

Positivity bounds in Effective Field Theories

Scott Melville

25 Jul 2022



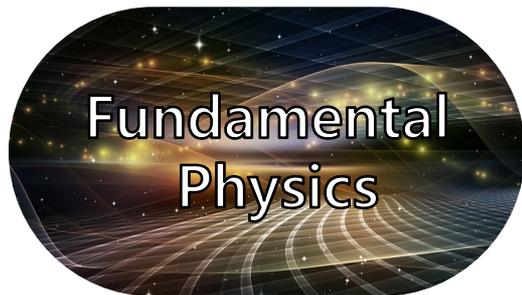
Big Picture

Big Picture

Energy ↑

UV

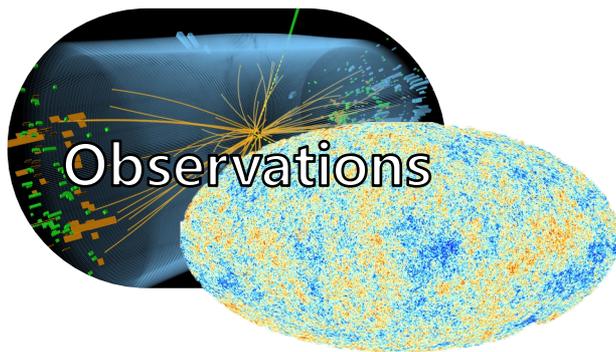
Inaccessible



Fundamental
Physics

IR

Accessible



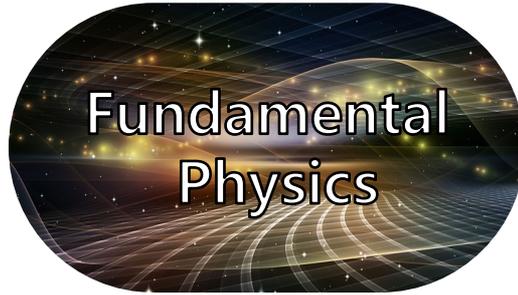
Observations

Big Picture

Energy ↑

UV

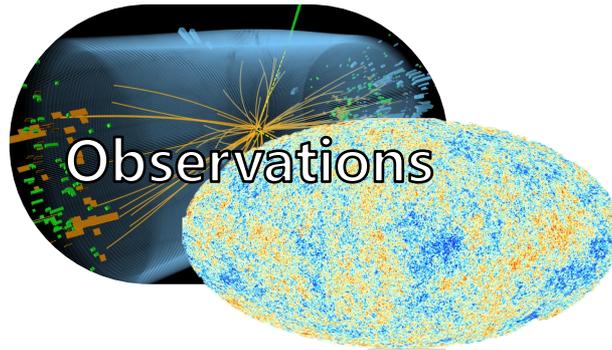
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Fundamental Physics

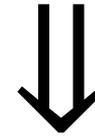
IR

Accessible



Observations

Basic UV properties
(causality, unitarity
symmetries, ...)



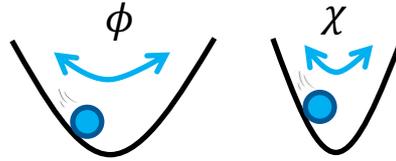
“Positivity”

Bounds on
low-energy
observables

Big Picture

Energy ↑

UV Inaccessible



All degrees of freedom

$$\mathcal{L}_{UV}[\phi, \chi]$$

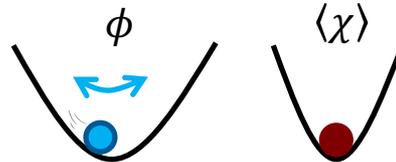
Fundamental



Effective

$$\mathcal{L}_{EFT}[\phi]$$

Only light degrees of freedom



IR Accessible



Basic UV properties
(causality, unitarity
symmetries, ...)

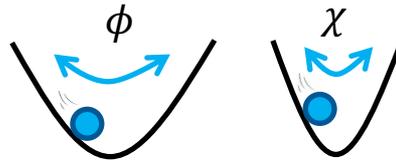
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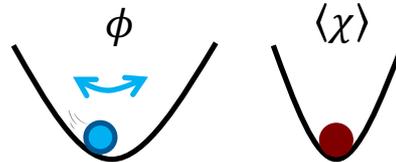
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Basic UV properties
(causality, unitarity
symmetries, ...)



**Constraints on
the RG flow**



Bounds on
low-energy
observables

Big Picture

Energy



$$\mathcal{L}_{\text{UV}}[\phi, \chi, \dots]$$

Λ

$$\mathcal{L}_{\text{EFT}}[\phi] = \mathcal{L}_{\text{renorm.}} + c_2(\Lambda)(\partial\phi)^4 + \dots$$

Heavy physics is captured by
non-renormalizable interactions

Big Picture

2 → 2 scattering
amplitude

$$A(s) = \left. \begin{array}{c} \phi \quad \phi \\ \diagdown \quad \diagup \\ \text{---} \text{---} \text{---} \text{---} \\ \diagup \quad \diagdown \\ \phi \quad \phi \end{array} \right\} \begin{array}{l} \text{Total energy,} \\ s = (p_1 + p_2)^2 \end{array}$$

Energy

$$\mathcal{L}_{\text{UV}}[\phi, \chi, \dots] \Rightarrow A_{\text{UV}}(s) = \begin{array}{c} m^2 < \Lambda^2 \\ \diagdown \quad \diagup \\ \text{---} \text{---} \text{---} \text{---} \\ \diagup \quad \diagdown \end{array} + \begin{array}{c} p^2 < \Lambda^2 \\ \diagdown \quad \diagup \\ \text{---} \text{---} \text{---} \text{---} \\ \diagup \quad \diagdown \end{array} + \begin{array}{c} p^2 > \Lambda^2 \\ \diagdown \quad \diagup \\ \text{---} \text{---} \text{---} \text{---} \\ \diagup \quad \diagdown \end{array} + \begin{array}{c} M^2 > \Lambda^2 \\ \diagdown \quad \diagup \\ \text{---} \text{---} \text{---} \text{---} \\ \diagup \quad \diagdown \end{array} + \begin{array}{c} M^2 > \Lambda^2 \\ \diagdown \quad \diagup \\ \text{---} \text{---} \text{---} \text{---} \\ \diagup \quad \diagdown \end{array} + \dots$$

Λ

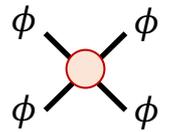
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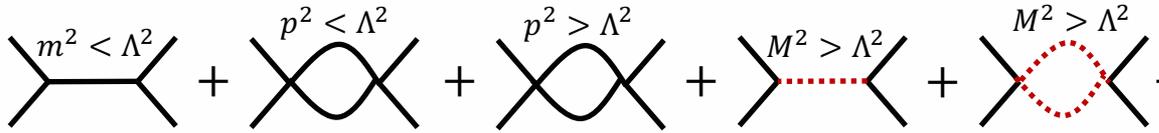
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where $c_n(\Lambda) \lesssim 1/\Lambda^{2n}$

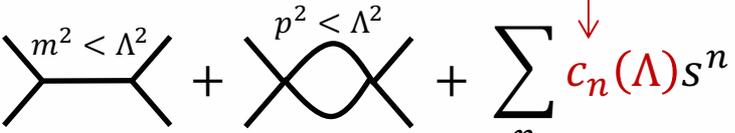
Big Picture

2 → 2 scattering amplitude $A(s) =$

 $\left. \vphantom{A(s)} \right\} \text{Total energy, } s = (p_1 + p_2)^2$

Energy ↑

$\mathcal{L}_{\text{UV}}[\phi, \chi, \dots] \Rightarrow A_{\text{UV}}(s) =$

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Λ -----

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Heavy physics is captured by non-renormalizable interactions

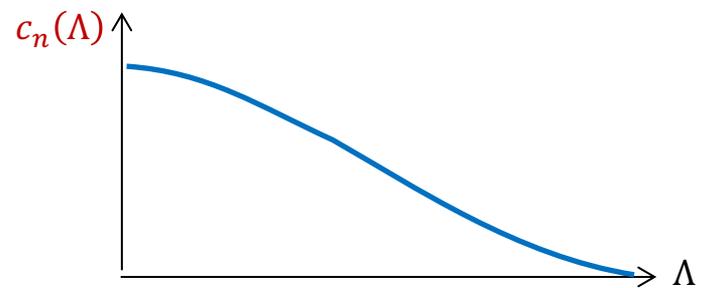
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Main result:

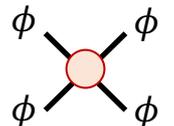
Causality+unitarity of heavy physics \Rightarrow

Constraints on RG flow

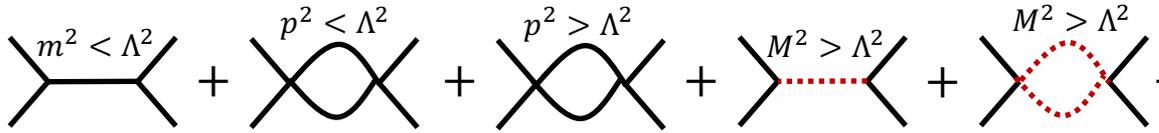
e.g. $\frac{\partial c_n(\Lambda)}{\partial \Lambda} \leq 0, \dots$



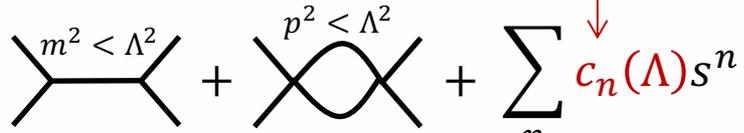
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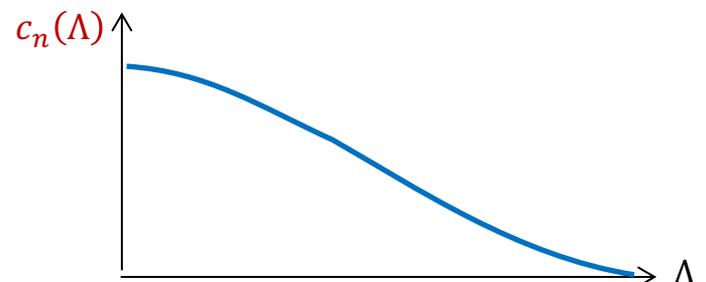
Useful because:

Renormalizable UV fixed point \Rightarrow

$c_n(\infty) = 0$

Bounded EFT coefficients

e.g. $c_n(\Lambda) \geq 0$



Outline

Positivity bounds
on $\partial_{\Lambda} c_n(\Lambda)$

GR as an EFT

Dark Energy
as an EFT

Positivity bounds

Integrating out heavy physics at the scale Λ produces terms like s^n/Λ^{2n} in $A(s)$.

$$\text{So expect } \Lambda \partial_\Lambda c_n(\Lambda) \sim \frac{1}{\Lambda^{2n}}$$

Positivity bounds

Integrating out heavy physics at the scale Λ produces terms like s^n/Λ^{2n} in $A(s)$.

$$\text{Causality of heavy physics} \Rightarrow \Lambda \partial_\Lambda c_n(\Lambda) = -\frac{\rho(\Lambda)}{\Lambda^{2n}} \text{ where } \rho(\Lambda) = \frac{2}{\pi} \text{Im } A_{UV}(s = \Lambda^2)$$

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Sketch of proof:

Causality dictates the
analytic structure of amplitude

Modern version of the
S-Matrix Program

[Bremermann, Bros, Epstein, Froissart,
Glaser, Gribov, Hepp, Jin, Kallen, Lehmann,
Mandelstam, Martin, Taylor, ..., 1960s]

The Analytic
S-Matrix

R.J. EDEN
P.V. LANDSHOFF
D.I. OLIVE
J.C. POLKINGHORNE

Cambridge University Press

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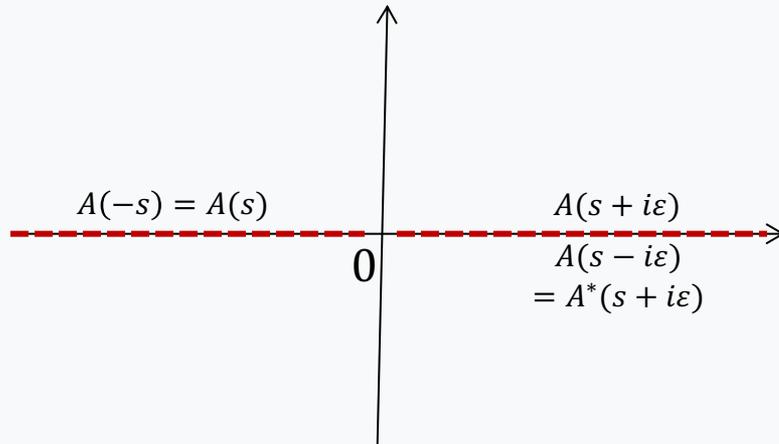
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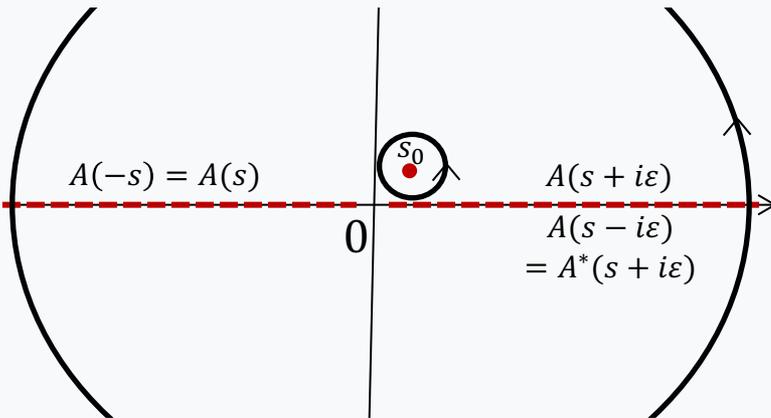
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Cauchy's theorem then relates the low- and high-energy amplitudes

$$A_{UV}(s_0) = \oint_{s_0} \frac{ds}{2\pi i} \frac{A_{UV}(s)}{s - s_0} = A_{UV}(\infty) + \int_0^\infty ds \frac{\rho(\sqrt{s})}{s - s_0}$$

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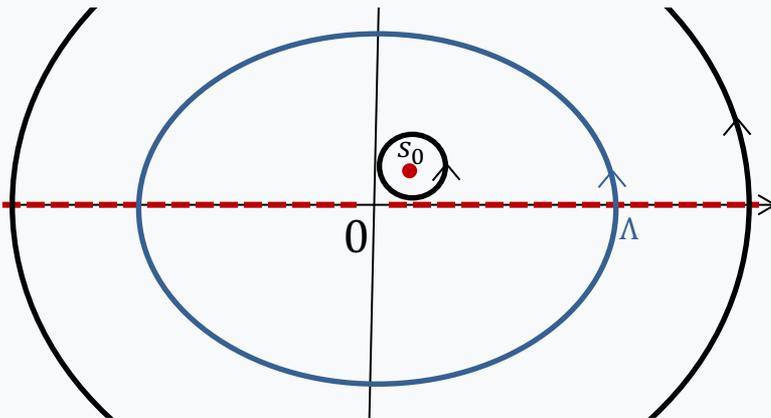
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EFT expansion corresponds to an intermediate contour

$$A_{EFT}(s_0) = \sum_n c_n(\Lambda) s_0^n + \int_0^{\Lambda^2} ds \frac{\rho(\sqrt{s})}{s - s_0}$$

$$\text{where } c_n(\Lambda) = c_n(\infty) + \int_{\Lambda^2}^\infty \frac{ds}{s} \frac{\rho(\sqrt{s})}{s^{2n}}$$

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Integrating out heavy physics at the scale Λ produces terms like s^n/Λ^{2n} in $A(s)$.

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Simple example: $\mathcal{L}_{UV} \supset g_1 \chi \phi^2 + g_2 \chi^2 \phi^2$

$$\Rightarrow A_{UV}(s) \supset \begin{array}{c} \phi \\ \diagdown \\ \chi \\ \diagup \\ \phi \end{array} \begin{array}{c} \phi \\ \diagup \\ \chi \\ \diagdown \\ \phi \end{array} + \begin{array}{c} \phi \\ \diagdown \\ \chi \\ \diagup \\ \phi \end{array} \begin{array}{c} \phi \\ \diagup \\ \chi \\ \diagdown \\ \phi \end{array} = \frac{g_1^2}{M_\chi^2 - s - i\varepsilon} + g_2^2 \sin^{-1} \left(\frac{\sqrt{s}}{2M_\chi} \right) \sqrt{\frac{4M_\chi^2}{s} - 1}$$

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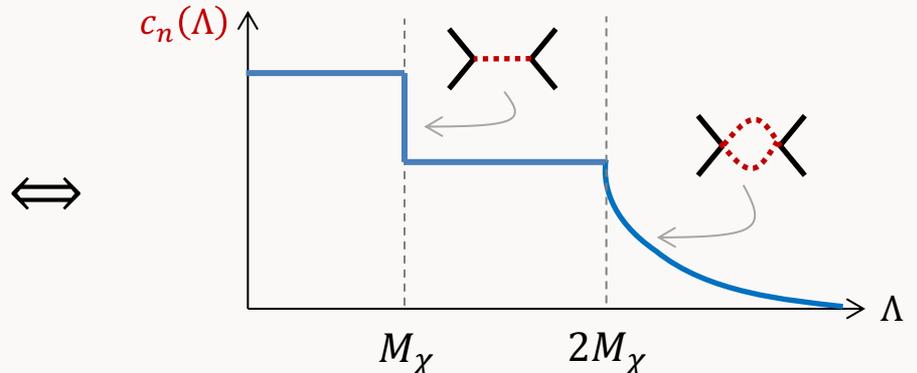
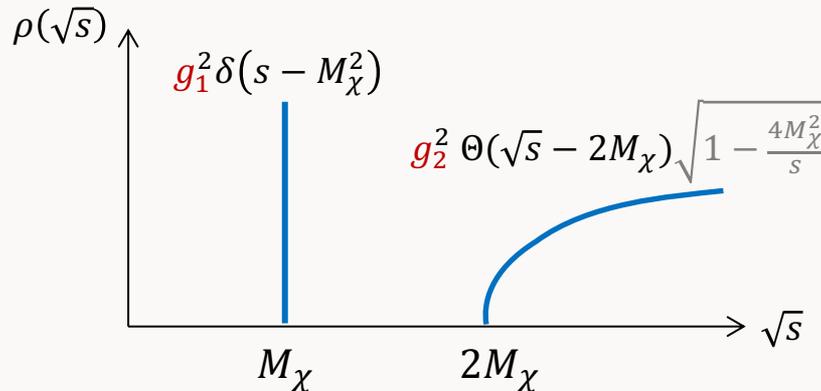
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The positivity of $\rho(\Lambda)$ is a general consequence of **unitarity** (probabilities add up to 1),

$$\text{since the probability } \langle p_1 p_2 | p_1 p_2 \rangle = |1 + i A(s)|^2 \leq 1 \Rightarrow 2 \text{Im } A_{UV}(s) \geq |A_{UV}(s)|^2 \geq 0$$

$$\text{Unitarity of heavy physics} \Rightarrow \rho(\Lambda) \geq 0 \Rightarrow \Lambda \partial_\Lambda c_n(\Lambda) \leq 0$$

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This monotonicity leads to a variety of bounds on the EFT coefficients $c_n(\Lambda)$

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e.g.

assuming UV
fixed point with

$$c_2(\infty) = 0,$$

$$c_4(\infty) = 0,$$

$$c_6(\infty) = 0$$

Causality
 \Rightarrow
Unitarity

$$c_2(\Lambda) \geq 0, \quad c_4(\Lambda) \geq 0, \quad c_2(\Lambda) \geq 0 \quad [\text{Adams++}, \text{ hep-th/0602178}]$$

$$\Lambda^8 c_6(\Lambda) \leq \Lambda^6 c_4(\Lambda) \leq \Lambda^4 c_2(\Lambda) \quad [\text{de Rham+SM+Tolley+Zhou 1702.06134}]$$

$$c_6(\Lambda) c_2(\Lambda) \geq c_4(\Lambda) c_4(\Lambda) \quad [\text{Bellazzini++}, 2011.00037] \\ [\text{Tolley+Wang+Zhou}, 2011.02400]$$

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[Adams++, hep-th/0602178]

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$$c_6(\Lambda) c_2(\Lambda) \geq c_4(\Lambda) c_4(\Lambda)$$

[Bellazzini++, 2011.00037]
[Tolley+Wang+Zhou, 2011.02400]

We can directly compare these with low-energy measurements of the $c_n(\Lambda)$

Positivity bounds

These bounds have been developed for **every** EFT operator that contributes to the $2 \rightarrow 2$ amplitude

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General
 $2 \rightarrow 2$ scattering
amplitude

$$A_{UV}(s, t) = \begin{array}{c} p_1 \quad p_3 \\ \diagdown \quad / \\ \text{red circle} \\ / \quad \diagdown \\ p_2 \quad p_4 \end{array}$$

depends on

Ingoing energy, $s = (p_1 + p_2)^2$
Transferred energy, $t = (p_1 - p_3)^2$
(note that $(p_1 - p_4)^2 = -s - t$)

$$\Rightarrow A_{\text{EFT}}(s, t) = \sum_{a,b} c_{2a,b}(\Lambda) (s + t/2)^{2a} t^b$$

Causality of
heavy physics

$$\Rightarrow \Lambda \partial_\Lambda c_{2a,b}(\Lambda) = \partial_t^b \left[\frac{-\rho(\Lambda, t)}{(\Lambda^2 + t/2)^{2a}} \right]_{t=0}$$

where $\rho(\Lambda, t) = \frac{2}{\pi} \text{Im} A_{UV}(s = \Lambda^2, t)$

Unitarity of
heavy physics

$$\Rightarrow \partial_t^b \rho(\Lambda, t) \Big|_{t=0} \geq 0$$

see e.g. [Nicolis+Rattazzi+Trincherini, 0912.4258]

Positivity bounds

These bounds have been developed for **every** EFT operator that contributes to the $2 \rightarrow 2$ amplitude

General
 $2 \rightarrow 2$ scattering
amplitude

$$A_{UV}(s, t) = \begin{array}{c} p_1 \quad p_3 \\ \diagdown \quad / \\ \text{red circle} \\ / \quad \diagdown \\ p_2 \quad p_4 \end{array}$$

depends on

Ingoing energy, $s = (p_1 + p_2)^2$
Transferred energy, $t = (p_1 - p_3)^2$
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$$\Rightarrow A_{\text{EFT}}(s, t) = \sum_{a,b} c_{2a,b}(\Lambda) (s + t/2)^{2a} t^b$$

Causality of
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$$\Rightarrow \Lambda \partial_{\Lambda} c_{2a,b}(\Lambda) = \partial_t^b \left[\frac{-\rho(\Lambda, t)}{(\Lambda^2 + t/2)^{2a}} \right]_{t=0}$$

where $\rho(\Lambda, t) = \frac{2}{\pi} \text{Im} A_{UV}(s = \Lambda^2, t)$

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Causal + Unitary \Rightarrow

Every $\Lambda \partial_{\Lambda} c_{2a,b}(\Lambda)$ is bounded

$$\text{e.g. } \Lambda^6 \partial_{\Lambda} c_{2a,1} + \frac{2a+1}{2} \Lambda^4 \partial_{\Lambda} c_{2a,0} \geq 0$$

[de Rham+SM+Tolley+Zhou, 1702.06134]

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Also implies non-linear bounds on the $c_{2a,b}$

[Bellazzini+, 2011.00037]

Using Bose symmetry ($s \leftrightarrow t \leftrightarrow u$ crossing) gives further bounds

[Tolley+Wang+Zhou, 2011.02400]

Bounds get even stronger for spinning particles

[Davighi+SM+You, 2108.06334]

Positivity bounds on $\partial_\Lambda c_n(\Lambda)$

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GR as an EFT

Dark Energy
as an EFT

GR as an EFT

Pure gravity in 4 spacetime dimensions is described at low energies by, (see e.g. [Donoghue, gr-qc/9512024])

$$\mathcal{L}_{\text{EFT}}[g_{\mu\nu}] = M_P^2 R + c_4(\Lambda) (R^{\mu\nu\rho\sigma} R_{\mu\nu\rho\sigma})^2 + \tilde{c}_4(\Lambda) (R^{\mu\nu\rho\sigma} \tilde{R}_{\mu\nu\rho\sigma})^2 + \dots$$

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Two distinct scattering amplitudes:

$$\begin{array}{cc} g_{\mu\nu}^+ & g_{\mu\nu}^+ \\ & \times \\ g_{\mu\nu}^+ & g_{\mu\nu}^+ \end{array}$$

$$A_{++++}(s) = s^4 c_4(\Lambda)$$

$$\begin{array}{cc} g_{\mu\nu}^+ & g_{\mu\nu}^+ \\ & \times \\ g_{\mu\nu}^x & g_{\mu\nu}^x \end{array}$$

$$A_{+x+x}(s) = s^4 \tilde{c}_4(\Lambda)$$

Renormalizable UV fixed point
+ causal, unitary RG flow

\Rightarrow

$$\begin{aligned} c_4(\Lambda) &\geq 0, \\ \tilde{c}_4(\Lambda) &\geq 0 \end{aligned}$$

[Bellazzini+Cheung+Remmen, 1509.00851]

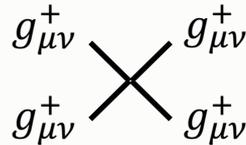
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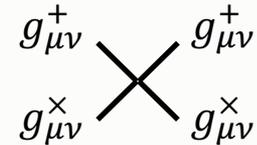
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Many recent developments:

- including matter fields
- stronger bounds (more UV input)
- different spacetime dimensions
- connections with string swampland

e.g. [Tokuda+Aoki+Hirano, 2007.15009]

[Alberte+de Rham+Jaitily+Tolley,, 2007.12667]

[Caron-Huot+Li+Parra-Martinez+Simmons-Duffins, 2201.06602]

[Bern+Herrmann+Kosmopoulos+Roiban, 2205.01655]

[Herrero-Valea+Koshelev+Tokareva, 2205.13332], ...

Dark Energy as an EFT

Scalar-tensor EFT

Metric \nearrow Light scalar field

$$\mathcal{L}_{EFT}[g_{\mu\nu}, \phi] \quad \text{with } \Lambda^3 \sim M_P H_0^2 \ll M_P^3$$

Many possible interactions. Two ways to proceed:

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(1) Assume more symmetries, e.g. $\phi \rightarrow \phi + c + c_\mu x^\mu$

$$\mathcal{L}_{EFT} = c_2 X + c_3 \mathcal{L}_3 + c_4 \mathcal{L}_4 + c_5 \mathcal{L}_5 \quad \text{with}$$

$$\mathcal{L}_3 = 2X (\nabla_\mu \nabla^\mu \phi), \quad \mathcal{L}_4 = 2X \left((\nabla_\mu \nabla^\mu \phi)^2 - (\nabla_\mu \nabla_\nu \phi)^2 \right)$$
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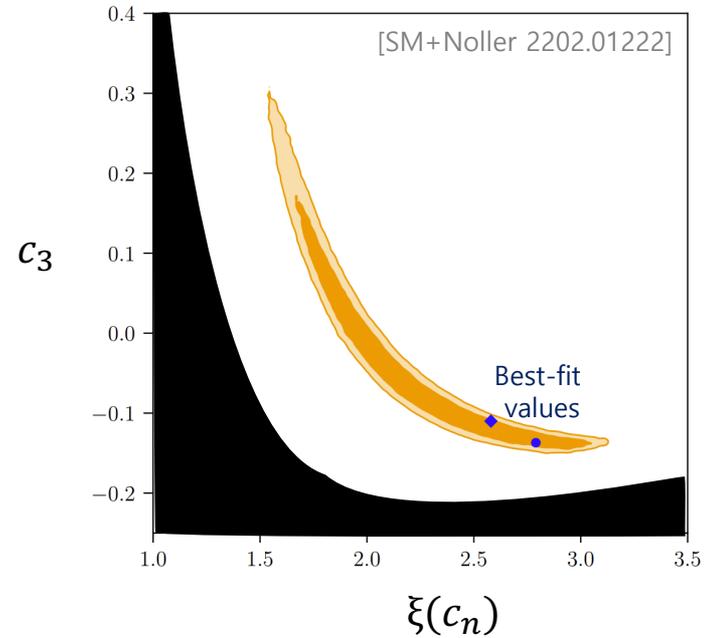
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This "covariant Galileon" EFT has a Λ CDM-like background,

$$\frac{\nabla\nabla\phi}{\Lambda^3} = -\xi(c_n) (\delta_\mu^\nu + q \delta_\mu^0 \delta_0^\nu) \quad [\text{Renk et al, 1707.02263}]$$

Fixed by eom \nearrow Deceleration parameter \nwarrow



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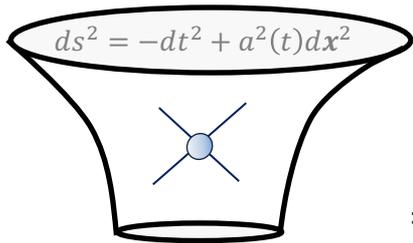
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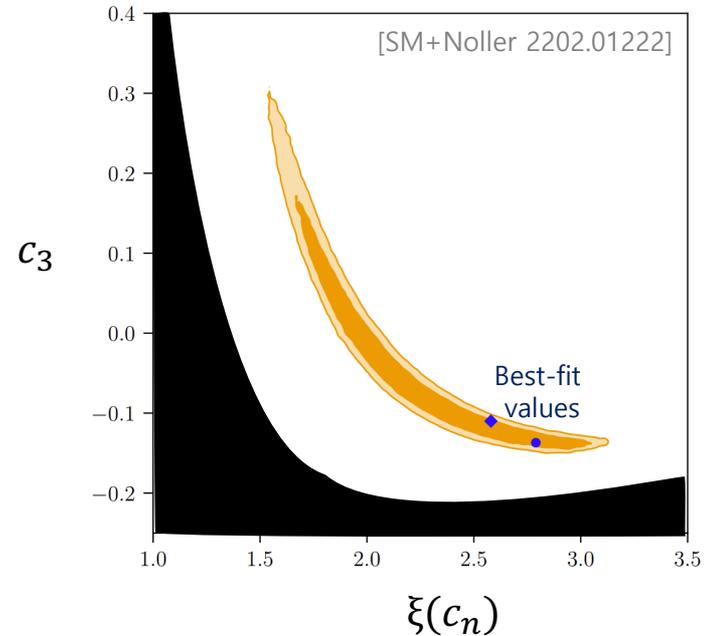
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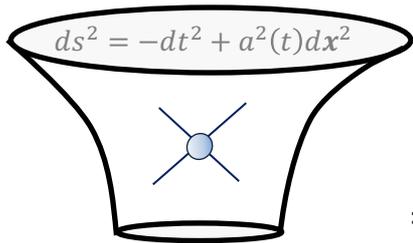
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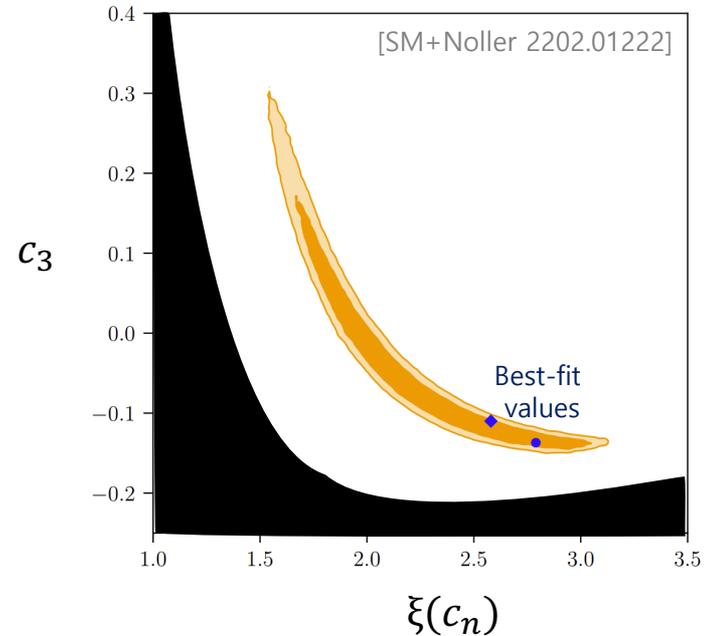
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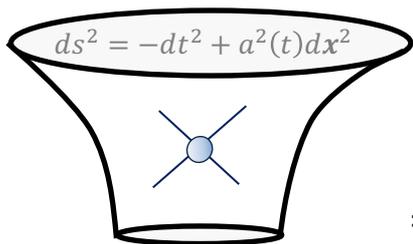
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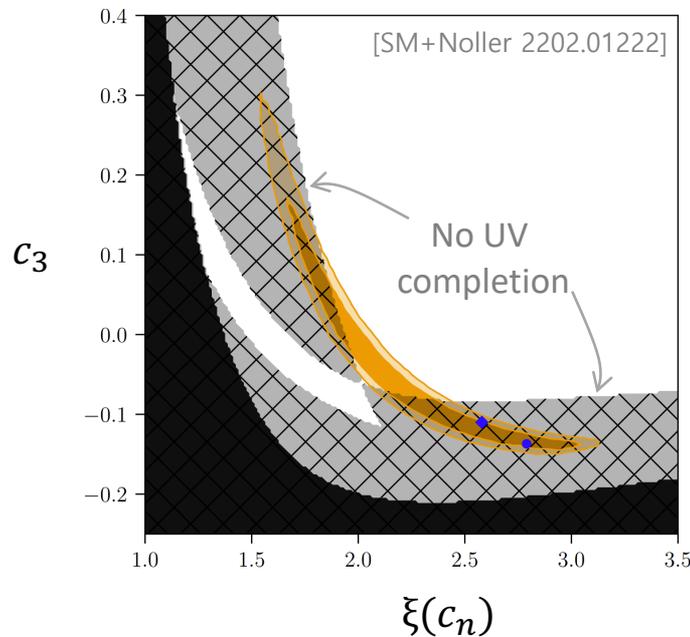
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This amplitude violates the positivity bounds required for a unitary UV completion in large regions of parameter space.



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Many possible interactions. Two ways to proceed:

(2) Work with general parametrisations

$$\mathcal{L}_{EFT} = M^2(X)R + P(X)$$

$$\text{with } X = (\partial\phi)^2$$

e.g. [SM+Noller, 1904.05874]

[Herrero-Valea+Timiryasov+Tokareva, 1905.08816]

[SM+Grall, 2102.05683]

[de Rham+SM+Noller, 2103.06855]

[Noumi+Tokuda, 2105.01436], ...

Some general lessons:

(i) **Subluminal GWs** are difficult* to UV complete

[de Rham+Tolley, 1909.00881]

[de Rham+SM+Noller, 2103.06855]

(ii) **Screening mechanisms** are difficult* to UV complete

[SM+Davis, 2107.00010]

* difficult = any operator that contributes to $2 \rightarrow 2$ scattering only satisfies these positivity bounds if they lead to superluminal GWs and no screening

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has $c_4(\Lambda) > 0$ and $\tilde{c}_4(\Lambda) > 0$

Dark Energy as an EFT

- (i) **Subluminal GWs** are difficult to UV complete
- (ii) **Screening mechanisms** are difficult to UV complete

