



Helmholtz Institute lena

Research Unit FOR 2783

DAAD German Academic Exchange Service

Vacuum birefringence and diffraction at XFEL: from analytical estimates to optimal parameters

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ExHILP 2021, September 16, Jena



Outline

- I. Introduction
- II. Theory
- III. Parameters
- IV. Results
- V. Conclusion & Outlook



Introduction

The quantum vacuum

(in this talk: the QED vacuum)



= vacuum fluctuations

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Introduction

The quantum vacuum + prescribed electromagnetic fields



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Introduction

Vacuum birefringence in laser fields



[Becker, Mitter: J. Phys. A **8** (1975)], [Aleksandrov, *et al.*: Sov. JETP **64** (1985)]

cf. also [Di Piazza, et al.: PRL 97 (2006)], [Dinu, et al.: PRD 89 & 90 (2014)], [Karbstein, et al.: PRD 92 (2015)], [Schlenvoigt, et al.: Phys. Scr. 91 (2016)], [King, Heinzl: HPL Sci. & Eng. 4 (2016)], [Karbstein, Sundqvist: PRD 94 (2016)], [Karbstein: PRD 98 (2018)], [Ataman: PRA 97 (2018)], [Briscese: PRA 97 (2018)], [Shen, et al.: Plasma Phys. CF 60 (2018)], [Huang, et al.: PRD 100 (2019)], [Karbstein, EM: PRD 100 (2019)]

 \rightarrow high-purity x-ray polarisation \mathcal{P} urity = $1.4 \cdot 10^{-11} @ \omega = 12914 \, eV$

[Marx, et al.: PRL **110** (2013)] [Marx, et al.: in preparation]



For our consideration we use the leading nonlinear correction to classical Maxwell theory

$$(\hbar = c = 1)$$

$$\mathcal{L}_{\text{HE}} = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} + \frac{m_e^4}{360\pi^2} \left[\frac{(\vec{B}^2 - \vec{E}^2)^2}{E_{\text{cr}}^4} + 7 \frac{(\vec{B} \cdot \vec{E})^2}{E_{\text{cr}}^4} \right] + \cdots$$

$$\text{[Euler, Kockel: Naturwiss. 32 (1935)]}_{[Euler: Ann. Phys. 26 (1936)]}$$
where
$$E_{cr} = \frac{m_e^2 c^3}{\hbar e} \simeq 1.3 \cdot 10^{18} \frac{V}{m}$$
Higher terms are
parametrically supressed by powers of:
Higher loops $\rightarrow \frac{\alpha}{\pi} \simeq 0.002 \ll 1$
Higher derivatives $\rightarrow \left(\frac{\partial}{m_e}\right)^2 \sim \left(\frac{\omega}{m_e}\right)^2 \ll 1$
Higher terms $\rightarrow \left(\frac{eF}{m_e^2}\right)^2 \sim \frac{I}{I_{\text{cr}}} \ll 1$

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Theory

We study QED vacuum birefringence and diffraction in the perfect head-on collision of a

loosely focused x-ray laser beam and a tightly focused high-intensity laser beam:

[[]Karbstein: PRD **98** (2018)] [Karbstein, Oude Weernink: PRD (2021)] **Poster of Ricardo Oude Weernink**



The effective vacuum-fluctuation-mediated interaction of the two laser beams gives rise to signal photons mainly emitted in the forward direction of the driving beams.

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For the driving laser fields we use solutions of the paraxial wave equation, namely linearly polarized fundamental Gaussian beam endowed with gaussian time envelope

- valid for small asymptotic beam divergences $\ \ \theta \ll 1$

$$\rightarrow \theta^2 = \left(\frac{w_0}{z_R}\right)^2 \simeq 0.02 \ll 1$$

• valid for large pulse durations

Theory

$$\rightarrow \frac{1}{\tau\omega} \simeq 0.01 \ll 1$$

cf., e.g., [L.W. Davis: PRA **19** (1979)] [J.P.Barton, D.R. Alexander: J.Appl. Phys. **66** (1989)] [Salamin, *et al.*: Phys. Rev. ST Acc. B. **5** (2002)]

cf., e.g., [Karbstein, EM: PRD 96 (2017)]





We aim at detecting signal photons far-outside the interaction region

- start configuration: vacuum plus driving laser fields $= |0\rangle$
- signatures of vacuum birefringence/diffraction are encoded in x-ray signal photons= $|\gamma_p(\vec{k})
 angle$
- transition amplitude of "vacuum emission"

 $\mathcal{S}_{(p)}(\vec{k}) = \left\langle \gamma_p(\vec{k}) \right| \underbrace{\sim}_{\mathbf{k}} \left\langle \gamma_p(\vec{k}) \right\rangle$

$$\circ$$
 direction \vec{k}
 \circ energy $|\vec{k}|$
 \circ polarization p

[Galtsov, Skobelev: Phys. Lett. **36** (1961)] [Karbstein, Shaisultanov: PRD **91** (2015)]

 $\mathrm{d}^{3}N_{(p)}(\vec{k}) = \frac{\mathrm{d}^{3}k}{(2\pi)^{3}} |\mathcal{S}_{(p)}(\vec{k})|^{2}$

Theory

• the signal photon numbers are dominated by (quasi-)elastic contributions

[Karbstein, Gies, Reuter, Zepf: PRD 92 (2015)]

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Theory

We are interested in two types of signals

• polarization-flipped photons $N_{(p)} = N_{\perp}$ becoming maximum for $\angle \phi = \frac{\pi}{4}$

Talk of Annika Tamara Schmitt

• total photon number $N_{
m tot}=\sum_p N_{(p)}$ in polarization insensitive measurement becoming maximum for $\angle\phi=rac{\pi}{2}$





The signal number is an integral expression depending on many experimental parameters





The signal number is an integral expression depending on many experimental parameters





Theory

The signal number is an integral expression depending on many experimental parameters

$$\begin{cases} \mathrm{d}N_{\mathrm{tot}} \\ \mathrm{d}N_{\perp} \end{cases} \simeq \begin{cases} 130 - 66\cos(2\phi) \\ 9\sin^2(2\phi) \end{cases} f(W,\Omega,\tau,w_0|N,\omega,T,w_{\mathrm{x}}|\vartheta)\mathrm{d}\cos(\vartheta) \end{cases}$$

In the experiment some of the given parameters depend on each other, e.g. at XFEL the photon number N per pulse is a function of both the pulse duration T and the photon energy ω .







The identification of the optimal experimental parameters is a complicated problem. Whithout additional simplifications this requires the use of numerical methods.

Building on several well-justified approximations and introducing the notion of an effective waist for the optical beam, we have constructed compact analytical approximations.

New approximation reproduces an exact numerical results for the signal driven by fundamental Gaussian beams on a one percent level accuracy in a wide range of parameters. [EM,Karbstein: PRD 104 (2021)]

Due to both its analytical nature and high accuracy, our model can even be straightforwardly differentiated and allows for the transparent tracing of parameter dependencies.



Parameters





Parameters

The incident probe pulse duration is nonlinearly increased by the silicon polariser.



Incident XFEL pulse duration

Pulse duration in the interaction region

Typical FWHM pulse durations after the polariser are

New J. Phys 23(2021)]

[Lindberg, Shvyd'ko, Phys. Rev. ST Accel.Beams **15**, (2012)] [Shvyd'ko, Lindberg, Phys. Rev. ST Accel.Beams **15**, (2012)] [Karbstein, Sundgvist, Schulze, Uschmann, Gies, Paulus:

 $T_{\mathcal{P}} \simeq 100 - 130 [\text{fs}]$

The temporal structure of realistic XFEL pulse is typically characterized by a large number of spikes; polarisers substantially smoothen the resulting pulse profile.



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Results: induced ellipticity

We identified the optimal parameters for maximizing the ellipticity of the outgoing x-ray beam.



The ellipticity ϵ is directly related to the ratio of the photon numbers $\epsilon^2 \simeq N_\perp/N$

The ratio is maximum at minimum pulse duration $T_i = 1.7 \, {
m fs} \,$ and x-ray waist $w_{
m x} \leq 0.1 w_0$

$$\frac{N_{\perp}}{N} \simeq \frac{4\alpha^4}{25(3\pi)^{\frac{3}{2}}} \left(\frac{W}{m_e}\frac{\omega}{m_e}\right)^2 \left(\frac{\lambda_{\rm C}}{w_0}\right)^4 F\left(\frac{4z_R}{\sqrt{T^2 + \frac{1}{2}\tau^2}}, \frac{T}{\tau}\right)$$

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[[]Karbstein: PRD 98 (2018)]

Results: induced ellipticity



The divergence of the signal photons essenctially fit the one of the driving x-ray beam with approximately constant ratio N_{\perp}/N

For the current HiBEF parameters and a polarization purity of $\mathcal{P} = 1.4 \times 10^{-11}$ the maximum value for the ratio is

 ${N_{\perp}\over N}={1\over 40}{\cal P}$ (above the background measurement requires ${N_{\perp}\over N}>{\cal P}$)

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The discernible number of polarization-flipped signal photons $N_{\perp>}$ is defined to fulfill the criterium



The discernible polarisation flipped photons appear due to the fact that vacuum birefringence is generically accompanied by a scattering phenomenon.



The dependence of the discernible photon number $N_{\perp>}$ on the parameters w_x and T_i is more complicated



- maximum signal for maximum duration
- for current HiBEF parameters there is a negligible amount of discernible photons for waists fulfilling $w_{\rm x}/w_0 \leq 1$
- there exists an optimal x-ray waist maximizing the discernible signal $N_{\perp>}$
- for HiBEF parameters the maximum number of discernible photos is $N_{\perp>}\simeq 0.01/{
 m shot}$ at $w_{\rm x}=2.1w_0$

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Our analytic formulas allow to analyse the dependence of the signal on different input parameters.

The dependence of the signal on the pump energy is demonstrated below.



The figure shows the **normalized** discernible signal.

The optimal x-ray waist is changing with pump energy.

At waist $w_{\rm x} = 2.1 w_0$ small variations of the energy will change the signal as

 $N_{\perp>} \sim W^{2\kappa} \simeq W^{2.3}$



The dependence of the discernible signal on the polarisation purity is less pronounced for current HiBEF parameters.



At waist $w_x = 2.1w_0$ small variations of the polarisation purity will change the signal as

$$N_{\perp >} \sim \mathcal{P}^{1-\kappa} \simeq \mathcal{P}^{-0.14}$$

The number of signal photons per background photon scattered outside the discernibility angle is

$$\frac{N_{\perp>}}{N_{>}\mathcal{P}} \simeq 8 \qquad @\mathcal{P}urity = 1.4 \cdot 10^{-11}$$

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Results: discernible number $N_{tot>}$

The discernible number of total signal photons $N_{tot>}$ is defined to fulfill the criterium



- No x-ray polariser is needed
- Probe pulse duration in the interaction region is the incident pulse duration of the XFEL $T_i \simeq 1.67 \dots 107 {
 m fs}$



Results: discernible number $N_{\rm tot>}$

The dependence of the discernible photon number $N_{tot>}$ on the parameters w_x and T_i is more complicated



- The optimal pulse duration is the maximum one
- There is an optimal x-ray waist maximizing the total discernible signal $N_{\rm tot>}$
- For current HiBEF parameters the maximum number of discernible photons is $N_{
 m tot>}\simeq 0.07/
 m shot$ at $w_{
 m x}=4.5w_0$



Results: discernible number $N_{tot>}$

The dependence of the signal on the pump energy



The figure shows the dependence of the **normalized** discernible signal on the pump energy.

At waist $w_x = 4.5w_0$ small variations of the energy will change the signal as

 $N_{\rm tot>} \sim W^{2\kappa} \simeq W^{2.06}$

The number of signal photons per background photon scattered outside the discernibility angle is

$$\frac{N_{\rm tot>}}{N_{>}} \simeq 32.6$$

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Results: discernible number $N_{tot>}$

Dependence of the signal photon number on different probe photon energies



The probe photon number at XFEL raises with decreasing photon energy. A smaller energy may increase the effect

$$N_{\rm tot>} \sim N(\omega) \times \omega^{2.06}$$



Conclusions & Outlook

Consider generic non-linear corrections to classical Maxwell theory

$$\mathcal{L} = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} + \frac{m_e^4}{360\pi^2} \left[a \, \frac{(\vec{B}^2 - \vec{E}^2)^2}{4E_{\rm cr}^4} + b \, \frac{(\vec{B} \cdot \vec{E})^2}{E_{\rm cr}^4} \right] + \cdots$$

Heisenberg-Euler at one loop:

$$a = 4, \quad b = 7$$

The corresponding signal photon distributions read:

$$\begin{cases} \mathrm{d}N_{\mathrm{tot}} \\ \mathrm{d}N_{\perp} \end{cases} \simeq \begin{cases} 2(a^2 + b^2) + 2(a^2 - b^2)\cos(2\phi) \\ (a - b)^2\sin^2(2\phi) \end{cases} \end{cases} f(W, \Omega, \tau, w_0 | N, \omega, T, w_{\mathrm{x}} | \vartheta) \,\mathrm{d}\cos(\vartheta)$$

Measurement of both signals (or total number at different polarization angles) allows to identify separately the coefficients *a* and *b*.

Talk of Luis Roso

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Conclusions & Outlook

- Analytic scalings are important to provide guidance for experiments.
- Fundamental Gaussian beams are certainly an idealisation More realistic beam models are to be considered (e.g., flattened-Gaussian beams)



- Quantitatively accurate predictions eventually require detailed numerical simulations of these quantum vacuum signatures in experimentally realistic laser fields
 [Blinne et al: PRD 99 (2019)]
- This could for instance be done with the vacuum emission solver of the Jena group, or an Adapted Yee scheme (FDTD) to solve the nonlinear Maxwell equations.

THANK YOU!





Appendix

Let us come back to the original expression for signal photons

$$\begin{cases} dN_{\text{tot}} \\ dN_{\perp} \end{cases} \simeq \vartheta \, d\vartheta \begin{cases} 2(a^2 + b^2) + 2(a^2 - b^2)\cos(2\phi) \\ (a - b)^2 \sin^2(2\phi) \end{cases} \frac{4\alpha^4}{225(3\pi)^{\frac{3}{2}}} N\left(\frac{W}{m_e}\frac{\lambda_{\text{C}}}{w_0}\right)^2 \left(\frac{\omega}{m_e}\right)^4 \\ \times \frac{\left(\frac{w_{\text{x}}}{w}\right)^2}{[1 + 2\left(\frac{w_{\text{x}}}{w}\right)^2]^2} e^{-\frac{1}{2}\frac{(\omega\vartheta w_{\text{x}})^2}{1 + 2\left(\frac{w_{\text{x}}}{w}\right)^2}} F\left(\frac{4z_R}{\sqrt{T^2 + \frac{1}{2}\tau^2}}, \frac{T}{\tau}\right) \end{cases}$$

$$F(\chi,\rho) := \sqrt{\frac{1+2\rho^2}{3}} \chi^2 e^{2\chi^2} \int_{-\infty}^{\infty} \mathrm{d}\kappa \, \mathrm{e}^{-\kappa^2} \Big| \sum_{\ell=\pm 1} \mathrm{e}^{2\ell\rho\kappa\chi} \mathrm{erfc}\big(\ell\rho\kappa+\chi\big) \Big|^2$$

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Appendix

For the high-intensity laser pump we consider the following field profile

$$\mathcal{E}(x) = \mathcal{E}_0 e^{-\frac{(z-t)^2}{(\tau/2)^2}} \frac{w_0}{w(z)} e^{-\frac{r^2}{w^2(z)}} \cos\left(\Omega(z-t) + \frac{\Omega r^2}{2 z \left[1 + (\frac{z_R}{z})^2\right]} - \arctan\left(\frac{z}{z_R}\right)\right),$$

pulse envelope

For the x-ray probe field is justified to use in the infinite Rayleigh length approximation

$$\mathfrak{E}(x) = \mathfrak{E}_0 e^{-\frac{(z+t)^2}{(T/2)^2}} e^{-\frac{r^2}{w_x^2}} \cos(\omega(z+t))$$

The peak field amplitudes can be expressed as

$$\mathcal{E}_{0}^{2} = 8\sqrt{\frac{2}{\pi}} \frac{W}{\pi w_{0}^{2} \tau}, \qquad \mathfrak{E}_{0}^{2} = 8\sqrt{\frac{2}{\pi}} \frac{N\omega}{\pi w_{x}^{2} T}$$

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September 16, 2021

[Karbstein, Mosman: PRD 96 (2017)]

