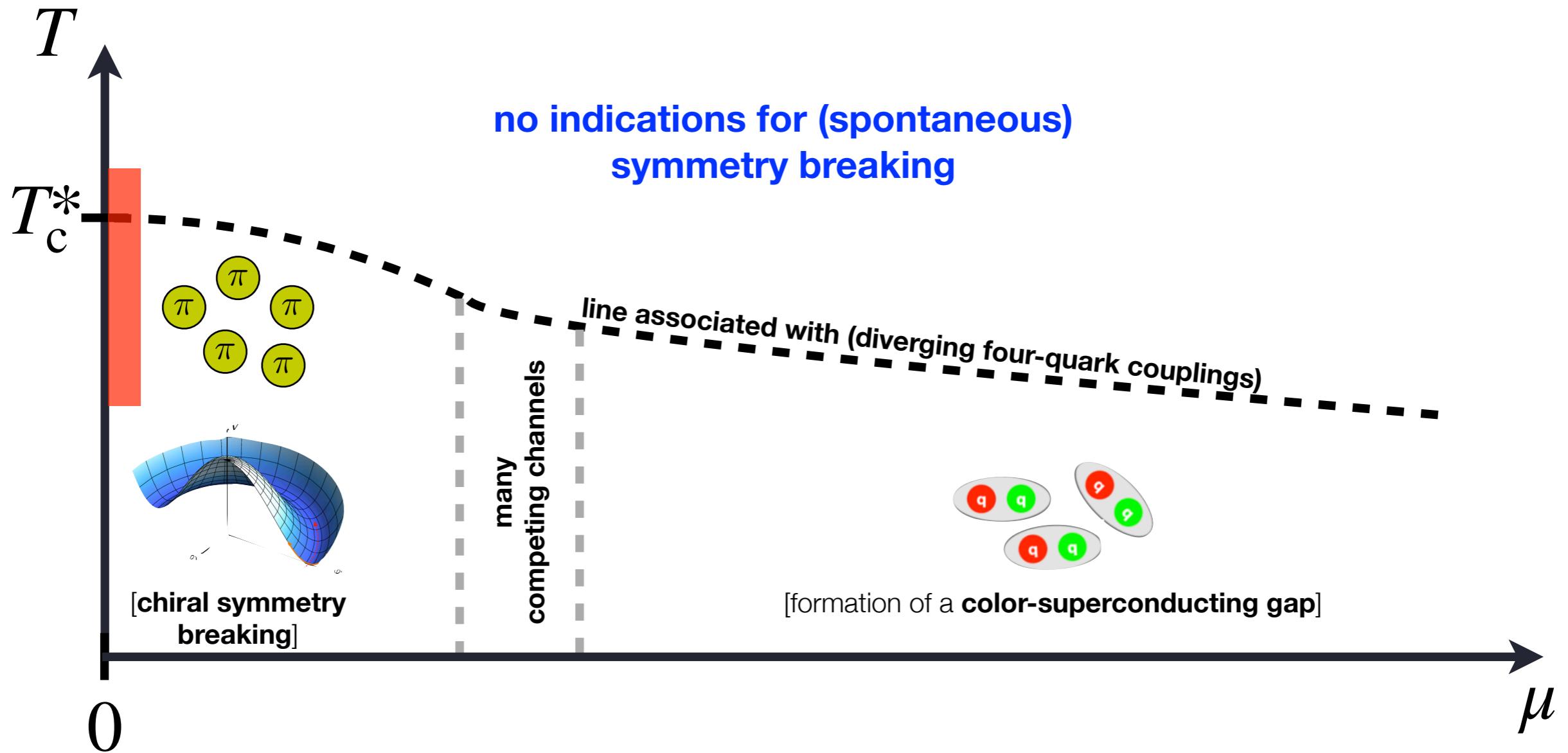
Phase structure: Vacuum



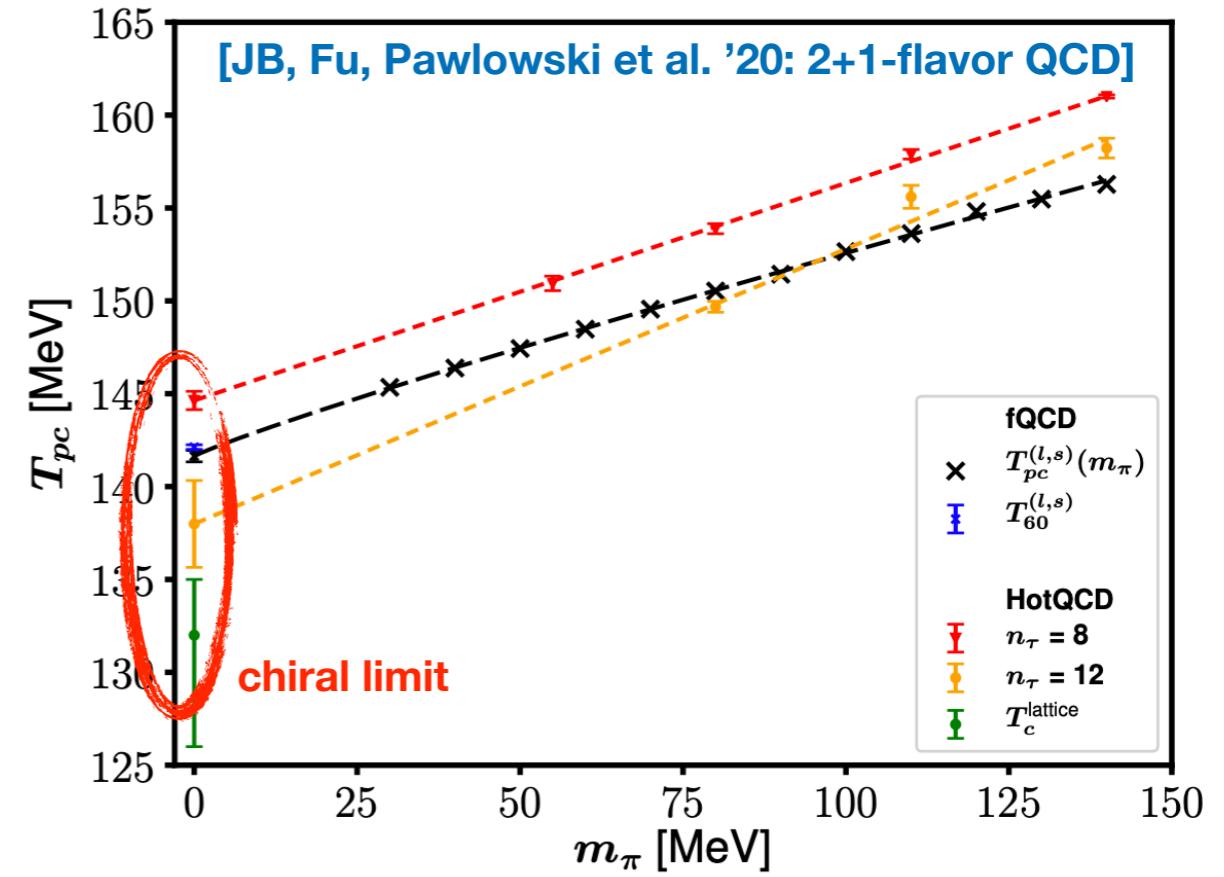
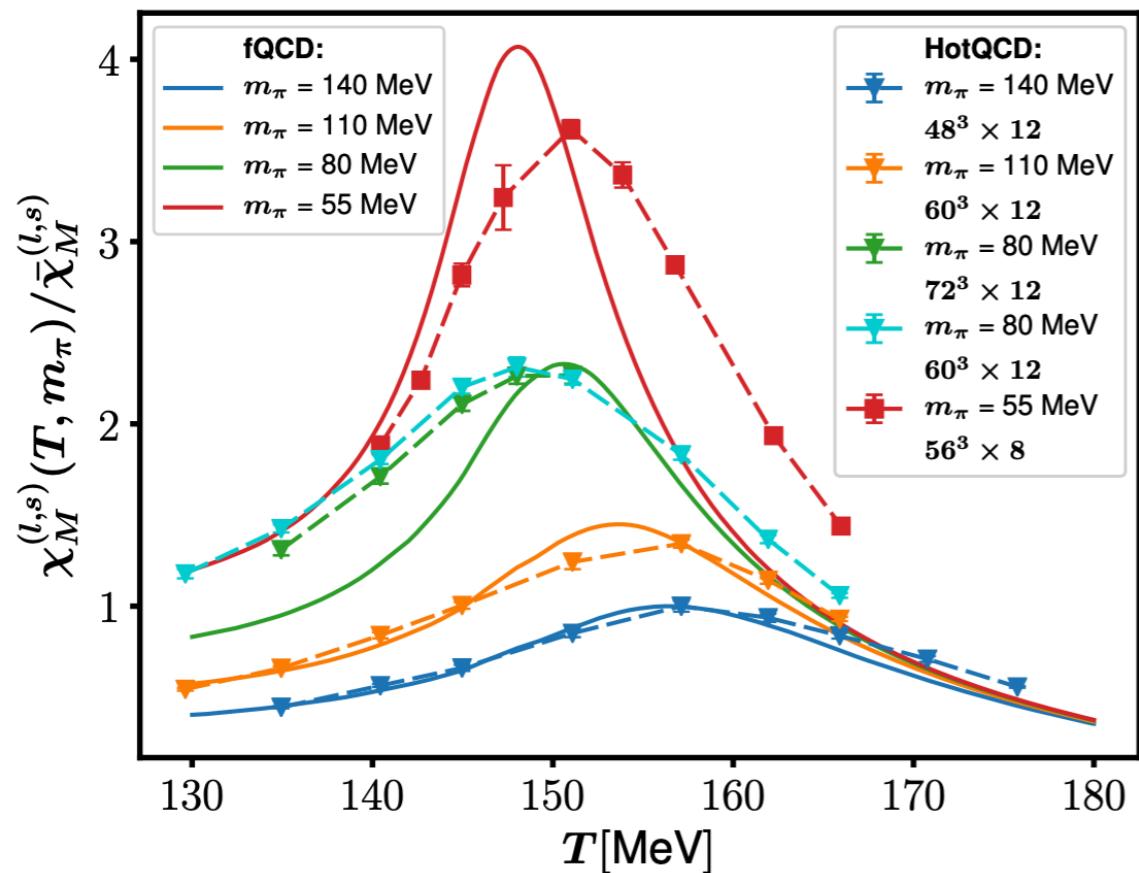
- **Impressive progress** on description of scattering processes [Eser & Blaizot '22] & spectral functions [Horak et al. '22]

ERG
2022
ERG
2022

Phase structure

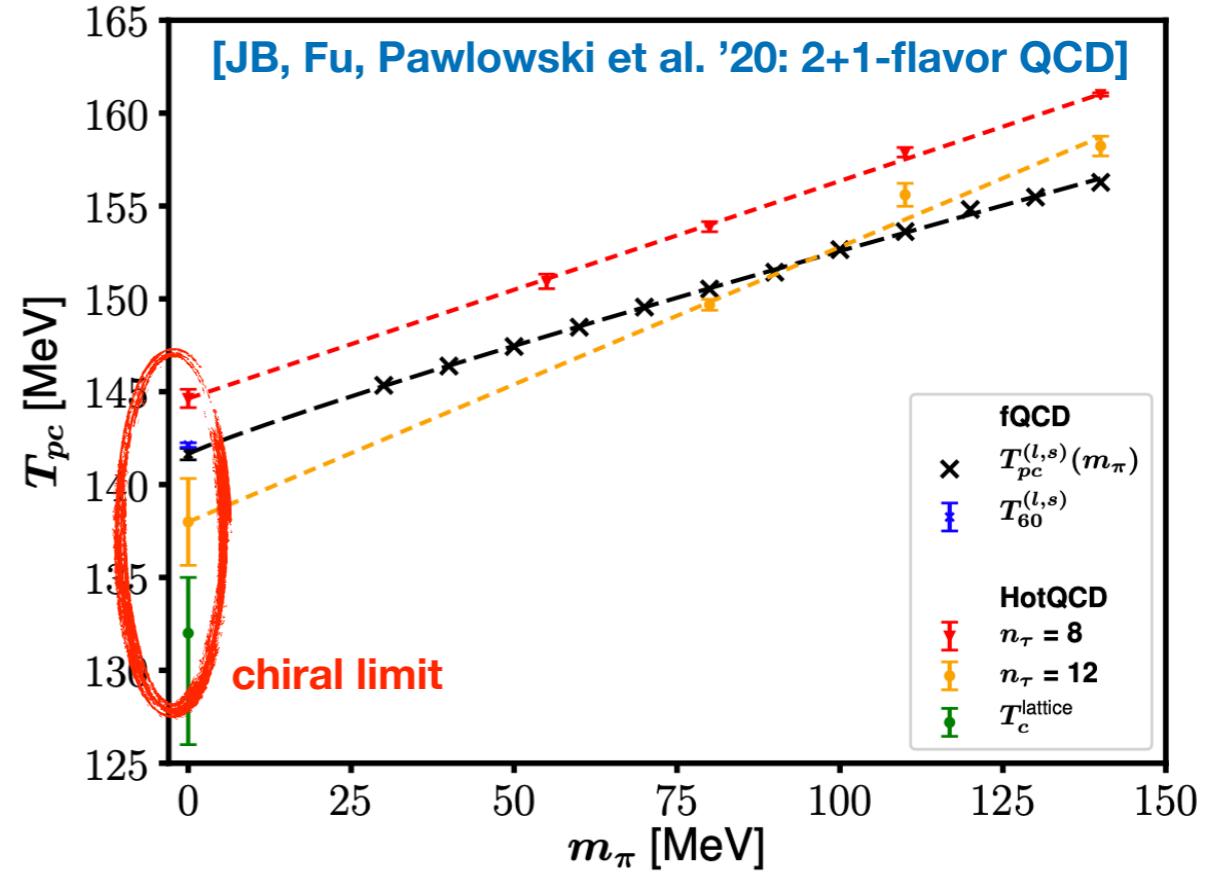
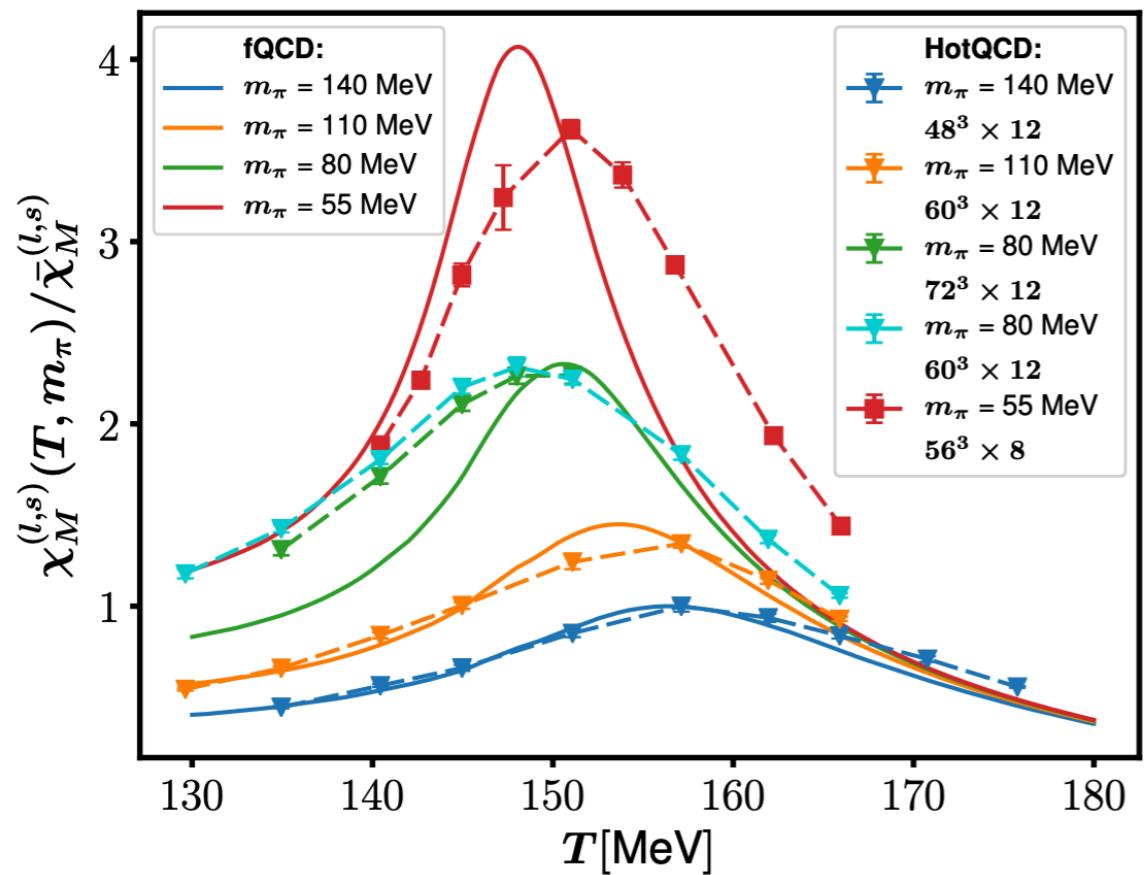


Nature of the chiral transition



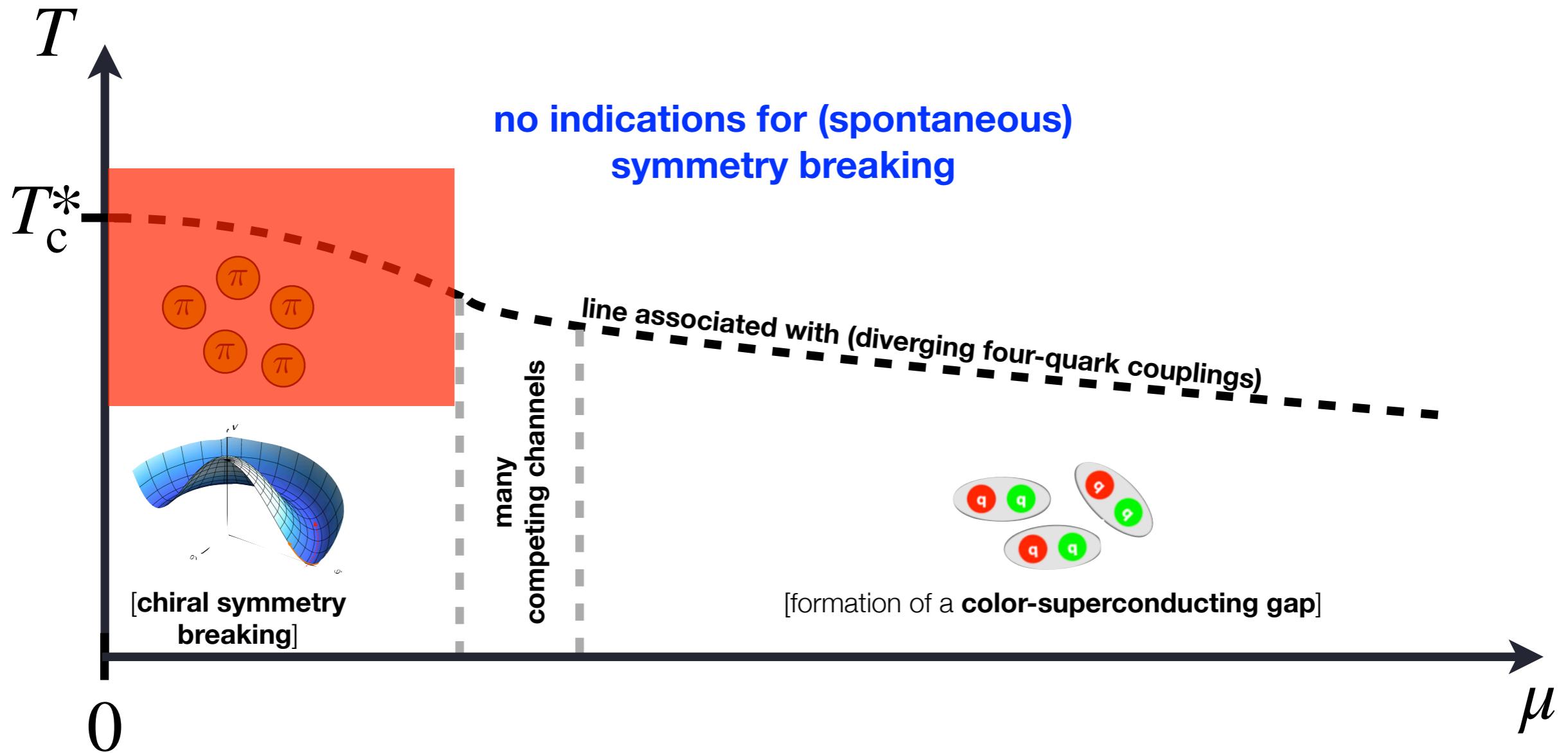
- Good agreement with lattice QCD: $T_c \sim 130 \dots 140 \text{ MeV}$
[Ding et al. '19; JB, Fu, Pawłowski et al. '20; Kotov, Lombardo, Trunin '21]
- Phase transition appears to fall into the O(4) universality class [JB, Fu, Pawłowski et al. '20; JB, Pawłowski, Rosenblüh '21]

Nature of the chiral transition

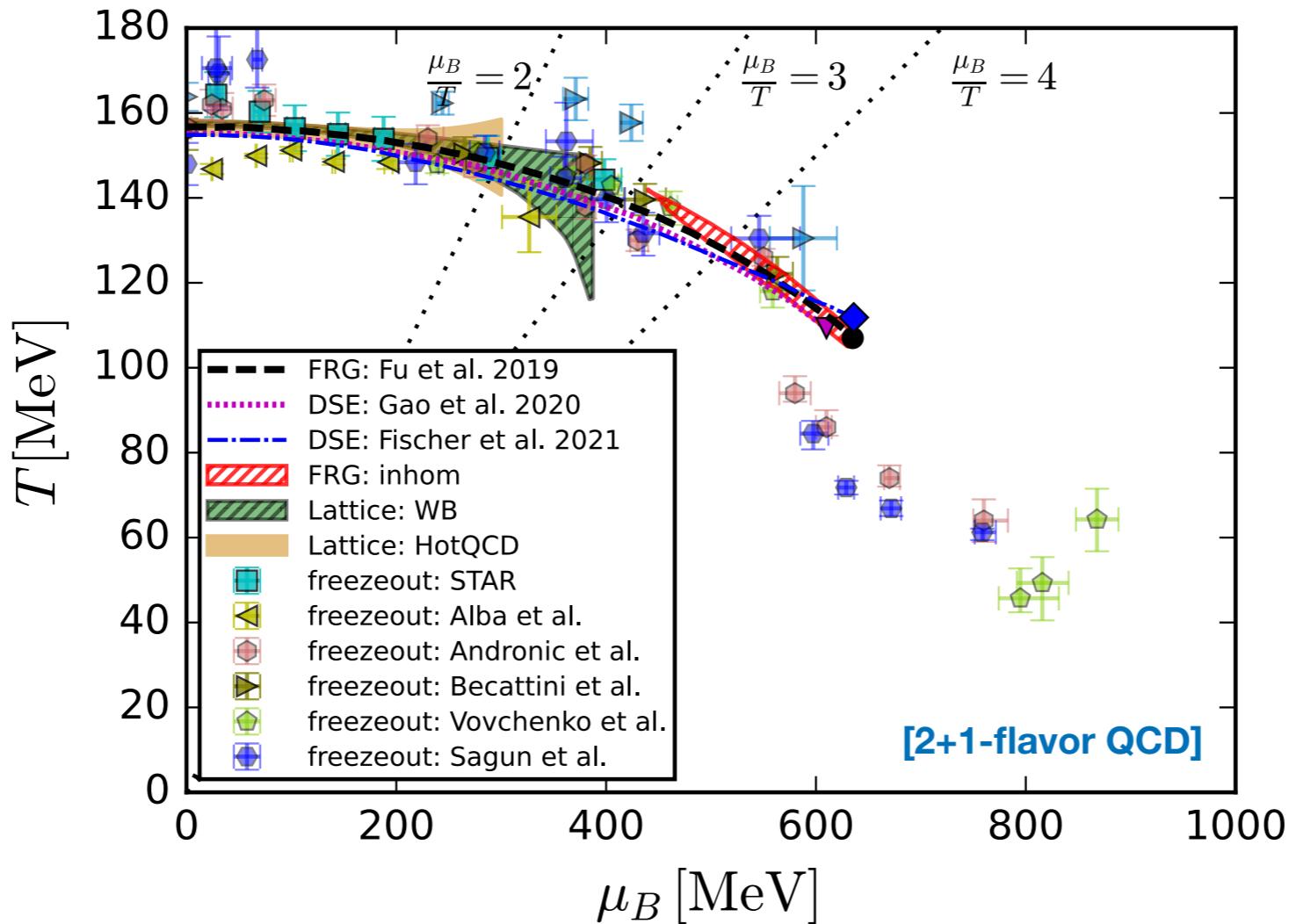


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- Scenarios: Nature of the transition in 3-flavor QCD?
[Fejos '22]

Phase structure

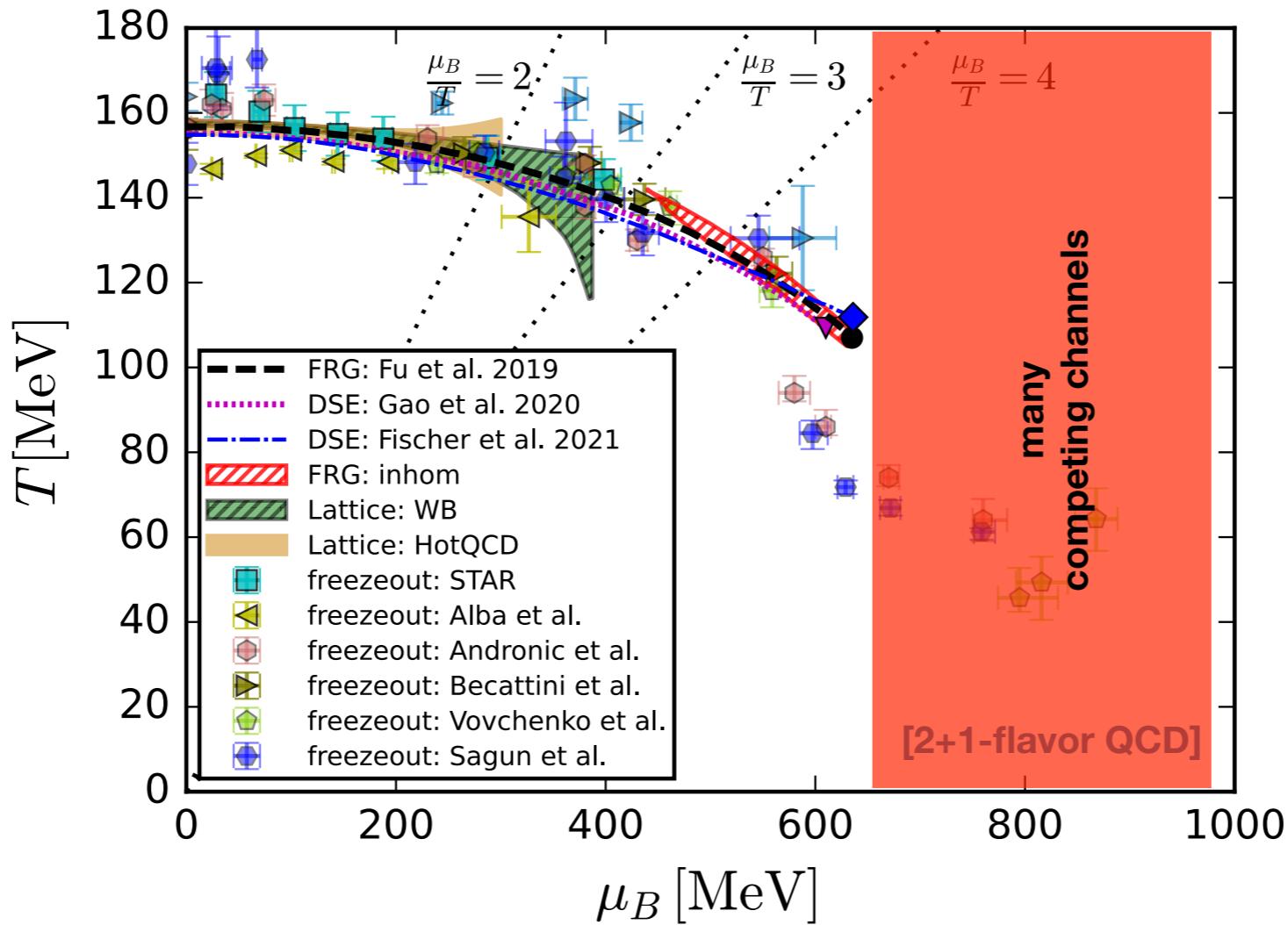


Phase structure at small chemical potential



- In very good agreement with lattice QCD results at small chemical potential [\[Fu, Pawłowski, Rennecke '20\]](#)
- **Prediction** for the critical point [\[Fu, Pawłowski, Rennecke '20\]](#)

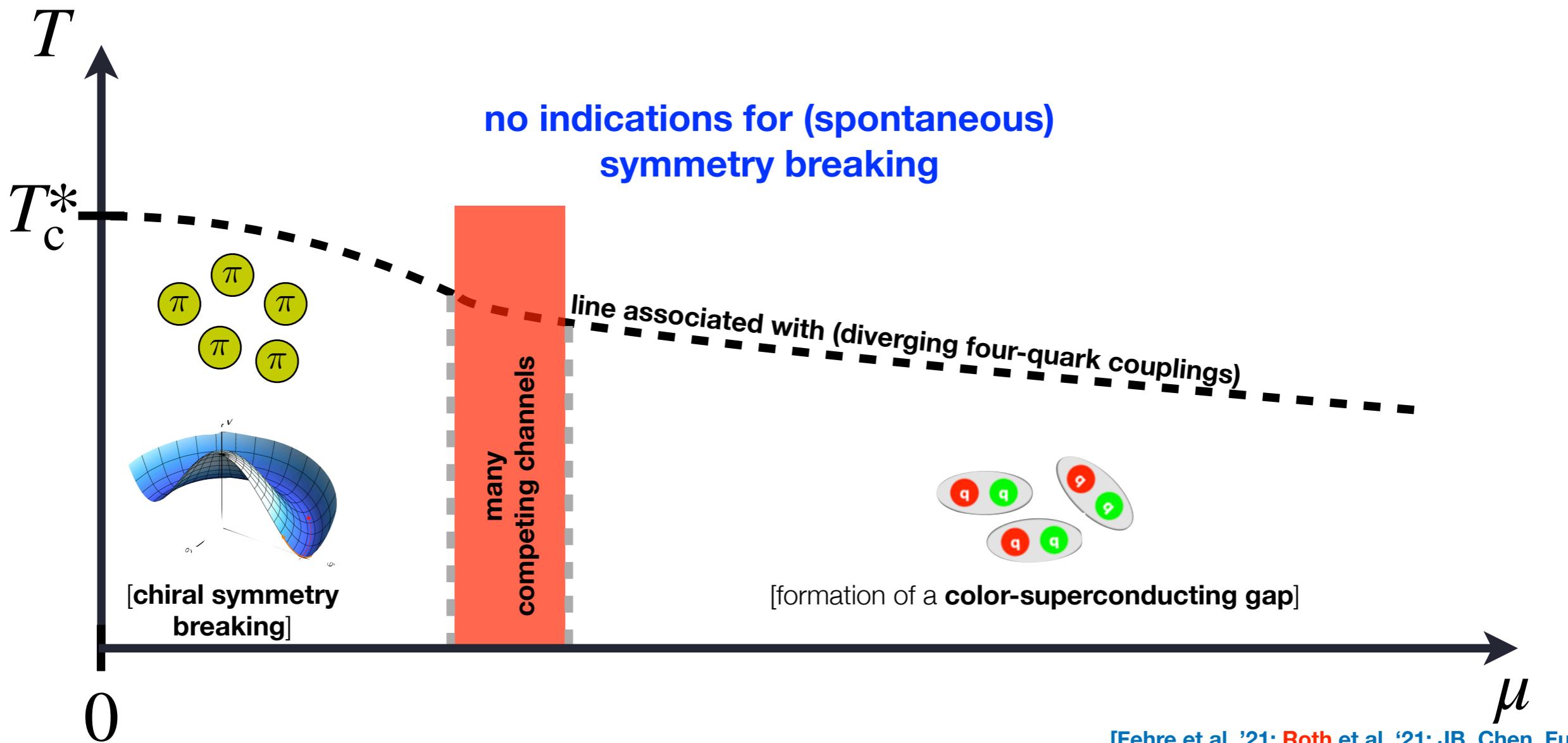
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Phase structure

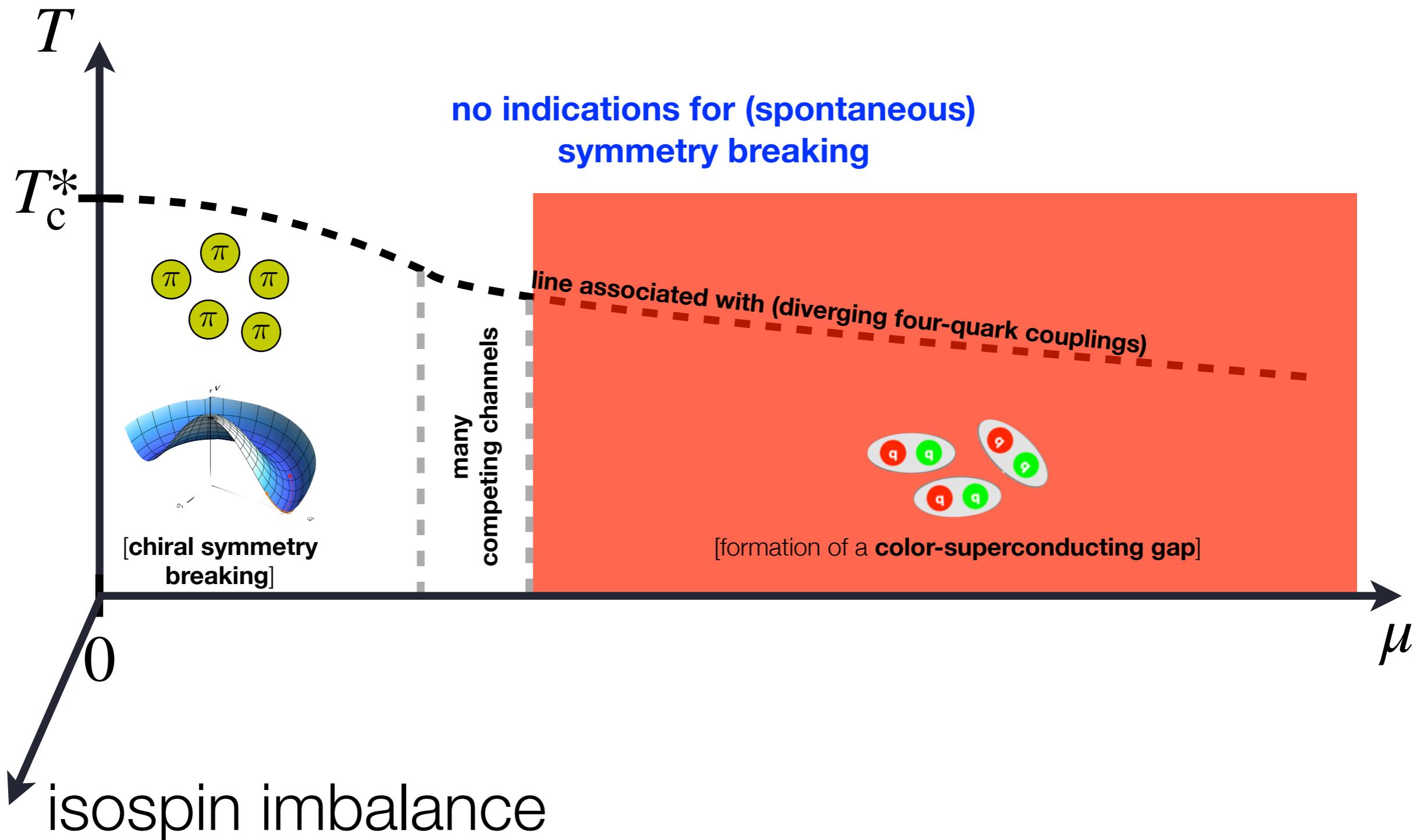
ERG
2022



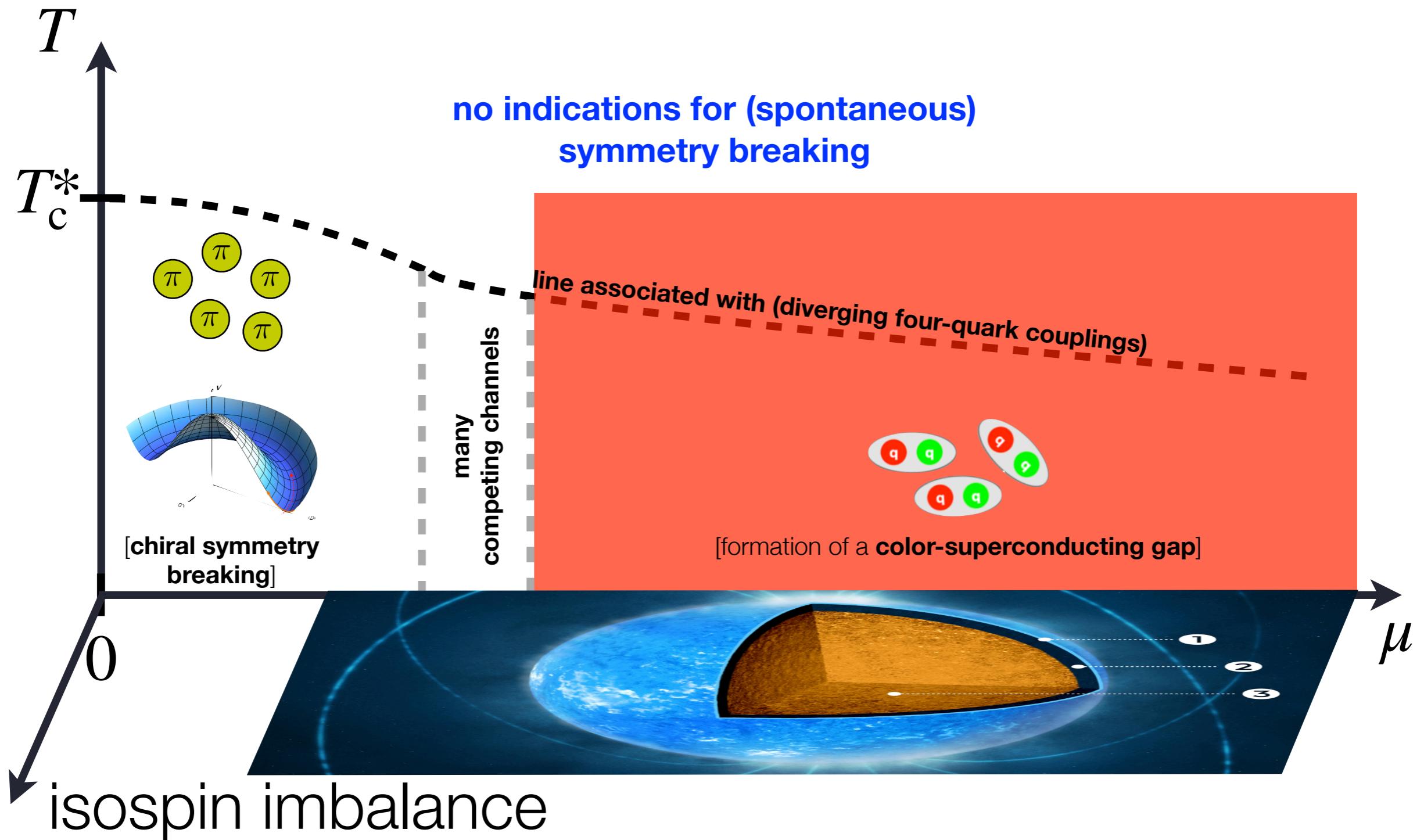
- Promising developments of regulators and numerical techniques

[Wink & Grossi '19; Koenigstein et al. '21; Stoll, Zorbach, Koenigstein et al. '21; Grossi, Ihssen et al. '22]

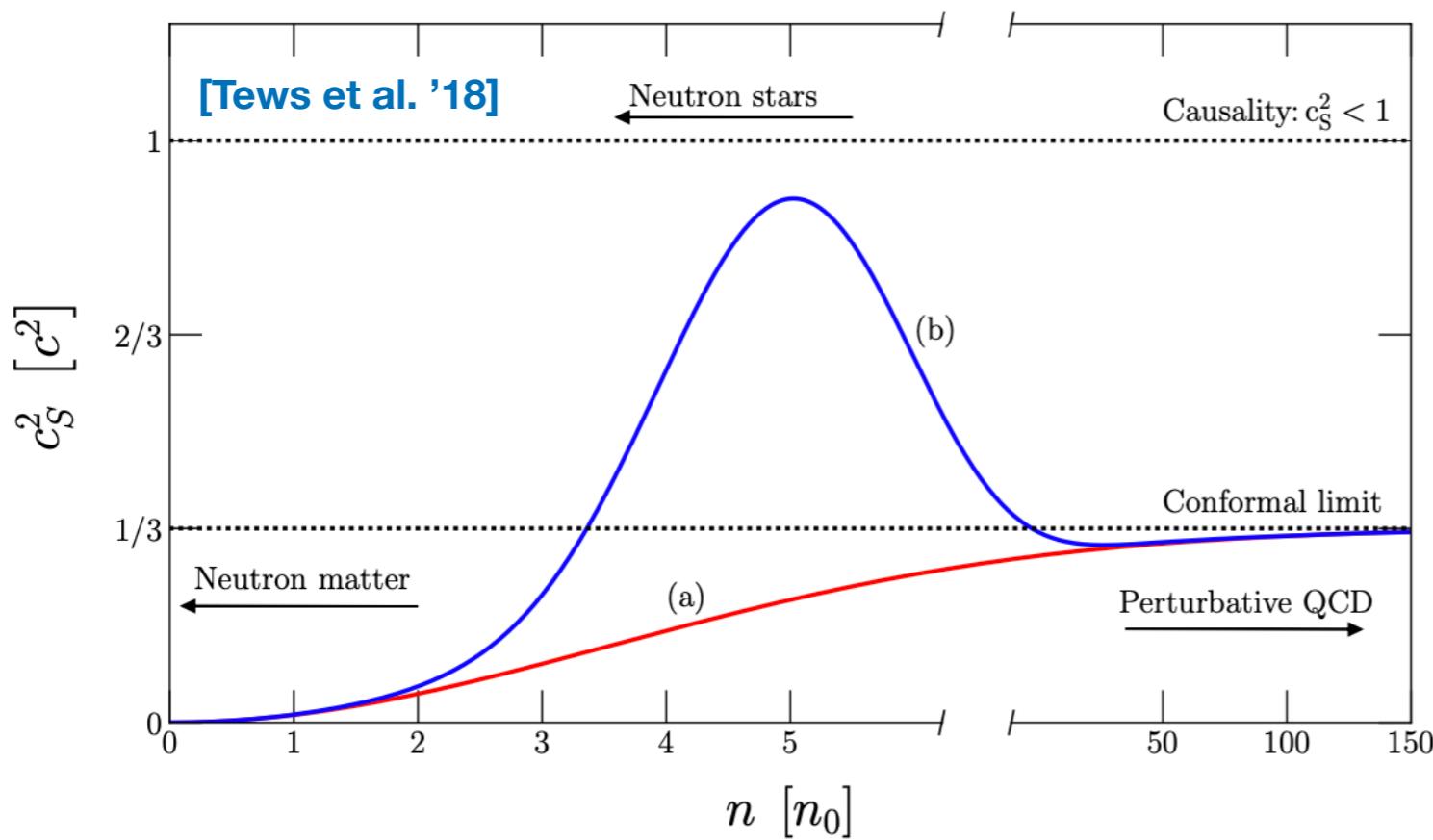
From the phase structure to **thermodynamics**



Thermodynamics of dense matter: why?

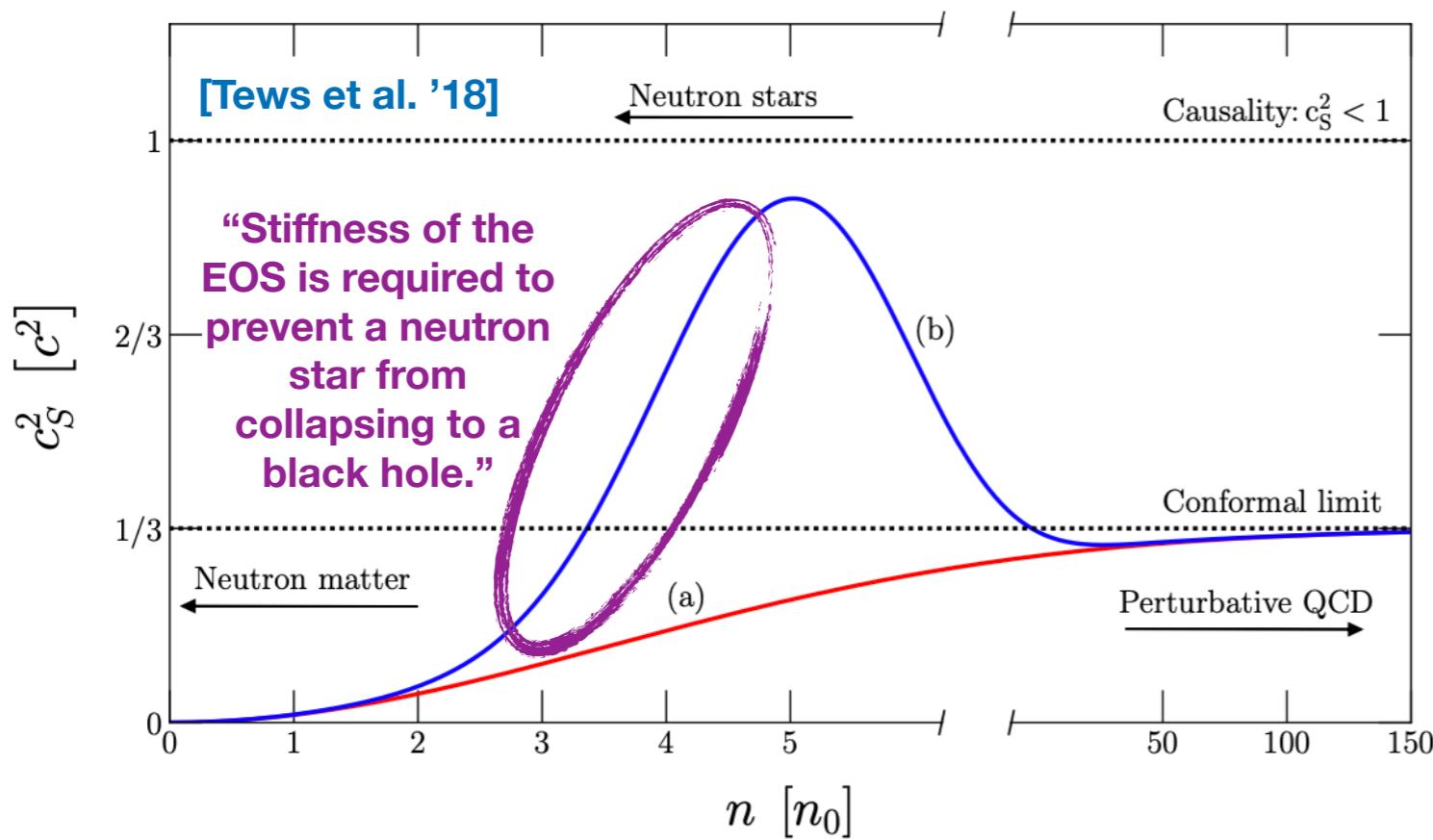


Constraints on dense matter: speed of sound



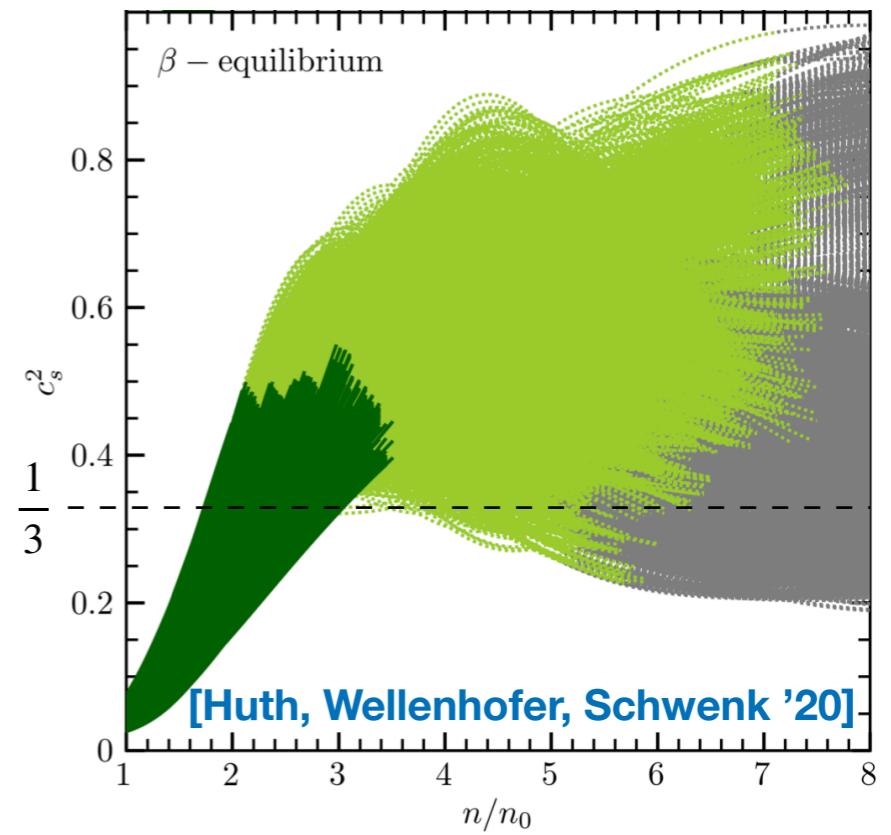
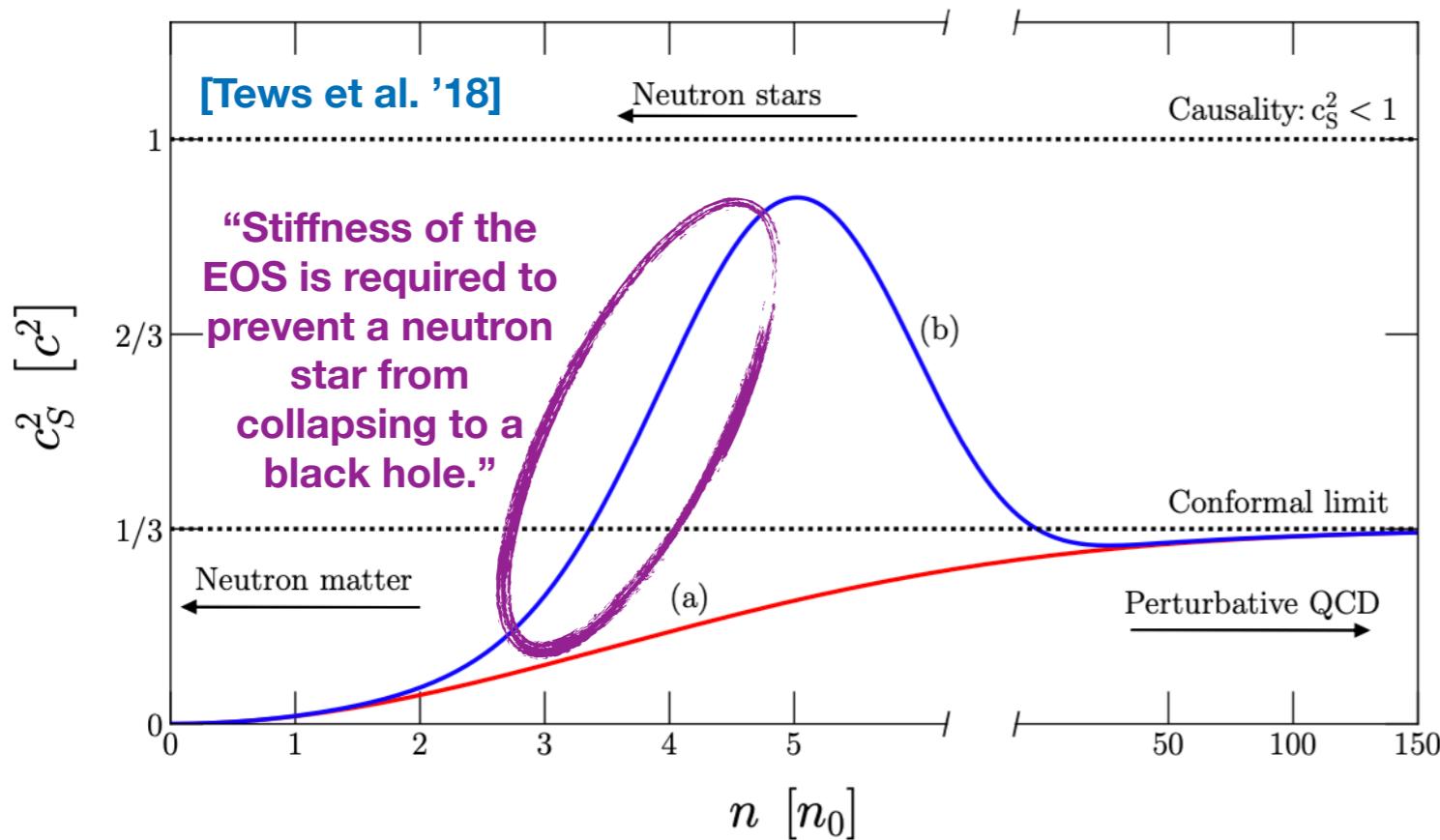
- **Speed of sound:** $c_s^2 = \partial P / \partial \epsilon$
- Constraints from neutron-star observations and nuclear physics suggest the existence of a global **maximum of the speed of sound** [Bedaque, Steiner '15; Tews et al. '18; Greif et al. '18; Raaijmaakers '19]

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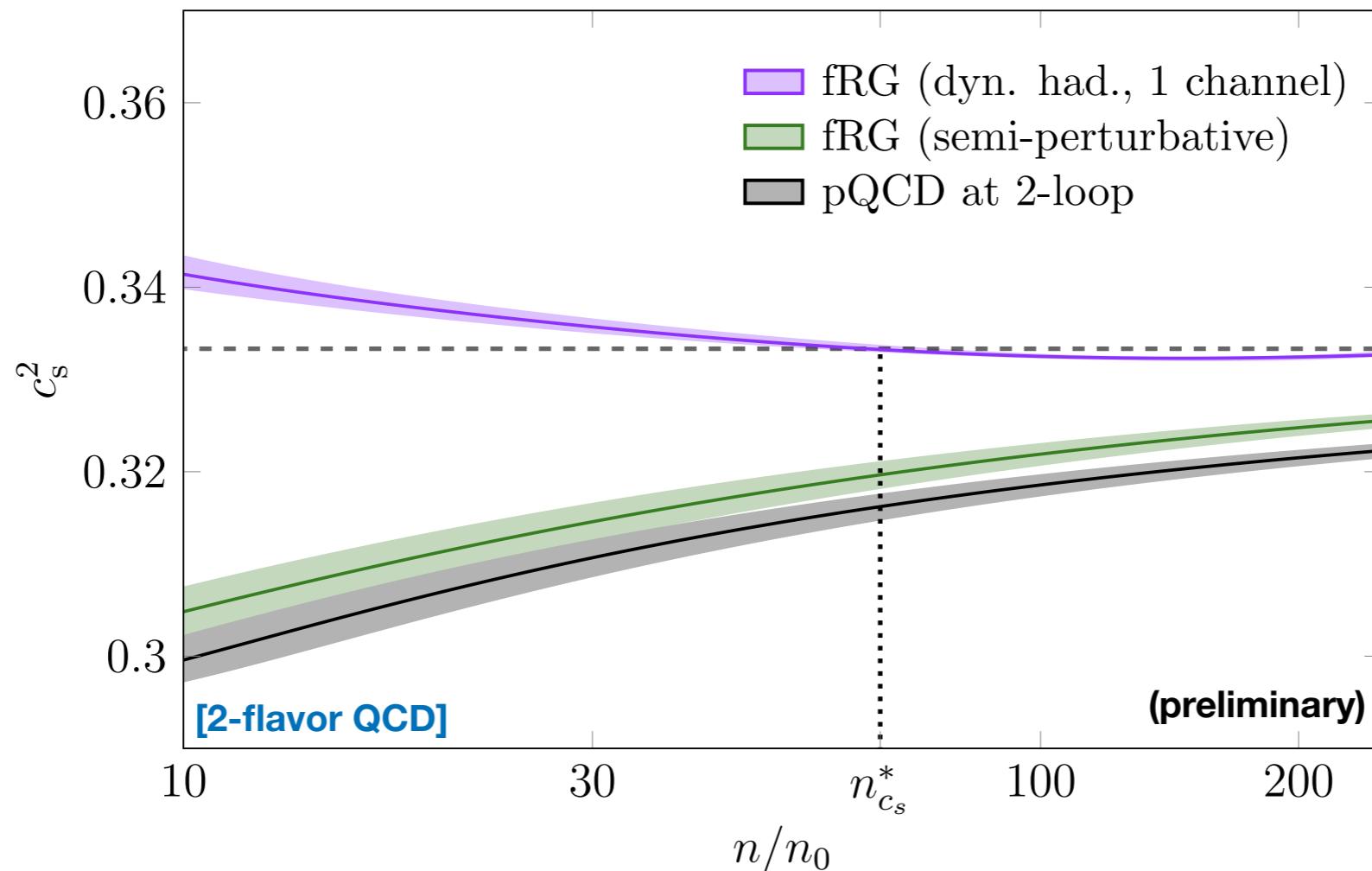
Constraints on dense matter: speed of sound



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Speed of sound of dense symmetric matter

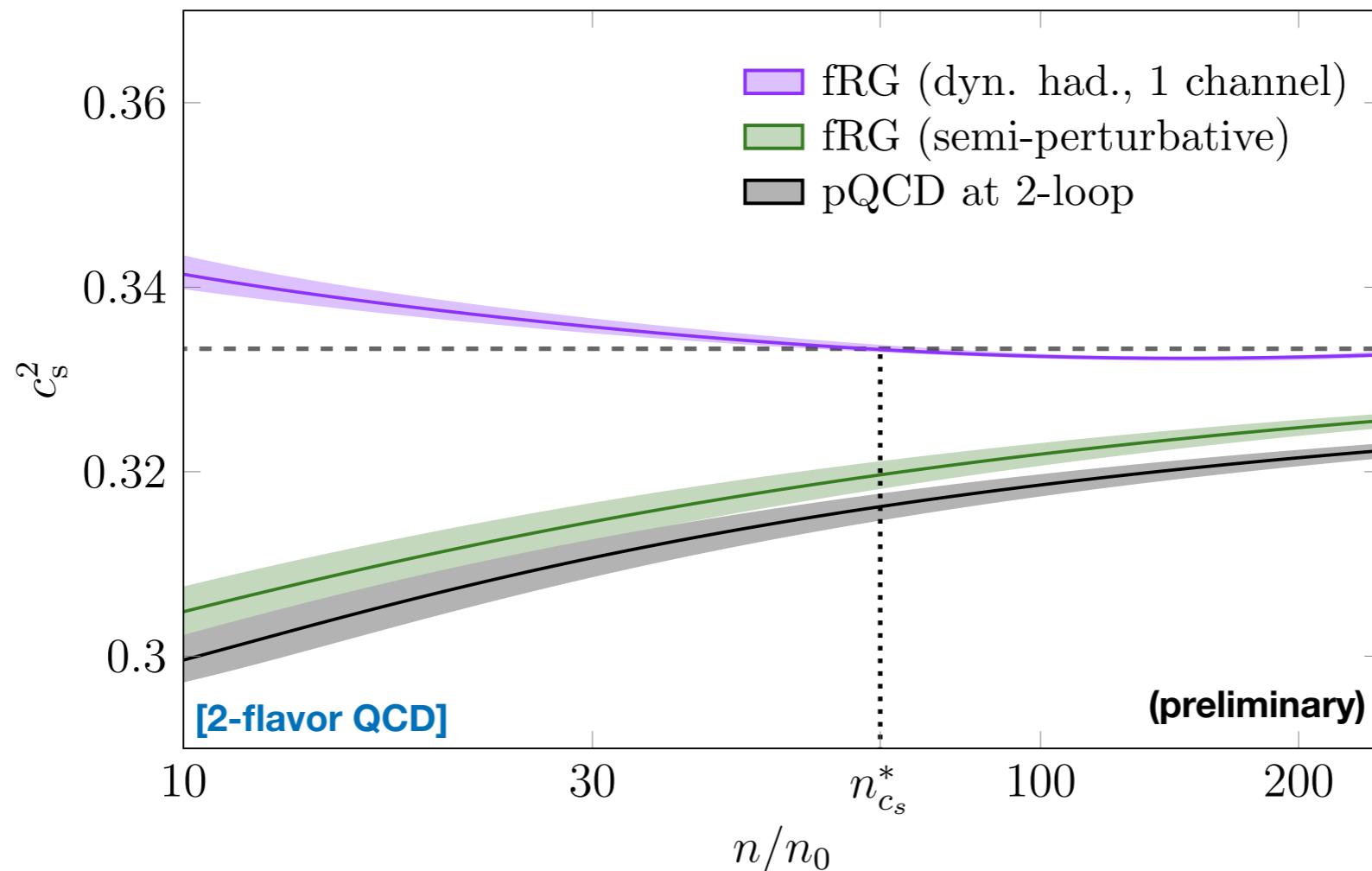
[JB & Schallmo '21; JB, Geissel, Schallmo '22; JB, Geissel, Schallmo (in prep.)]



- Consistent with **perturbative QCD at high densities**
- Deviations from pQCD because of non-perturbative form of the **gap**: $\Delta_0 \sim \exp(-c/(g^4 \mu^2))$

Speed of sound of dense symmetric matter

[JB & Schallmo '21; JB, Geissel, Schallmo '22; JB, Geissel, Schallmo (in prep.)]

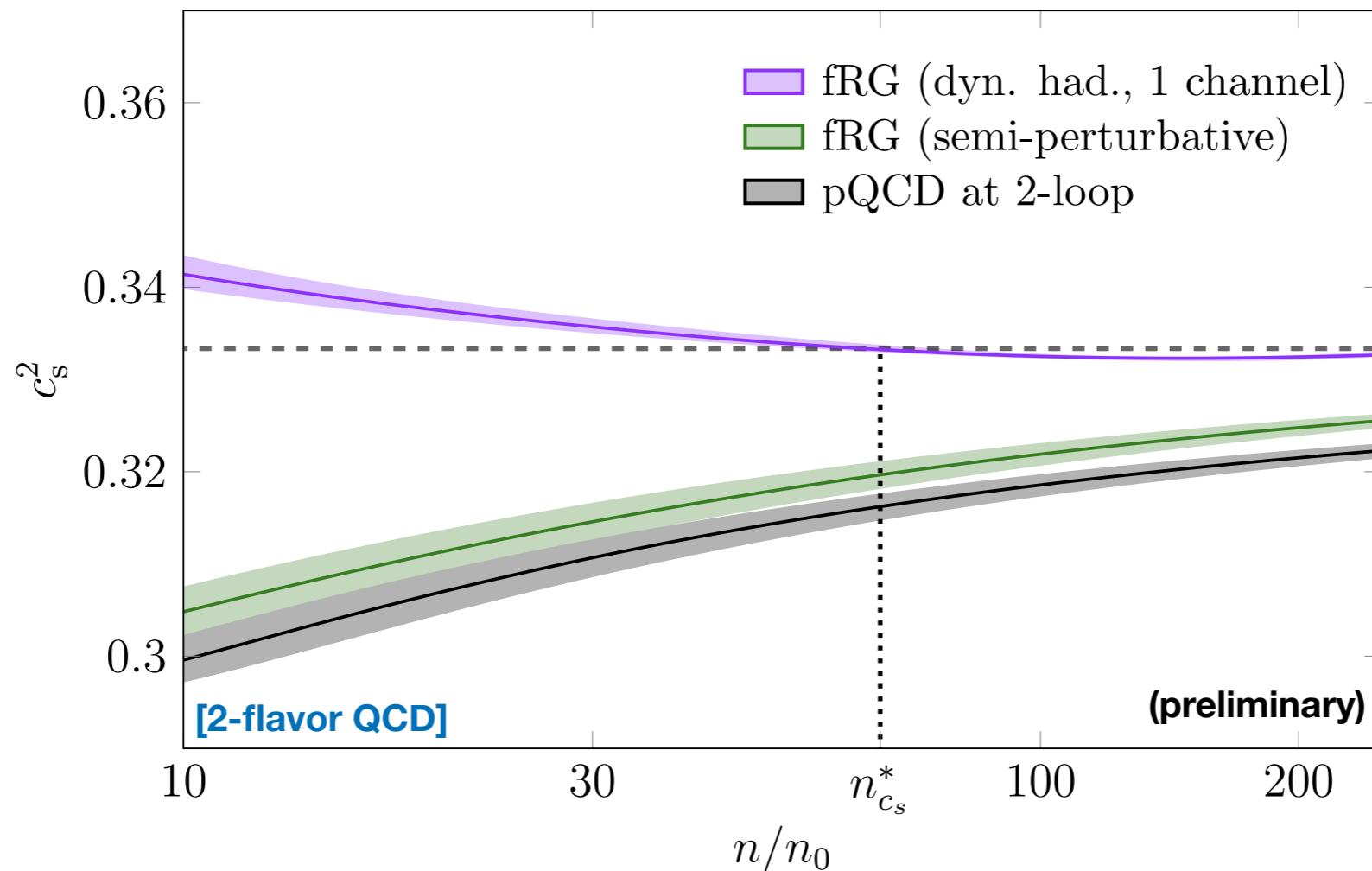


- **Scaling law** for the speed of sound: [JB, Geissel, Schallmo '22]

$$c_s^2 = \frac{1}{3} + \bar{c}_0(1+\sigma)n^{\frac{2(\sigma-1)}{3}} - \bar{c}_1/\ln(\bar{c}_2 n^{\frac{1}{3}})$$

Speed of sound of dense symmetric matter

[JB & Schallmo '21; JB, Geissel, Schallmo '22; JB, Geissel, Schallmo (in prep.)]



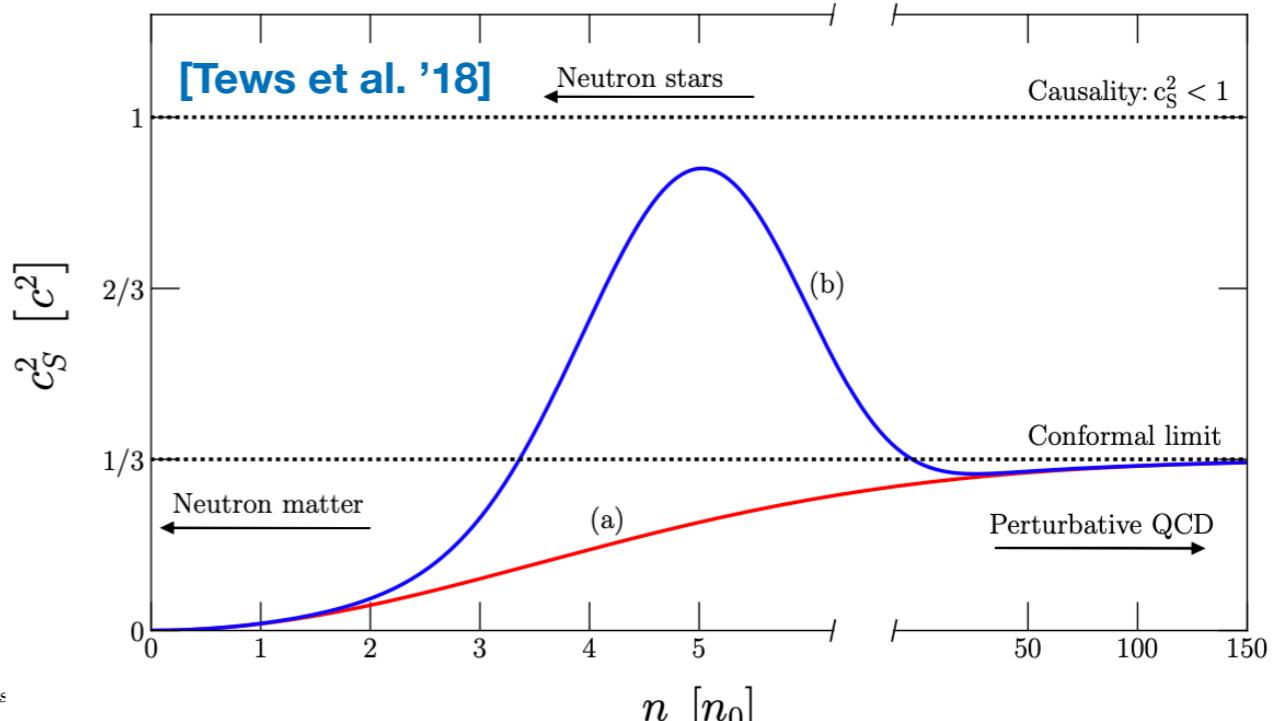
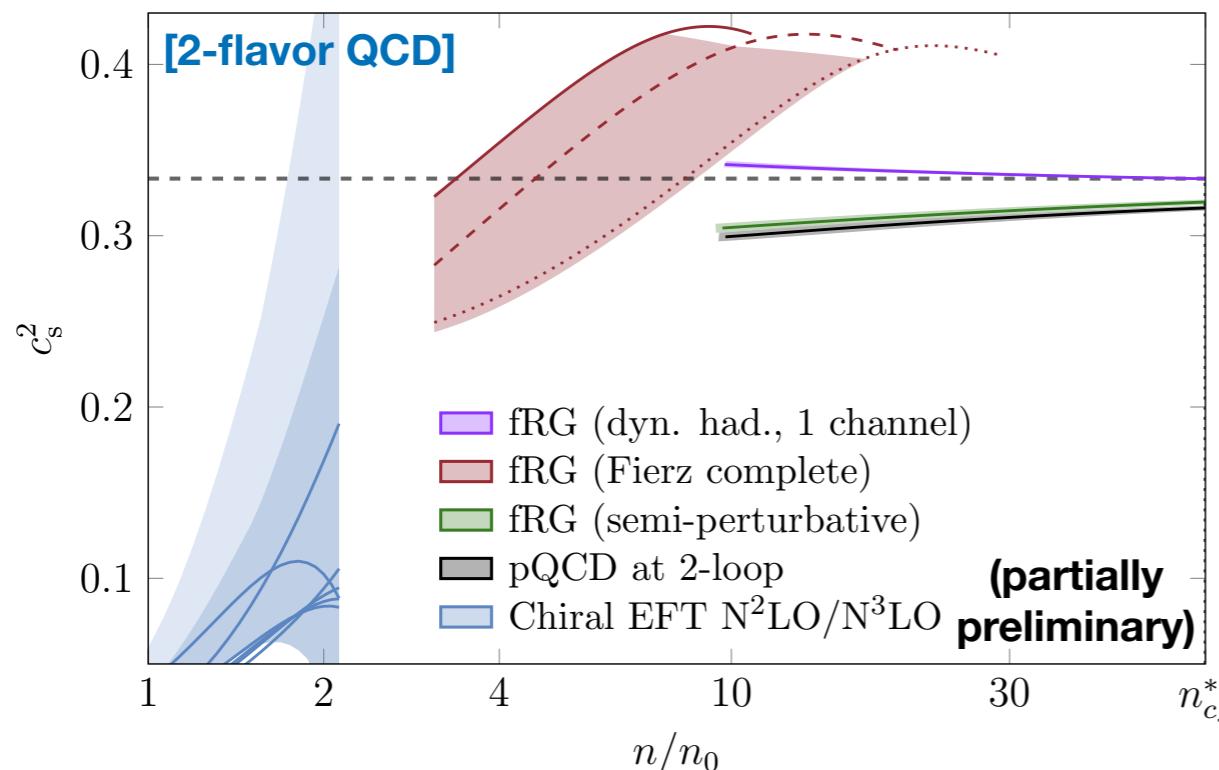
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The equation shows the scaling law for the speed of sound squared, c_s^2 , as a function of density n . The terms are:
1. **free gas** (blue circle): $\frac{1}{3}$
2. **gap-induced correction** (purple circle): $\bar{c}_0(1+\sigma)n^{\frac{2(\sigma-1)}{3}}$
3. **perturbative QCD** (red circle): $-\bar{c}_1/\ln(\bar{c}_2 n^{\frac{1}{3}})$

Cautious comparison to astrophysical constraints

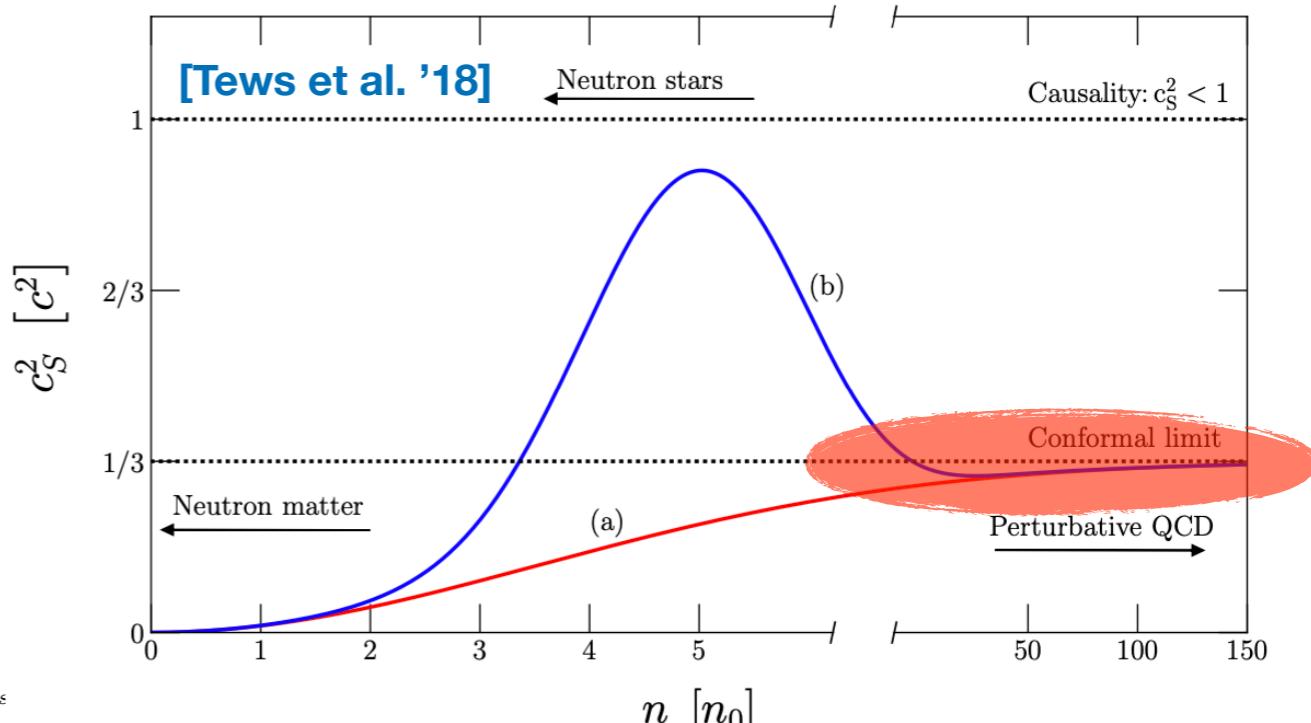
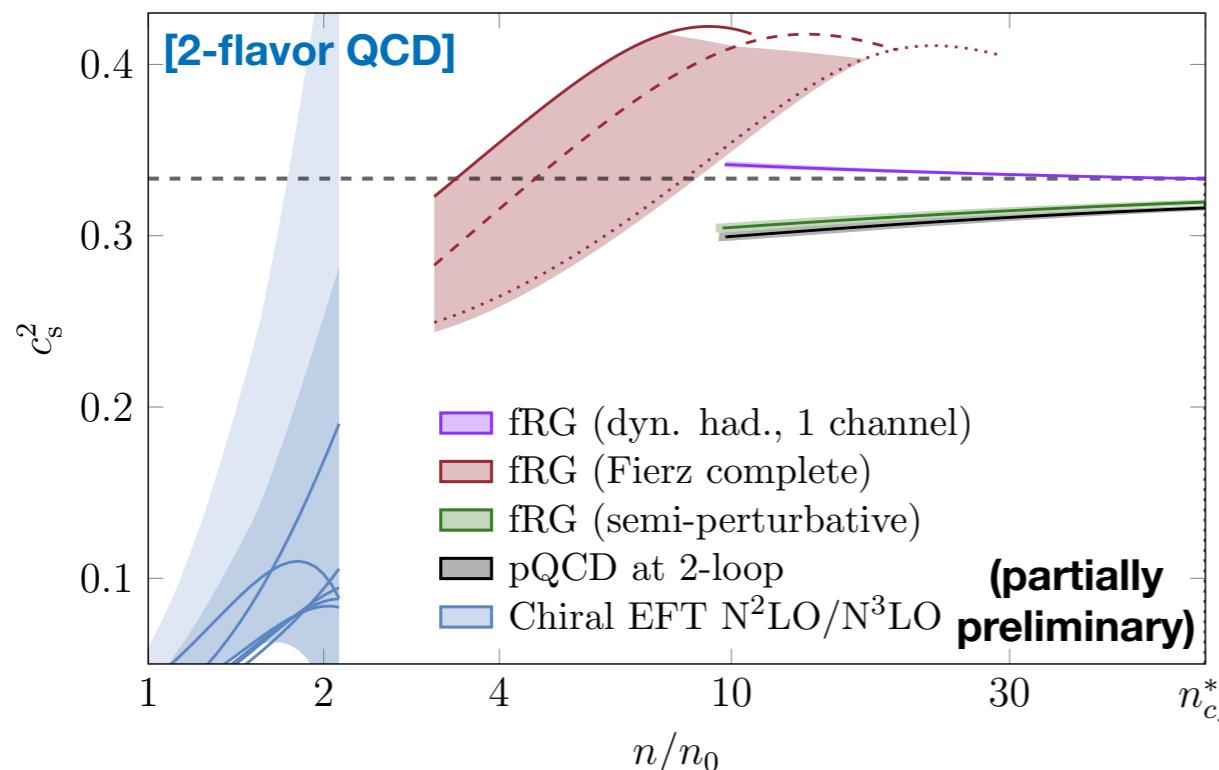
[Leonhardt et al. '19; JB & Schallmo '21; JB, Geissel, Schallmo '22; JB, Geissel, Schallmo (in prep.)]



- Existence of a maximum is in accordance with neutron-star observations and nuclear physics [Bedaque, Steiner '15; Tews et al. '18; Greif et al. '18; Raaijmaakers '19; Huth et al. '20; Huth et al. '22; Altiparmak, Ecker, Rezzolla '22; Ecker & Rezzolla '22]

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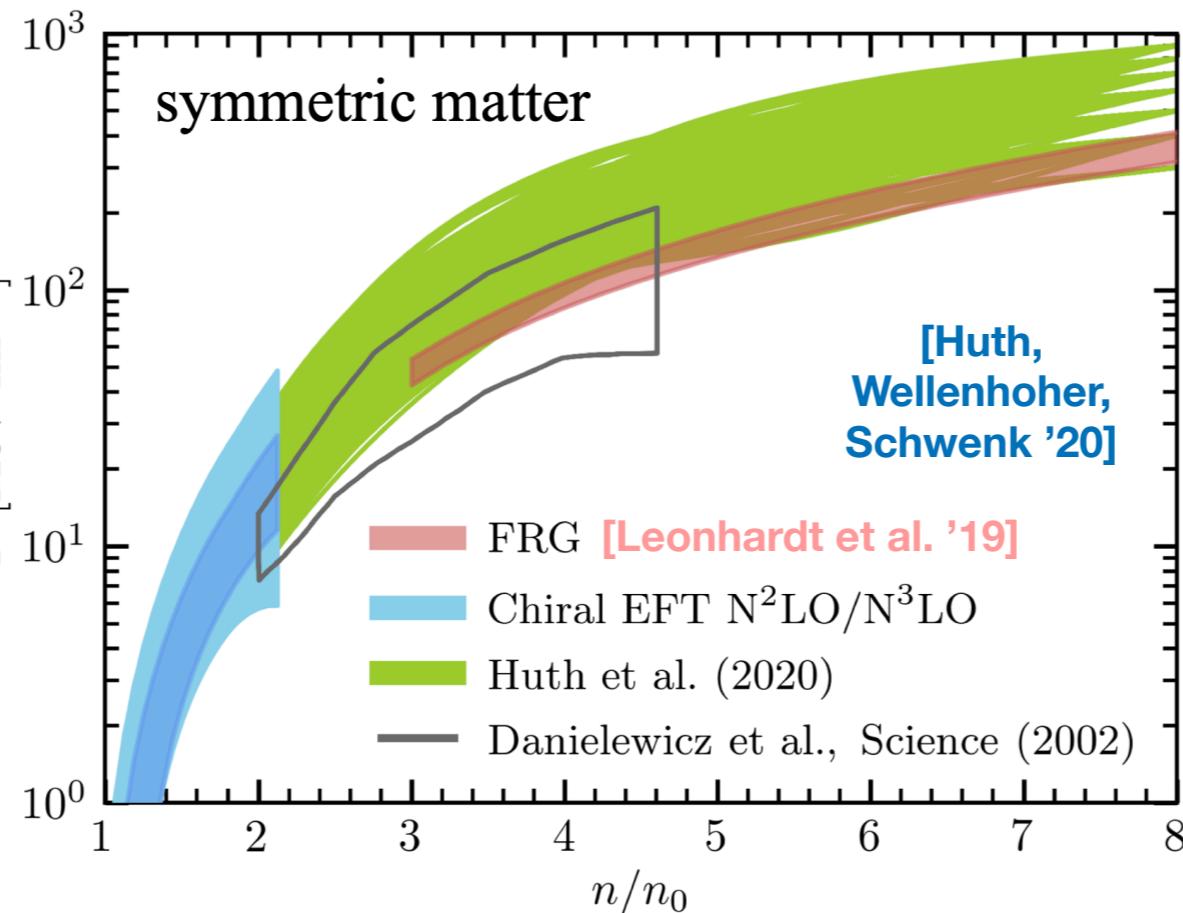
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- **“High-density crossing point”** appears to be directly related to the **color-superconducting gap**

Constraints on the EOS of dense symmetric matter

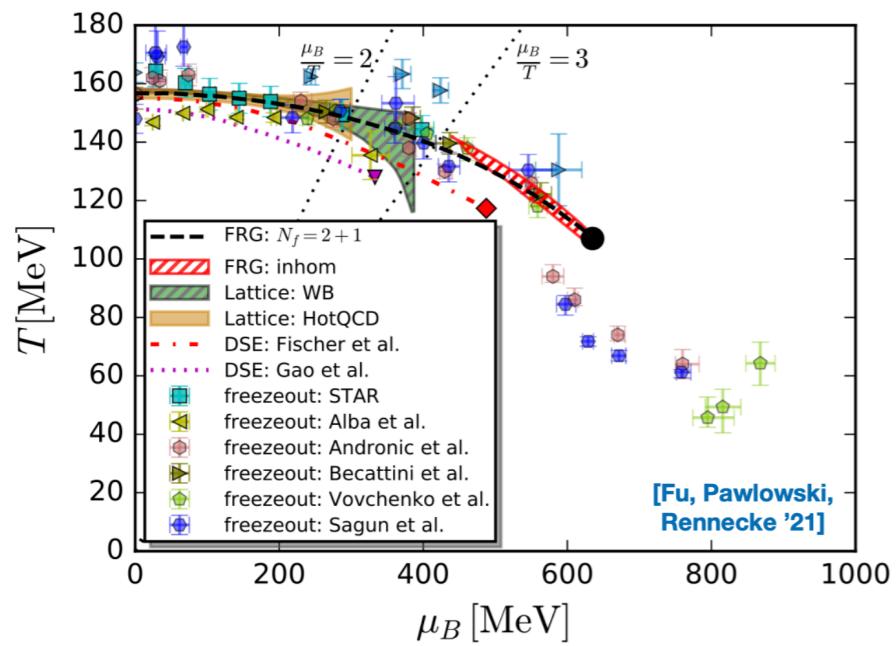
[incomprehensive list
of studies
of low-density thermodynamics:
Floerchinger & Wetterich '12
Wellenhofer et al. '14;
Drews, Weise '15;
Hu et al. 16;
**Drischler, Hebeler,
Schwenk '17,**
Otto, Oertel, Schaefer '20;
Tripolt et al. '21;
**Lu et al. (NLEFT
collaboration) '20;]**



- FRG results can be used to constrain **new density functionals** together with constraints coming from neutron-star observations & chiral EFT [Huth, Wellenhofer, Schwenk '20]
- Consistent with constraints from **heavy-ion collision experiments**

Conclusions

- **Exciting times:** Constraints from heavy-ion collision experiments and neutron-star observations



- Good agreement of FRG results & lattice QCD data at low densities
- **FRG predictions** for the finite-temperature phase structure of QCD at finite density

- **EOS results** are consistent with **results from studies based on chiral EFT** at low densities and with **neutron-star observations**

